## What We Think We Know About CIRs/SIRs

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From:
Borovsky \& Denton, 2010.

Satellite View in RTN Coordinate System


## From: <br> Borovsky \& <br> Denton, 2010.

- Is an SIR a rapidly evolving CIR?
- Or is a CIR a steady-state SIR?
- Does it matter?


## What is a CIR? An "Artist's" View



From:
Tsurutani et al., 2006.

## What is a CIR? <br> An Analytical View



From:
Lee, 2000.

## What is a CIR? The In-Situ View



From:
McPherson \&
Weygand, 2006.

## What is a CIR? White-Light View



From:
Wood et al., 2010.


From:
Wood et al., 2010.



From:
Tappin et al., 1984.

## The Observational 'How'

- Much more can be said about CIRs
- deflections in solar wind flow direction at the stream interface
- formation and orientation of shock fronts
- Most of what we know has been learned from in-situ measurements
- A recent, thorough review can be found in Richardson [LRSP, 2018]
- PUNCH will use remote sensing
- PUNCH will measure scattered brightness from line-of-sight integrated mass
- PUNCH observations will be used to reconstruct the 3D structure of CIRS



## Simple, Ideal Coronal Holes



(b)
$\Delta=60 \mathrm{~km} / \mathrm{s}$

## Simple, Ideal Coronal Holes




TIME $=180.0 \mathrm{~h}$
(a) Total Brightness


## Synthetic Movies

Two movie files can be viewed

- Solar wind from an equatorial circular coronal hole
- Solar wind from a tilted streamer belt

Note: Dark is high brightness; white is low brightness


## 2008-11-19 stereo <br> observation <br> From: <br> Tappin \& Howard, 2009.

- PUNCH will provide a new observational perspective on CIRS
- Will PUNCH confirm our understanding of CIRs, tweak our understanding, or revolutionize our understanding?

