The Sun, Earth Climate, and Society "If It's Not Open, It's Not Science"

Greg Kopp Univ. of Colorado / LASP

4th Eddy Cross-Disciplinary Symposium 30 Oct. 2023 Golden, CO

Sun, Climate, and Society



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Visualizations by: Lori Perkins Written by: Sofie Bates Scientific consulting by: Gavin A. Schmidt Produced by: Kathryn Mersmann & Katie Jepson

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1880 - 1884

No, It's Not the Sun, Despite What You Might Read



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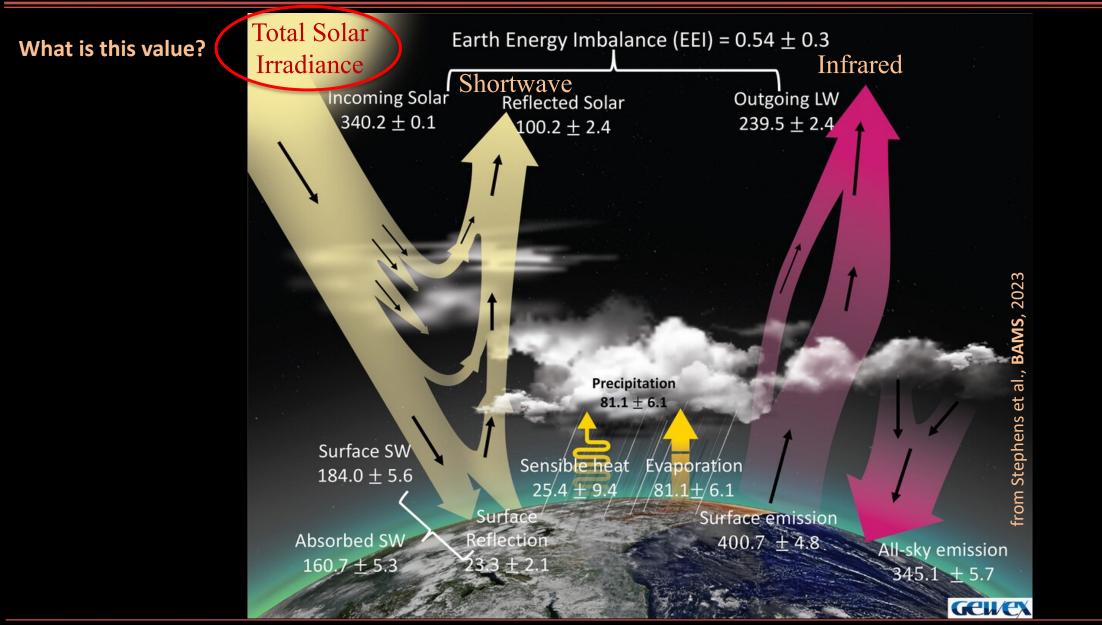
courtesy of Thierry Dudok de Wit, Univ. of Orleans

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Total Solar Irradiance Is Nearly All of the Earth's Energy Input





What Is <u>The Answer</u>?

BBC

"A simple answer?" "Yes" **1361 W m⁻²** The solar constant... ...that isn't

"There really is an answer?"

The Answer to the Ultimate Question *The Hitchhiker's Guide to the Galaxy* (BBC)

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Deep Thought

99.974 % of the Energy Heating the Earth Comes from the Sun

| | Heat Flux* | Uncertainty or Range | | | | | | |
|---|----------------------|----------------------|----------------|------------------------|-----------|--|--|--|
| Energy Source | [W m ⁻²] | [W m ⁻²] | Relative Input | | | | | |
| Solar Irradiance | 340.2 | 0.0000% | (1.000E+00) | Total Input (relative) | 1.000E+00 | | | |
| | | | | Temperature [°K] | 278 | | | |
| Secondary Sources of Solar Origin (Total) | 0.0268 | | 7.90E-05 | [°F] | 42 | | | |
| Infrared Radiation from the Full Moon | 0.01 | - | 2.90E-05 | | | | | |
| Combustion of Coal, Oil, and Gas (in U.S.) | 0.0052 | - | 1.50E-05 | | | | | |
| Dissipation of Magnetic Storm Energy | 0.00362 | 1.0E-05 to 1.0E-03 | 1.10E-05 | | | | | |
| Airglow Emission | 0.0036 | - | 1.10E-05 | 4000 X | | | | |
| Sun's Radiation Reflected from Full Moon | 0.0018 | - | 5.30E-06 | | | | | |
| Energy Generated by Solar Tidal Forces in the Atmosphere | 0.00168 | - | 4.90E-06 | Λ | | | | |
| Energy Dissipated in Lightning Discharges | 4.95E-04 | 9.0E-05 to 9.0E-04 | 1.50E-06 | /1 | | | | |
| Auroral Emission | 3.70E-04 | 1.0E-05 to 1.0E-03 | 1.10E-06 | 3000 X | | | | |
| Zodiacal Irradiance | 5.67E-05 | 5.65E-05 to 5.68E-05 | 1.70E-07 | | | | | |
| Earthshine | 1.93E-07 | - | 5.70E-10 | | | | | |
| | | | | | | | | |
| Secondary Sources of Non-Solar Origin (Total) | 0.0900 | | 2.60E-04 | ۷ | | | | |
| Heat Flux from Earth's Interior | 0.09 | ± 0.006 | 2.60E-04 | | | | | |
| Energy Generated by Lunar Tidal Forces in the Atmosphere | 1.96E-05 | - | 5.80E-08 | | | | | |
| Galactic Cosmic Rays | 8.50E-06 | 7.0E-06 to 1.0E-05 | 2.50E-08 | | | | | |
| Total Radiation from Stars | 6.78E-06 | 5.62E-06 to 7.94E-06 | 2.00E-08 | | | | | |
| Cosmic Microwave Radiation Background | 3.13E-06 | ±2.62E-09 | 9.20E-09 | | | | | |
| Dissipation of Mechanical Energy from Micrometeorites | 1.10E-06 | 1.9E-08 to 2.0E- 06 | 3.20E-09 | | | | | |
| | | | | / | | | | |
| Total of All Secondary Energy Sources | 0.1169 | | 3.39E-04 | | | | | |
| * global average | | | | | | | | |
| Greenhouse gases are not an energy source. | | | | | | | | |
| from "Where does Earth's atmosphere get its energy?" by A.C. Kren, P. Pilewskie, and O. Coddington, Space Weather | | | | | | | | |
| and Space Climate, 2017 | | | | | | | | |

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Why You Might Expect the Sun to Affect Earth's Climate – and Why It Doesn't (Much)

Fortunately, for an 800-lb gorilla, the Sun is very placid

| | Heat Flux* | Uncertainty or Range | | | | |
|--|----------------------|----------------------|----------------|--|--|--|
| Energy Source | [W m ⁻²] | [W m ⁻²] | Relative Input | | | |
| Solar Irradiance | 340.2 | 0.0000% | 1.000E+00 | | | |
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| * global average | | | | | | |
| | | | | | | |
| rom "Where does Earth's atmosphere get its energy?" by A.C. Kren, P. Pilewskie, and O. Coddington, Space Weather | | | | | | |

from "Where does Earth's atmosphere get its energy?" by A.C. Kren, P. Pilewskie, and O. Coddington, *Space Weather* and Space Climate, 2017



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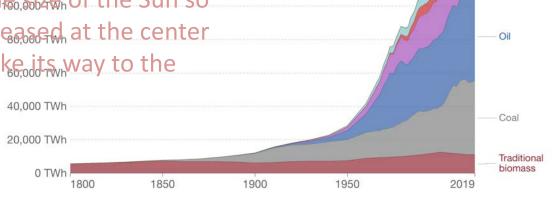
The Sun's Energy

Source: Vaclav Smil (2017) & BP Statistical Review of World Energy

losses as fossil fuels

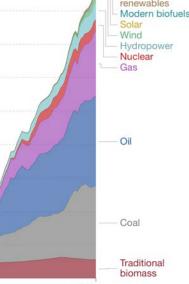
140.000 TWh

- Energy output
 - Energy heating the Earth 1361 W/m²
 - Sun's total output 3.8x10²³ kW
- Energy trivia
 - The total output of the Sun in one second would provide the world with enough energy, at its current usage rate, for the next ~600,000 years Global primary energy consumption by source Primary energy is calculated based on the 'substitution method' which takes account of the inefficienci
- Unfathomable trivia
 - The Sun converts 6x10¹¹ kg of H into He^{160,000} TWh second
- More fun (and unfathomable) trivi
 - The core is so dense (150 g/cm³) and the size of the Sun so great (7x10¹⁰ cm radius) that energy released at the center takes 100,000 to 1,000,000 years to make its way to the surface



convection zone

radiative zone core Our World in Data fuel production by converting non-fossil energy into the energy inputs required if they had the same conversior Other renewables



OurWorldInData.org/energy · CC BY

Means of Energy Input from the Sun

electromagnetic radiation (solar irradiance) 1361 W/m²

> energetic particles (protons, electrons) ~ 10⁻⁴ W/m²

> > solar wind (plasma) ~ 10⁻⁴ W/m²

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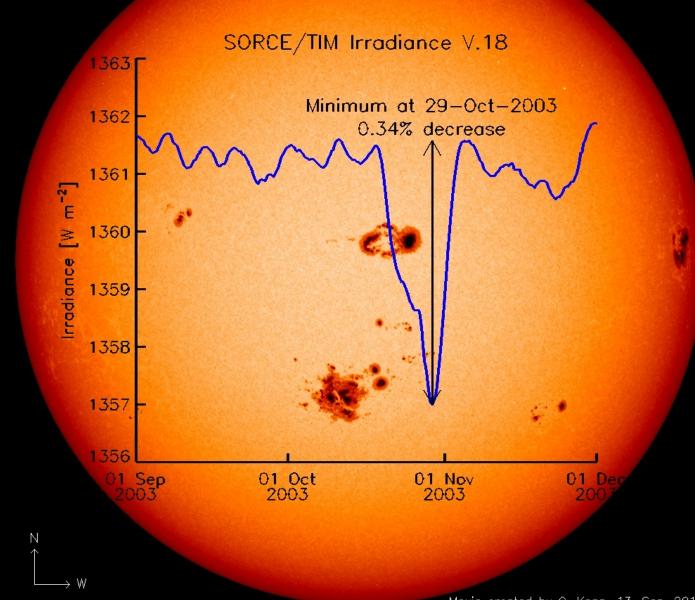
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galactic cosmic rays (particles)

10⁻⁵ W/m²

MDI Intensitygram

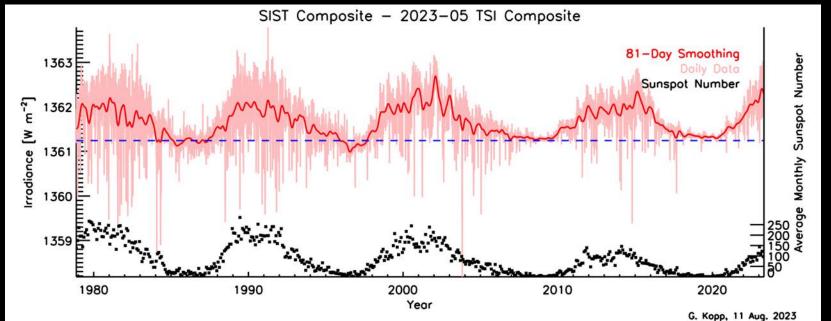
29-Oct-2003





The Earth's Incoming Energy

- TSI provides **99.974%** of Earth's incoming energy
 - Mean TSI: 1361 \pm 0.5 W m $^{\text{-2}}$
 - = $340.2 \pm 0.1 \text{ W m}^{-2}$ in "today's talk units" (not an SI unit)
 - Accepted by IAU
- "Community-Consensus TSI Composite" provides time series and updated absolute values
- SSI determines where energy is absorbed and thus affects outgoing radiation
- The "solar constant" is not a constant





Where Does Energy Go?

• Blackbody temperature of Earth

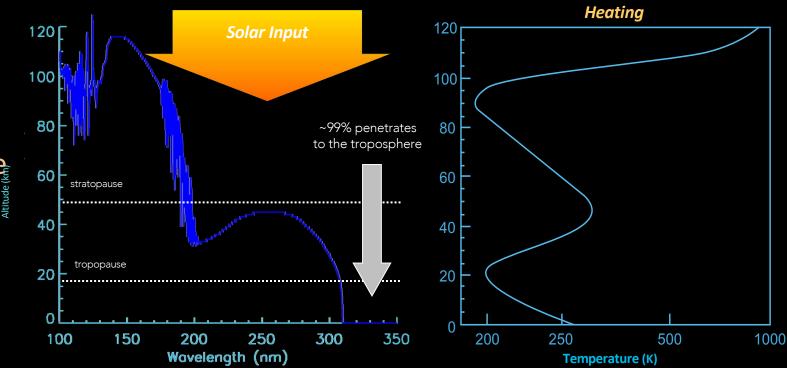
Forcing sensitivity κ

 $\Delta T = \kappa \Delta S$

- < ¹/₂ is absorbed by surface
- 1/3 is reflected
- ¼ is absorbed by atmosphere
- Feedback mechanisms

Incoming Energy =
$$\pi R^2 \cdot (1 - A) \cdot S$$

Outgoing Energy = $4\pi R^2 \cdot \varepsilon \cdot \sigma T^4$
Energy Balance $\Rightarrow T = \sqrt[4]{\frac{1-A}{\varepsilon} \frac{1}{4\sigma}S} = 280 \text{ K}$



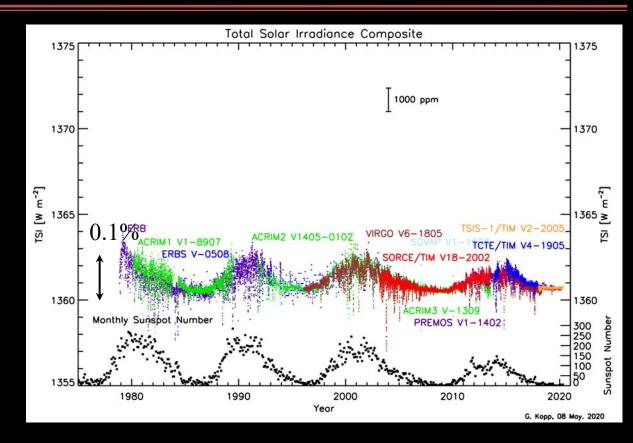




What Are the Timescales of Solar Variability?

- 0.01% over minutes
- <0.3% over a few days
 - Short duration causes negligible climate effect
- 0.1% over 11-year solar cycle
 - Small but detectable effect on climate
- 0.05-0.3% over centuries (unknown)
 - Direct effect on climate (Maunder Minimum and Europe's Little Ice Age)
- 10⁻¹⁰/yr on evolutionary timescales

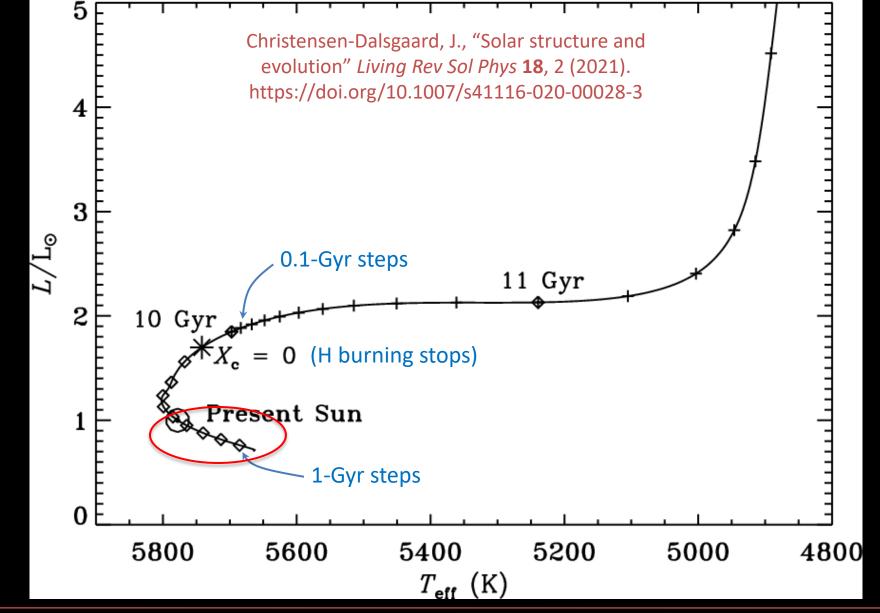
<u>Misnomer</u> It is <u>not</u> the "solar constant"



- An unequivocal link between climate change and TSI has been established
 - Magnitude of natural climate forcing needs to be known for setting present and future climate policy regulating anthropogenic forcings



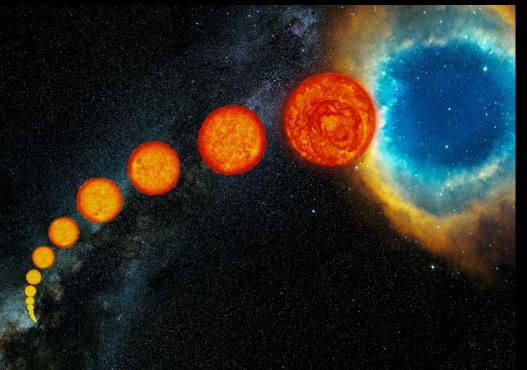
Solar Evolution and Faint Young Sun

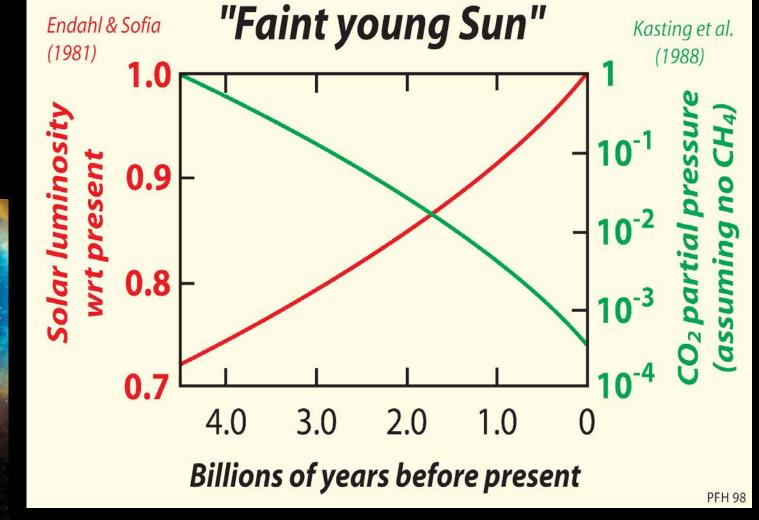




Faint Young Sun Paradox

- How did liquid H₂O exist?
 - Greenhouse gases
 - Radiogenic heating
 - Tidal heating



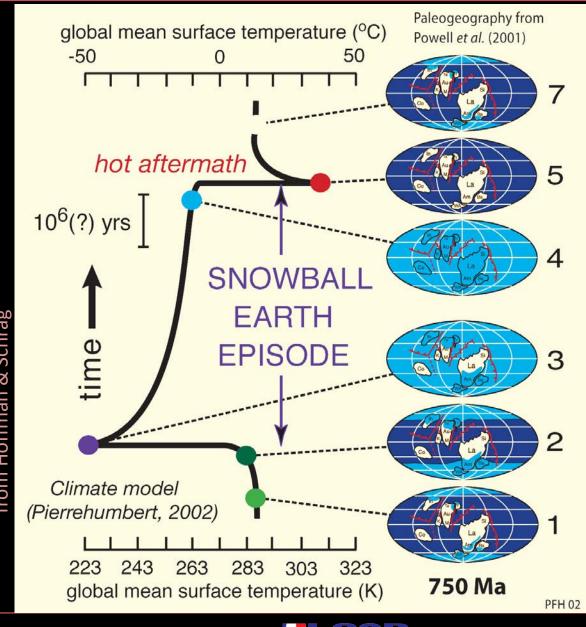




Snow/Slush-Ball Earth Cycle

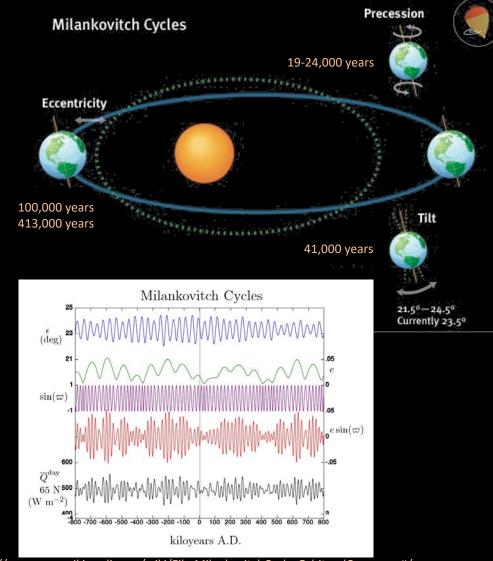
- Polar-ice extent has negative CO₂ feedback
 - Exposed areas are CO₂ sinks: silicate weathering, CO₂ ocean-uptake, photosynthesis
 - CO₂ sources: volcanoes & geothermal
- Ice extent has positive albedo feedback
 - Once more than ½ Earth's surface becomes ice covered, albedo feedback is unstoppable →
 "Snowball Earth"
- Volcanic activity continues to release CO₂ → Greenhouse gas warming
- Once ice melts in tropics, continued melt is unstoppable → warm, ice-free Earth
- Increased silicate weathering and photosynthesis \rightarrow lower CO₂ \rightarrow cooling

~4 periods 750 million to 580 million years ago

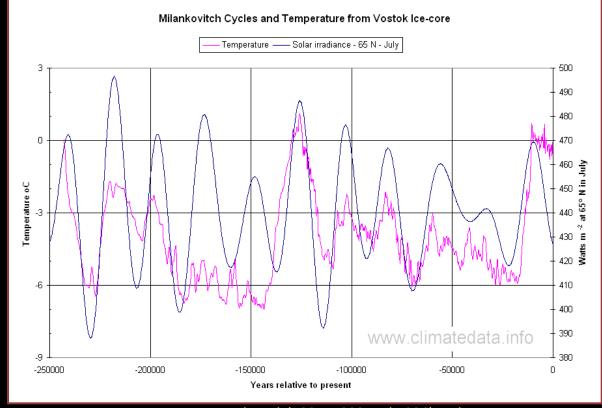


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Milankovitch Cycles Cause Direct Radiative Forcing



http://commons.wikimedia.org/wiki/File:MilankovitchCyclesOrbitandCores.png#/me dia/File:MilankovitchCyclesOrbitandCores.png "Observations show climate behavior is much more intense than the calculated variations." Joanna Haigh, "Solar Influences on Climate," 2011



Jouzel *et al*. (1987, 1993 and 1996) and Wahlen, M.*, et al*, (2000), Vostok Ice Core CO₂ Data, 1105-2856 m

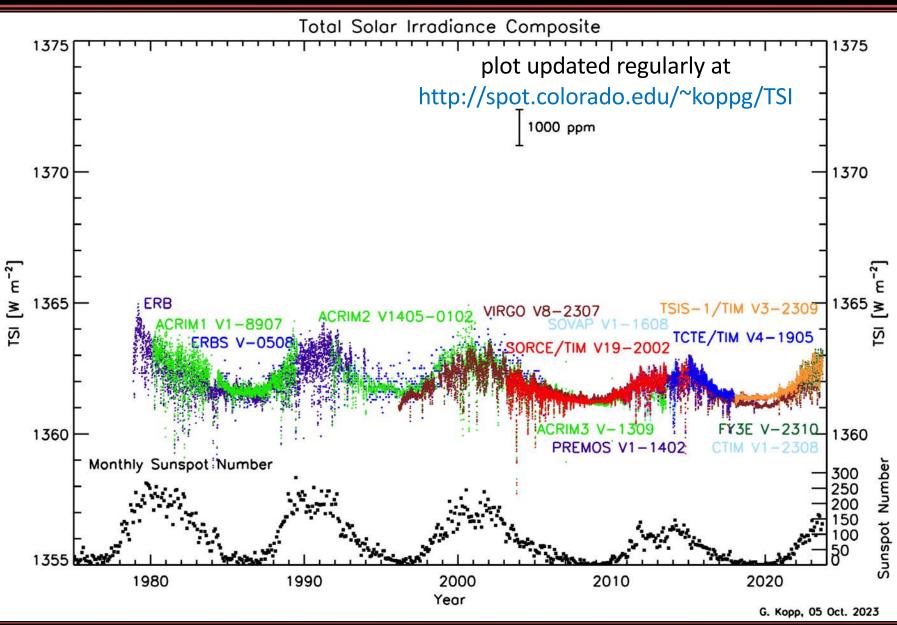


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The Current Total Solar Irradiance Record

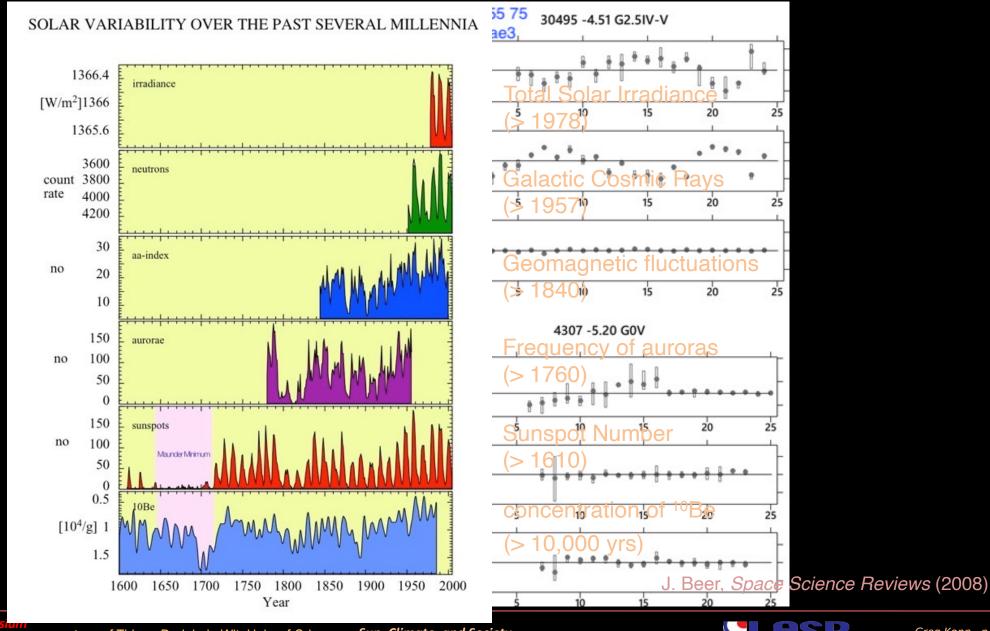
- The uninterrupted, 45year-long spaceborne TSI record includes contributions from more than 15 NASA, ESA, and NOAA instruments
 - Improvements have been made to absolute accuracy
 - The record continues to rely on continuity and stability







Proxies of Solar Activity

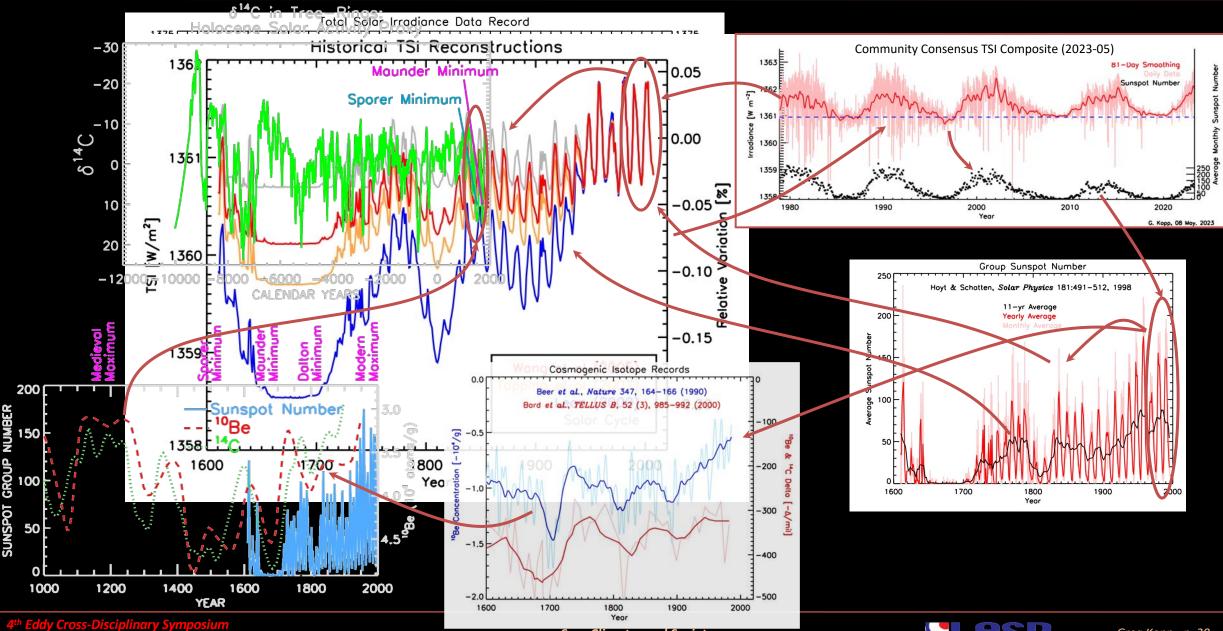


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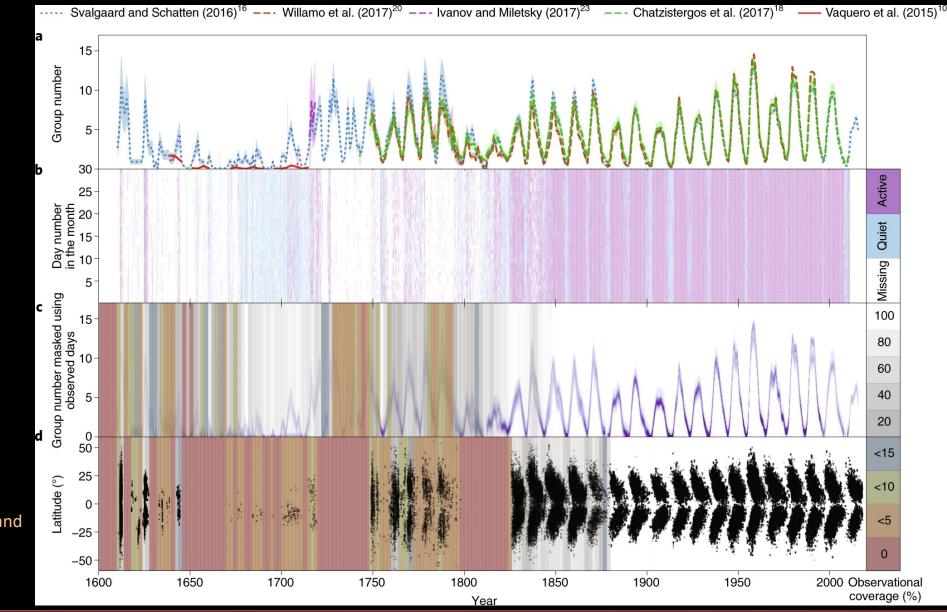
Proxies Reconstruct Historical TSI Based on Measurement Record



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400-Year Sunspot Record (and Reconstructions)



Muñoz-Jaramillo and Vaquero, Nature Astronomy, 2019

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Sunspots Are Linked to Climate

1610-1801 - Explanations of sunspots

- Galileo Galilei (1564-1642) cloud-like structures in the solar atmosphere
- Christoph Scheiner (1575-1650) intra-Mercurial objects; dense objects embedded in the Sun's luminous atmosphere
- René Descartes (1596-1650) floating aggregates of etheral matter accreted along the Sun's rotational axis, where centrifugal forces are negligible
- William Herschel (1738-1822) & A. Wilson in 1774 openings in the Sun's luminous atmosphere, allowing a view of the underlying, cooler surface of the Sun (which was likely inhabited)



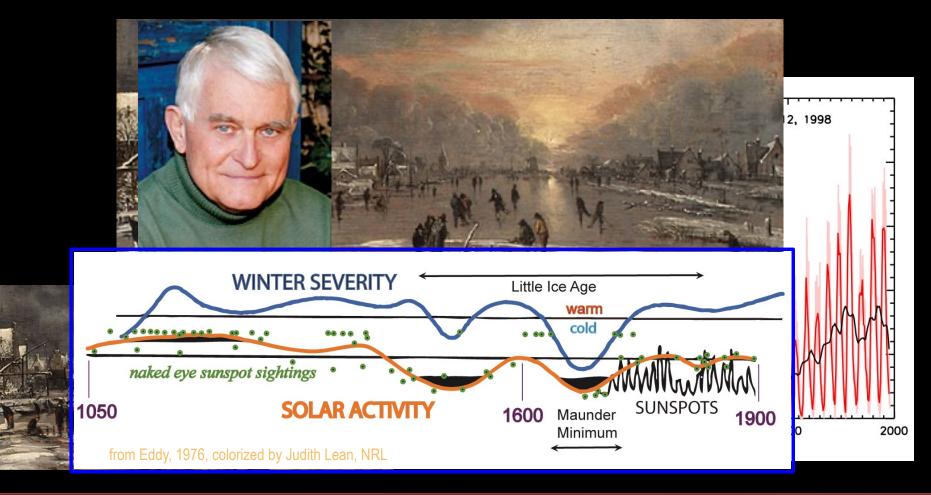
Herschel [1801]: Correlated the price of wheat in London with the number of visible sunspots, attributing the connection to reduced rainfall when the Sun was less spotted



Sunspots – and Europe's Little Ice Age

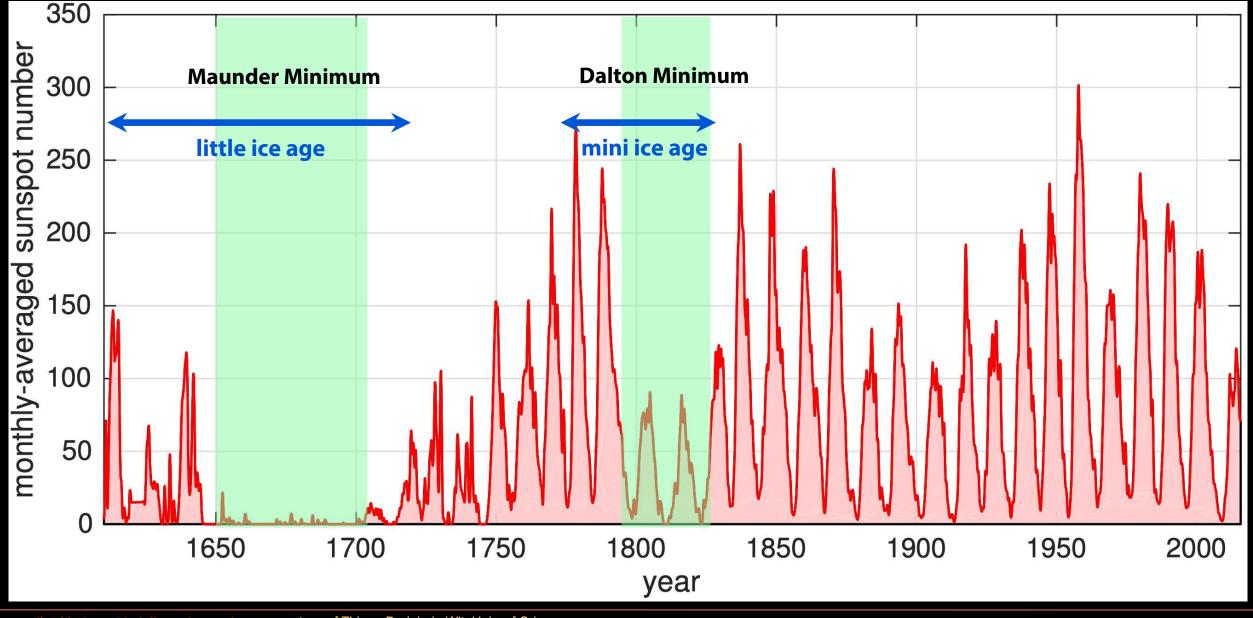
1645-1715 – Maunder Minimum

- Solar output decreased 0.1-0.3% for 70 years
- Earth temperatures were ~0.2-0.4 C colder than the early 1900s





Solar Sunspot Activity



ymposium courtesy of Thierry Dudok de Wit, Univ. of Orleans *Golden, CO*



The Little Ice Age

- Star
 Norse colonies in Greenland starved and vanished by the early 15th century
 - Crop practices throughout **Europe** had to be altered
 - There were many years of *scarcity and famine*
 - Famines in **France** 1693–94, **Norway** 1695–96 and **Sweden** 1696–97 claimed roughly *10 percent of the population* of each country
 - In **Estonia** and **Finland** in 1696–97, losses have been estimated at a *fifth and a third* of the national populations
- 15
 Some storms and flooding resulted in the permanent *loss of large areas* of land from the **Danish, German, and Dutch** coasts
 - Swedish army marched across the Great Belt to Denmark to attack Copenhagen in 1658
 - The colder climate may have caused the wood that was used in Antonio Stradivari's violins to be denser than in warmer periods and contribute to the tone of his instruments

bly triggered or enhanced by olcanic winter y glaciation



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End:

Mau

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Wikipedia



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ISSI Team Reevaluating Sunspot-Number

Team Leaders: Frederic Clette (Belgium) Royal Observatory of Belgium Mathew Owens (UK) University of Reading

Team Members: Rainer Arlt (Germany) Leibniz-Institut für Astrophysik Potsdam Ed Cliver (USA) National Solar Observatory Thierry Dudok De Wit (France) LPC2E Greg Kopp (USA) Univ. of Colorado / LASP Laure Lefèvre (Belgium) Royal Observatory of Belgium Mike Lockwood (UK) University of Reading Andrés Muñoz-Jaramillo (USA) SouthWest Research Institute Ilya Usoskin (Finland) Sodankylä Geophysical Observatory and ReSoLVE Centre of Excellence Lidia Van Driel-Gesztelyi (UK, Hungary, France) University College London José Vaguero (Spain) Departamento de Física

Young Scientist: Theodosios Chatzistergos (Germany) Max-Planck-Institut fur Sonnensystemforschung

Self-supported experts:

Thomas Friedli (Switzerland), Dean Pesnell (USA), and Alexei Pevtsov (USA)

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- 1. Recovered historical data
- 2. Evaluated corrections for specific observers
- 3. Explored combination methodologies

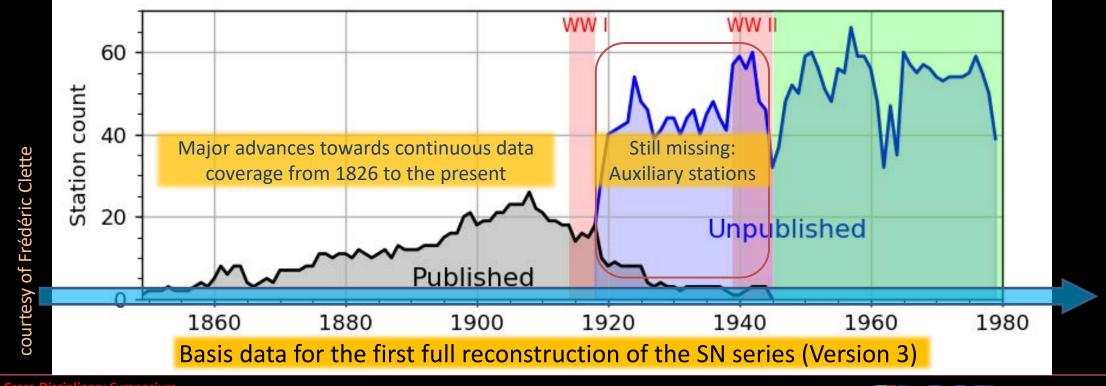
Recalculating TSI reconstructions by incorporating this forthcoming sunspot record into the reconstruction models will be important for estimating historical solar forcing and its uncertainties





Acquired Much More Historical Sunspot Data

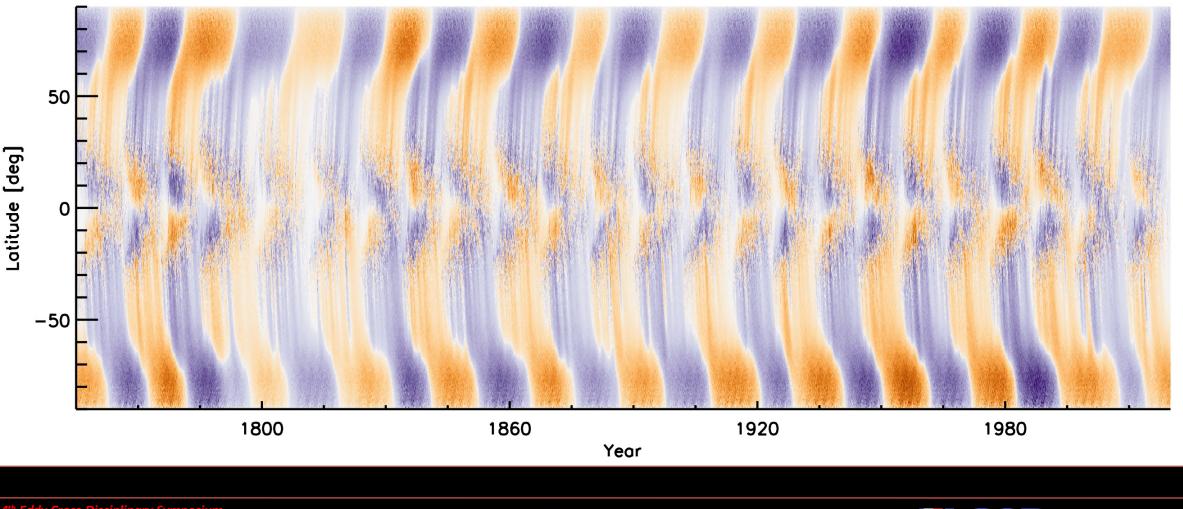
- Long-lost archive of Zürich SN source data from 1945 to 1980 fully recovered in 2018
- Full digitization of the SN source data in the Zürich *Mittheilungen*
- Corrections of misinterpreted GSN data in the Hoyt & Schatten 1998 original database
- Newly-recovered GSN series and drawings in 17th and 18th centuries (incl. observers in Asia)
- New GSN version (V 1.3) (Vaquero, Carrasco, Gallego) and Hisashi data recoveries





Improving the Historical TSI Record Based on Sunspot Records

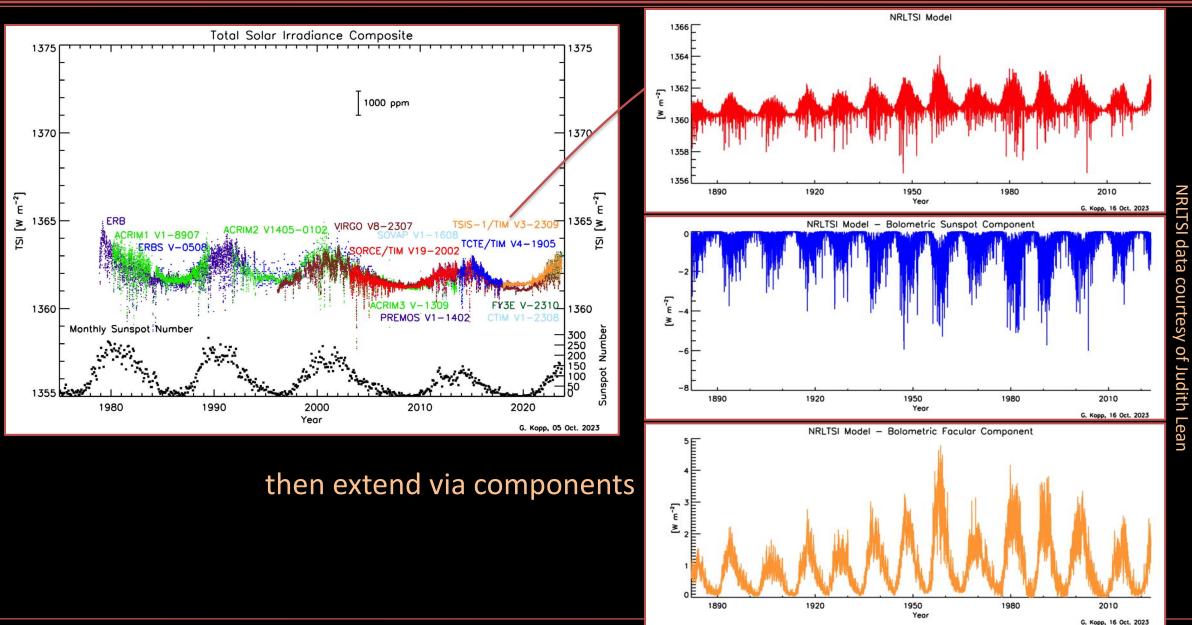
Created continuous simulations of magnetic flux from 1750 to near-present using Advective Flux Transport Model (Hathaway and Upton)



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Regress Solar-Activity Components to TSI Over Measurement Era

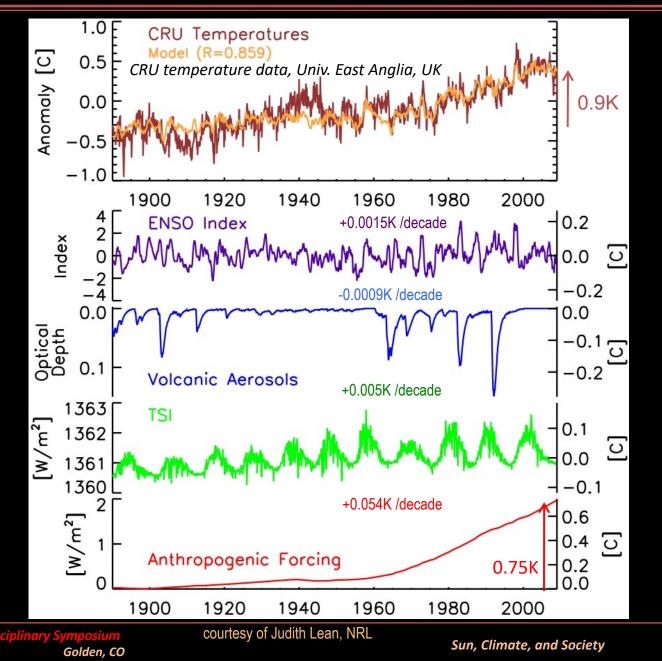


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"Components" of Global Temperature Changes



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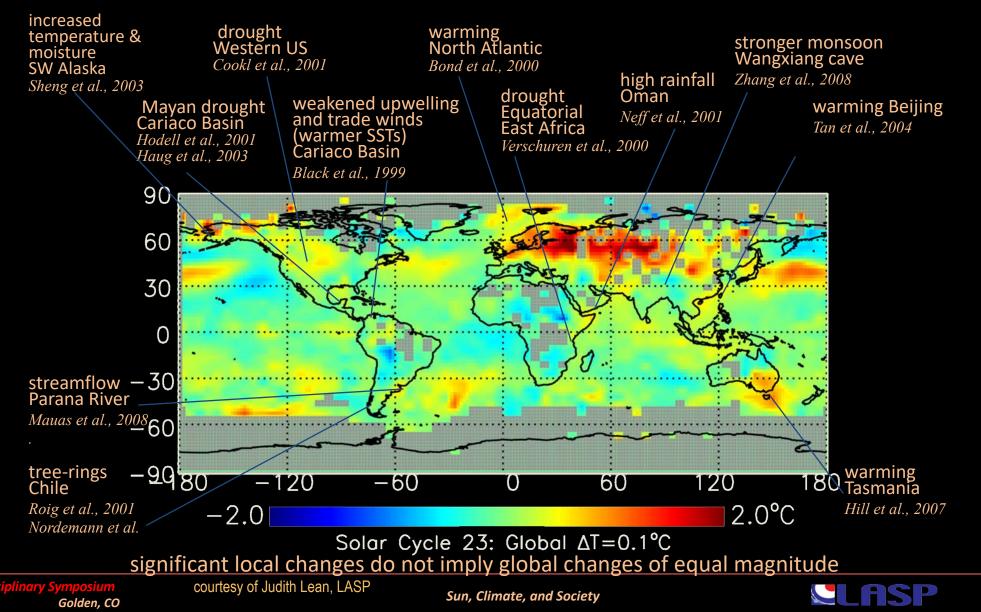
Decompositions of historical and recent global surface temperatures give consistent individual natural and anthropogenic components:

Natural components account for <15% of warming since 1890



Paleo Sun–Climate Synopsis

...when solar activity is high...

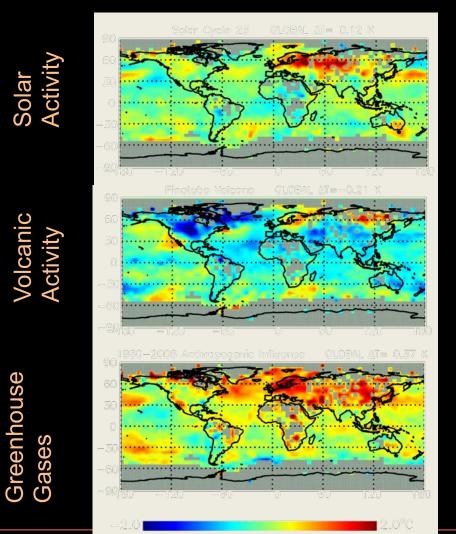


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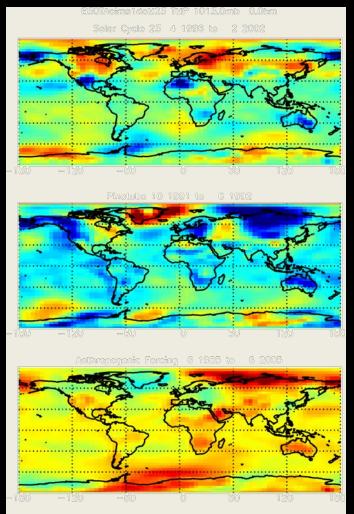
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Regional Temperature Change – Surface

Determined Statistically from Observations *Lean and Rind, GRL, 2008*



Modeled with GISS Model 3 Rind et al., JGR, 2008



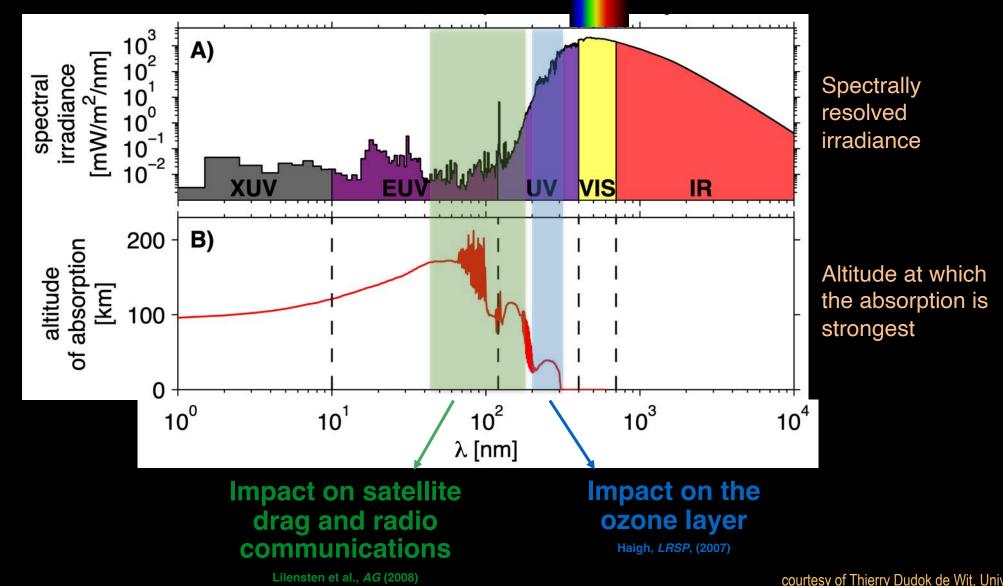
courtesy of Judith Lean, LASP

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Spectral Distribution of Incoming Energy – SSI

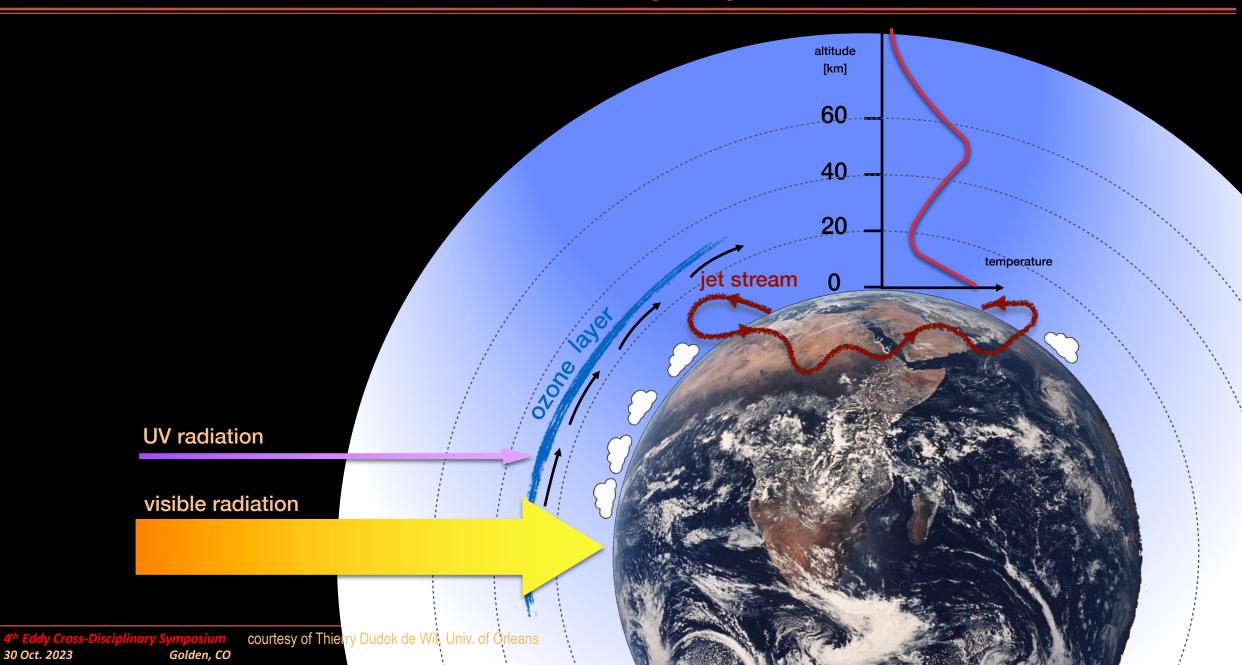


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courtesy of Thierry Dudok de Wit, Univ. of Orleans



How Solar UV Variability Impacts Climate



Solar Influences on Climate

"Bottom Up" (Vis-IR "TSI" Influence : Meehl et al.) Winter pole Summer pole Equator **TSI** THEFTALLSNIERE 201 Scien Planetary Increased Air-sea coupling and Subsidence enhances direct Less clouds solar signal More rain

ncreased Later

Involves solar radiation being absorbed over the oceans

• Produces increases in evaporation

Progress in Earth

et al.,

Seppälä,

 $\overline{\Delta}$

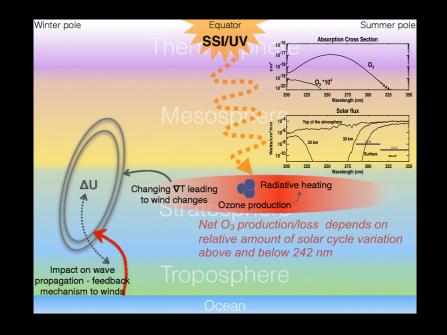
• Increased moisture converging in the precipitation zones

rade winds carry moisture to precipitation zones

This leads to further changes in precipitation patterns and vertical motions, influencing the trade winds and ocean upwelling.

• Produces stronger Hadley and Walker circulations and associated colder sea-surface temperatures at solar maximum

"Top Down" (UV "SSI" Influence : Haigh et al.)



Originates in the stratosphere where UV radiation modulates local radiative heating at the tropical stratopause

- Direct effect on O₃ production rates in upper stratosphere via UV photolysis of $O_2(\lambda < 242 \text{ nm})$
- Radiative heating through O_3 absorption and dissociation (λ >242 nm)

Changes in equatorial stratospheric heating affect meridional temperature gradient

- Modulates the zonal winds
- Influences the planetary wave propagation (positive feedback mechanism)





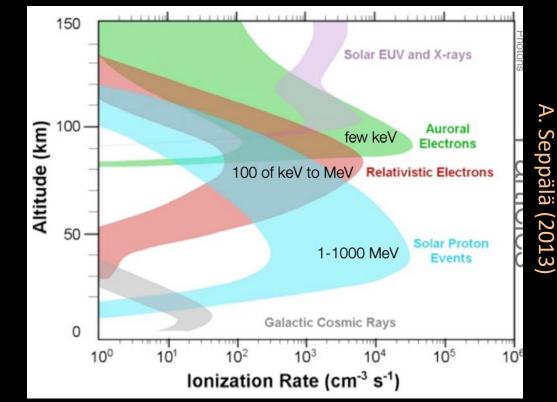
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courtesy of Erik Richard, LASP

Other (Lesser) Drivers

- Galactic cosmic rays and cloud nucleation
- Solar wind and CMEs
 - Solar wind transients store energy in the Earth's magnetosphere
 - Geomagnetic storms accelerate particles and generate intense currents
 - Energetic particles generate NO_{x} in upper atmosphere and alter dynamics
- Solar Flares
 - UV can ionize upper atmosphere and ozone

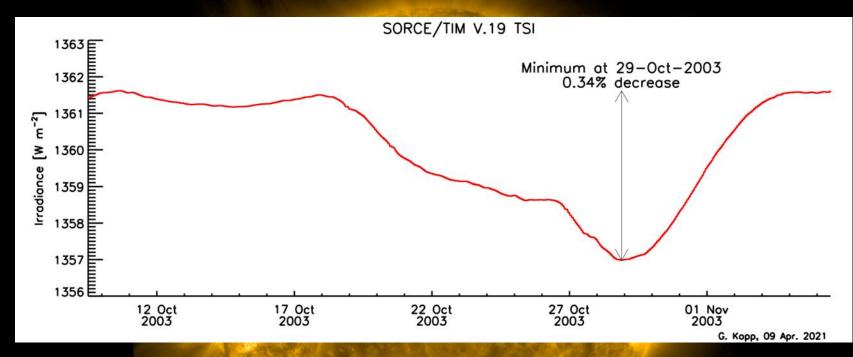
<u>Climate</u> depends on long-term effects (Tomorrow we'll talk about these short-term non-climatic effects)





What About Those Headline-Grabbing Energetic Solar Flares?

- Flares are largest explosive events in the solar system
 - Equivalent to ~ 40 billion
 Hiroshima-type atomic bombs
- The SORCE/TIM recorded the first measurement of a solar flare in TSI
 - This quantified the net radiative energy released by the flare



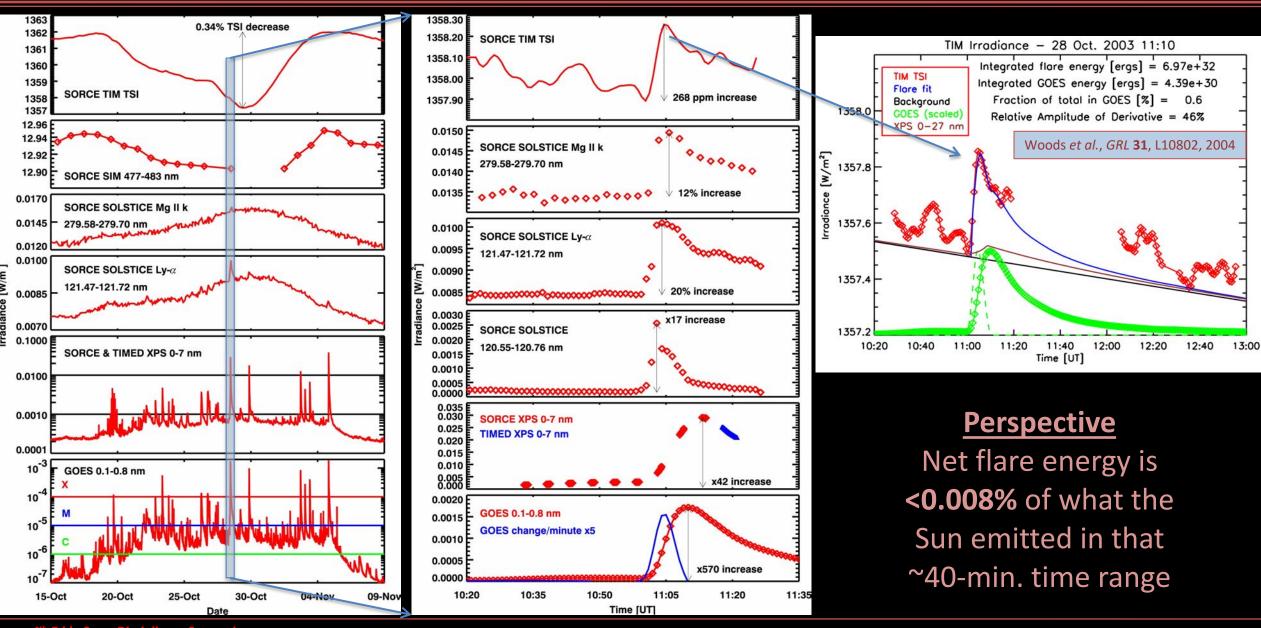
plots available at

http://spot.colorado.edu/~koppg/TSI

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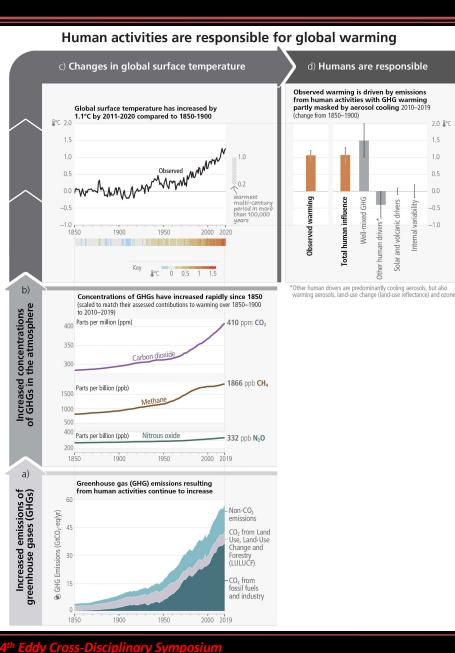
It's Not Flares. Or CMEs. 28 Oct. 2003 X17 SORCE Flare Observations

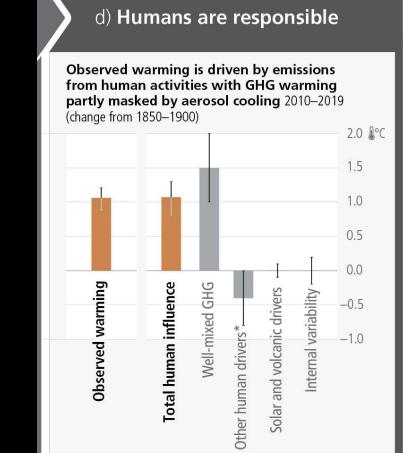


30 Oct. 2023 Golden, CO



IPCC AR6





- It's not the Sun
- But it's important to discriminate natural from anthropogenic effects



Sun, Climate, and Society

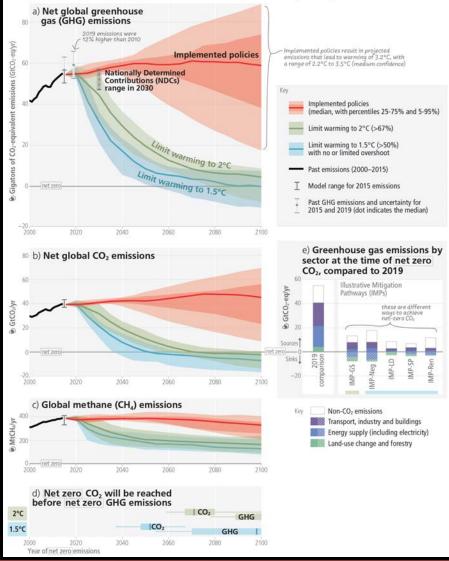
Golden, CO

The Future – <u>Our</u> Future

- Can we limit increase to 2 C?
- CO₂ lasts 1000's of years
 - Ocean uptake and response can be longer
 - We won't get back to pre-industrial environment
- Approaches to climate change
 - Mitigate
 - Adapt (proactively or responsively)
 - Suffer
- (Hint: We'll do all three)

Limiting warming to 1.5°C and 2°C involves rapid, deep and in most cases immediate greenhouse gas emission reductions

Net zero CO2 and net zero GHG emissions can be achieved through strong reductions across all sectors





Who's Involved in This Cross-Disciplinary Sun-Climate Topic?



- Spacecraft and ground-based
- Climate and solar modelers
- Ocean and land monitoring r
- Long-term observers (suns
- Stellar-variability researche
- History buffs, archeologists

1050

- Society
 - Policy makers
 - All nations
 - IPCC
 - Media
 - You!!!

