

NASA 5<sup>th</sup> Eddy Symposium  
UCAR, Boulder CO  
May 05, 2026

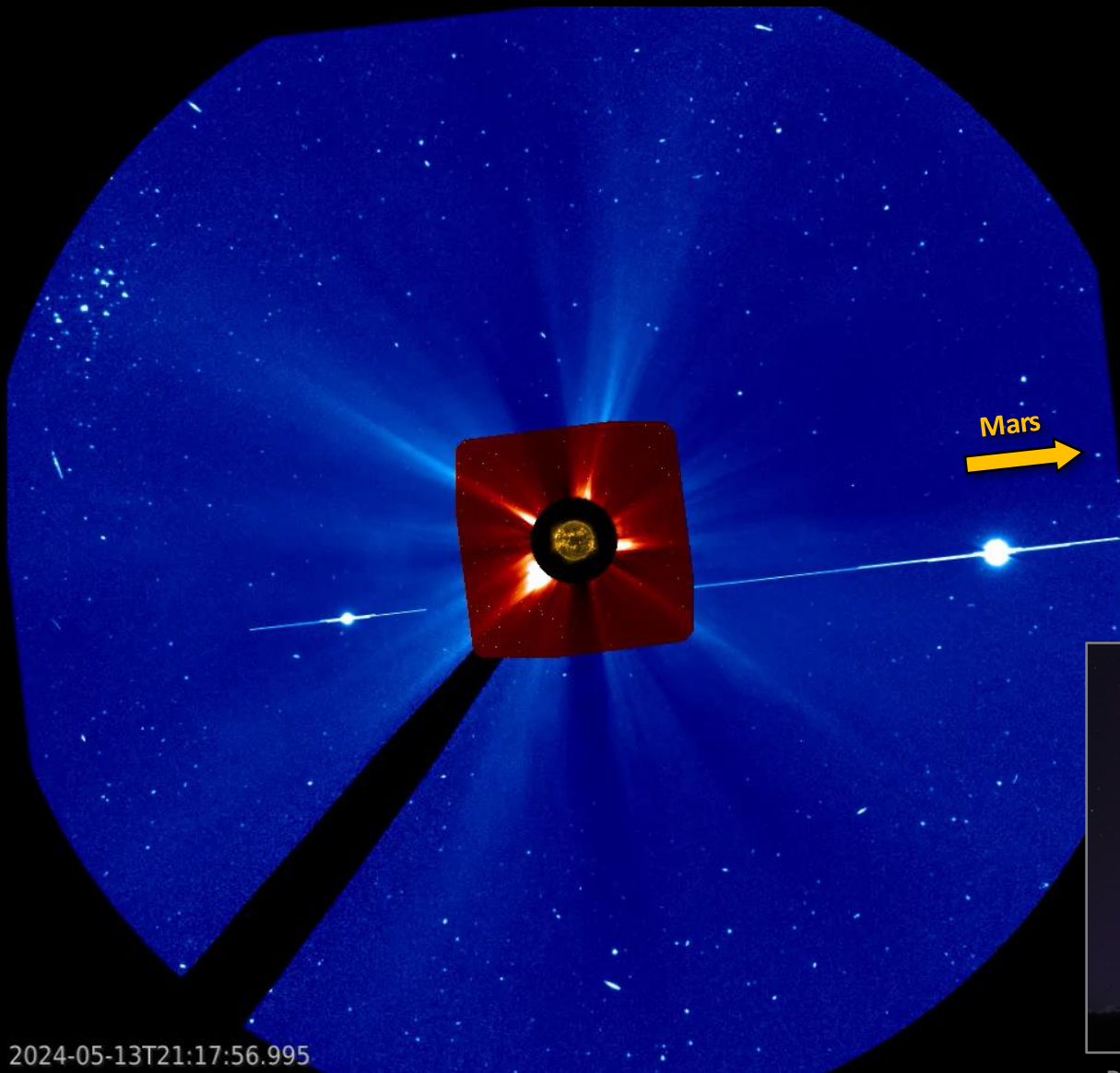
# The Evolving State of the Ion Escape Process During the May 17, 2024 ICME Impact on Mars

ROBIN RAMSTAD, MATS HOLMSTRÖM, JAMES MCFADDEN,  
JARED ESPLEY, DAVID BRAIN, AND SHANNON CURRY



# May 2024 events

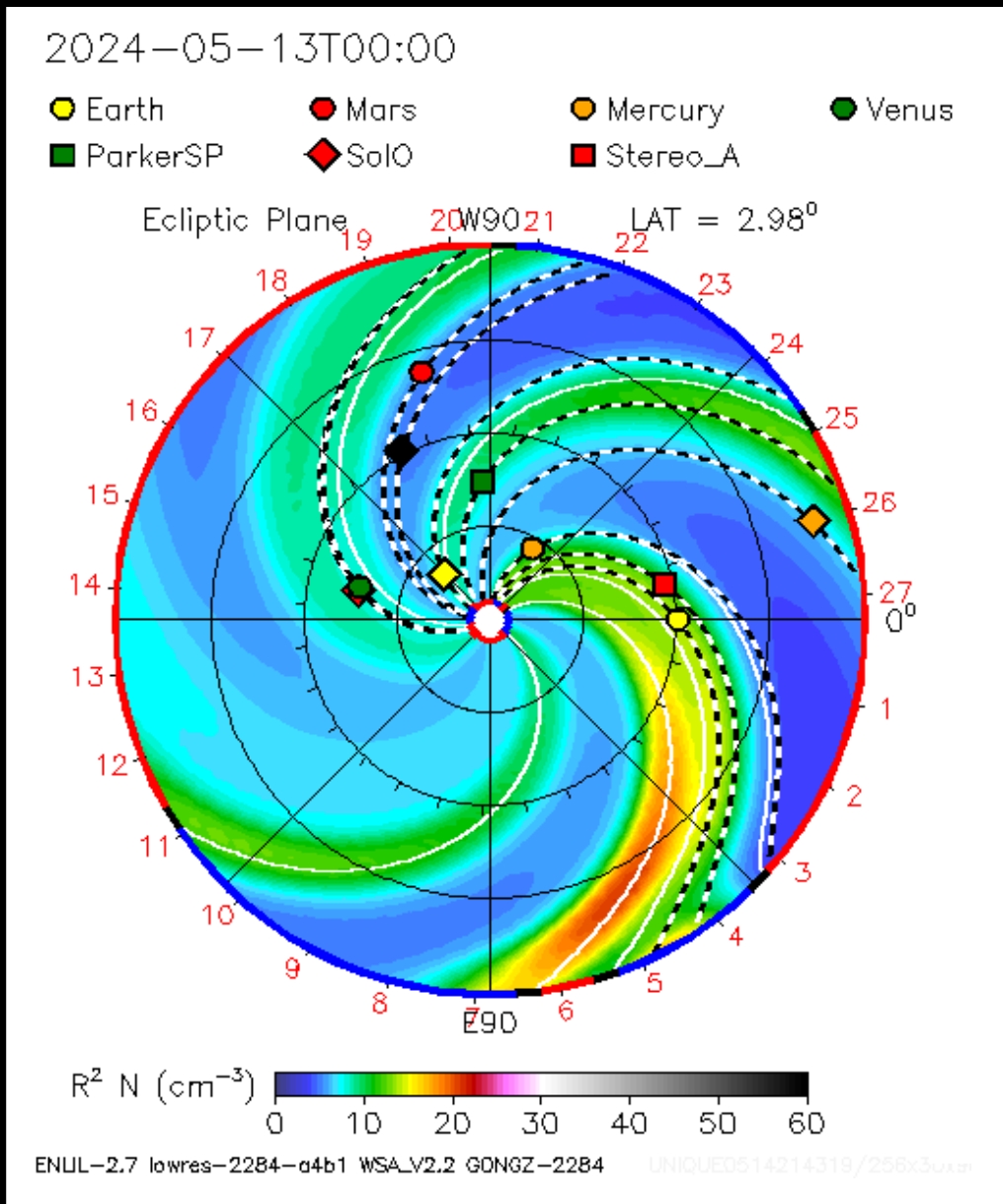
- The Sun produced a series of ICMEs in May 2024
- ICMEs towards Earth created the May 10-13 Gannon storm
- ICMEs also launched towards Mars



2024-05-13T21:17:56.995

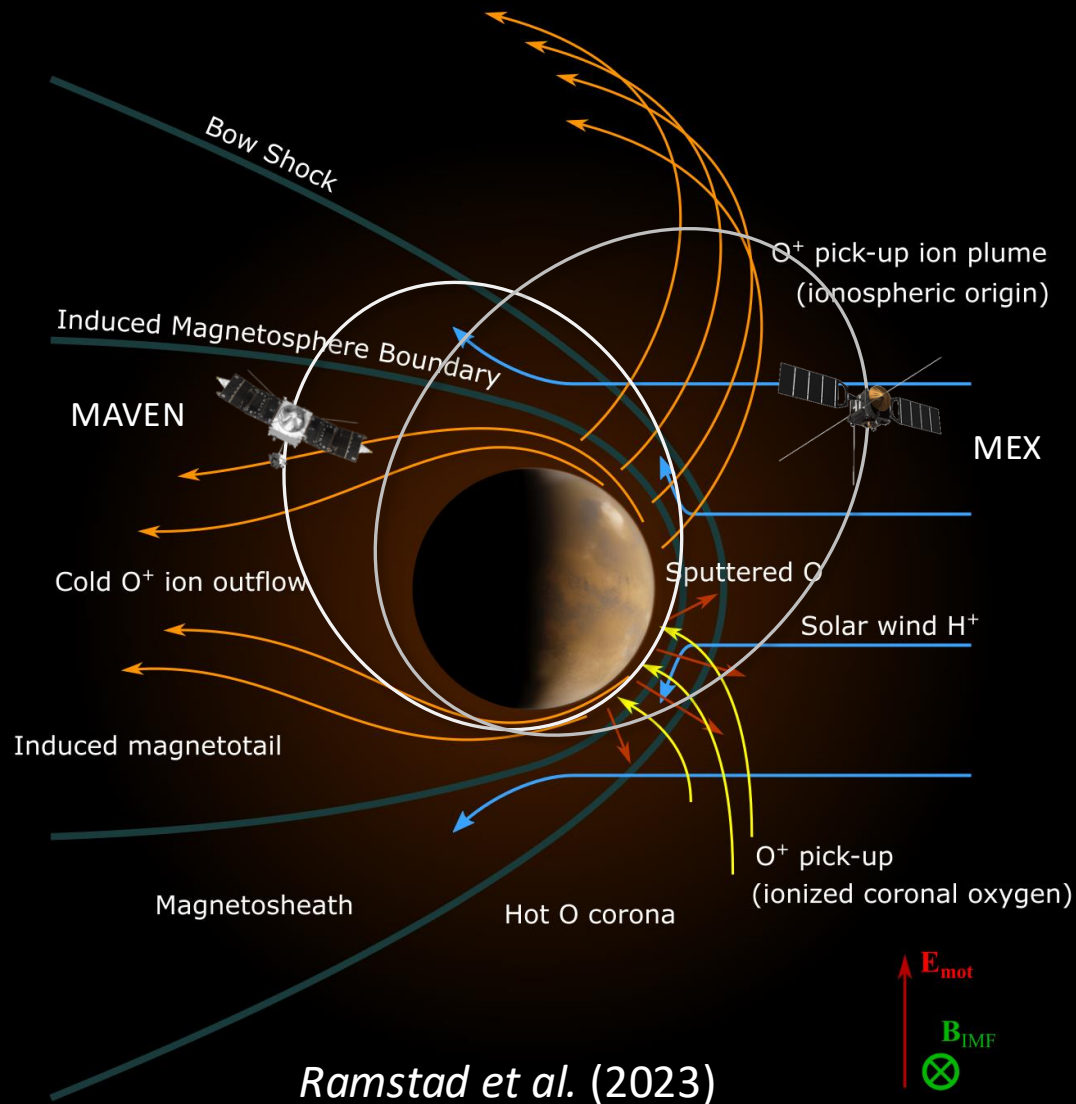
Robin's photo of Gannon storm aurora from Flagstaff mountain in Boulder (May 11)

# May 2024 events



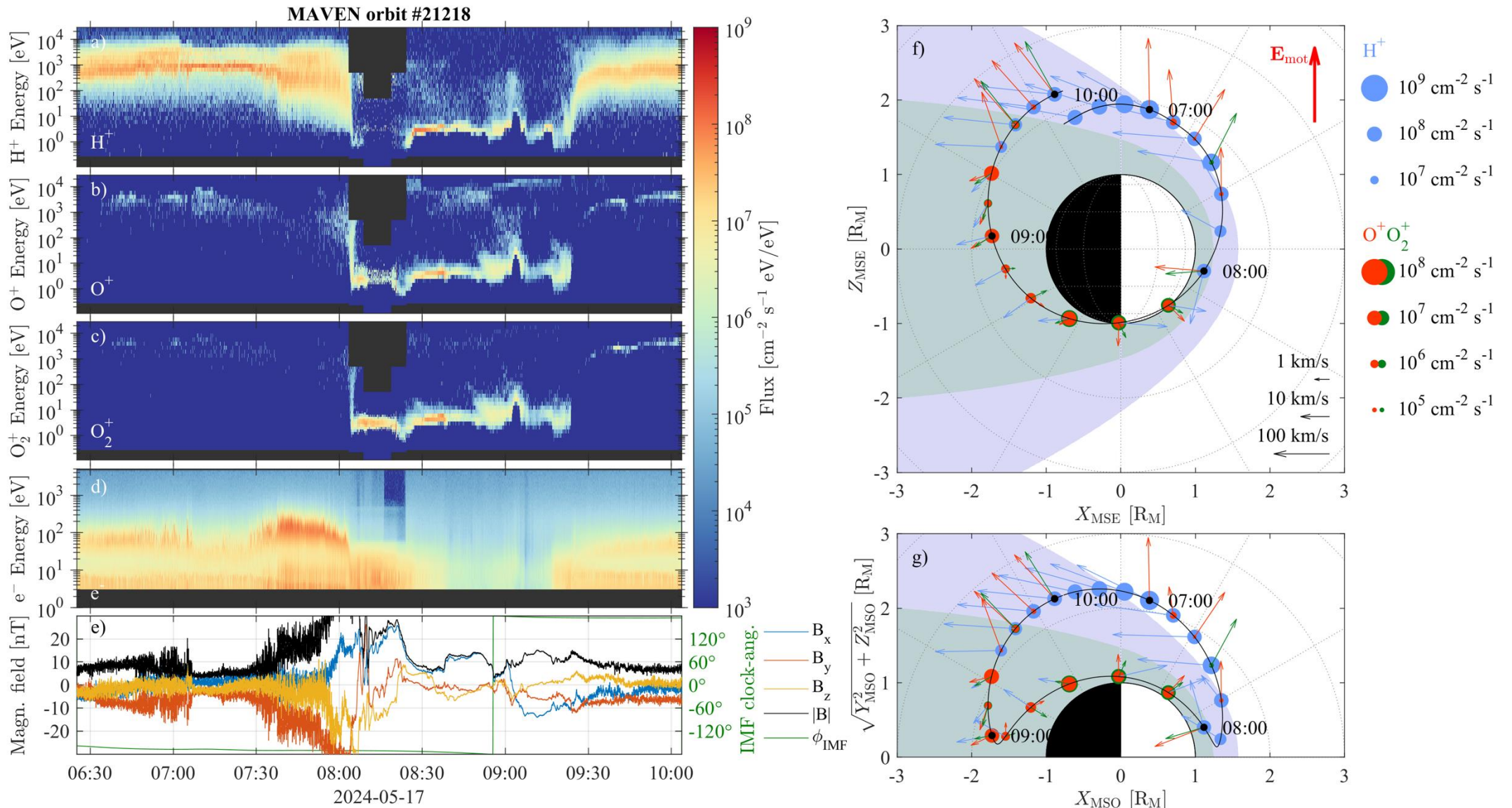
- ICME launched towards Mars on May 14, arrived May 17

# Observations at Mars

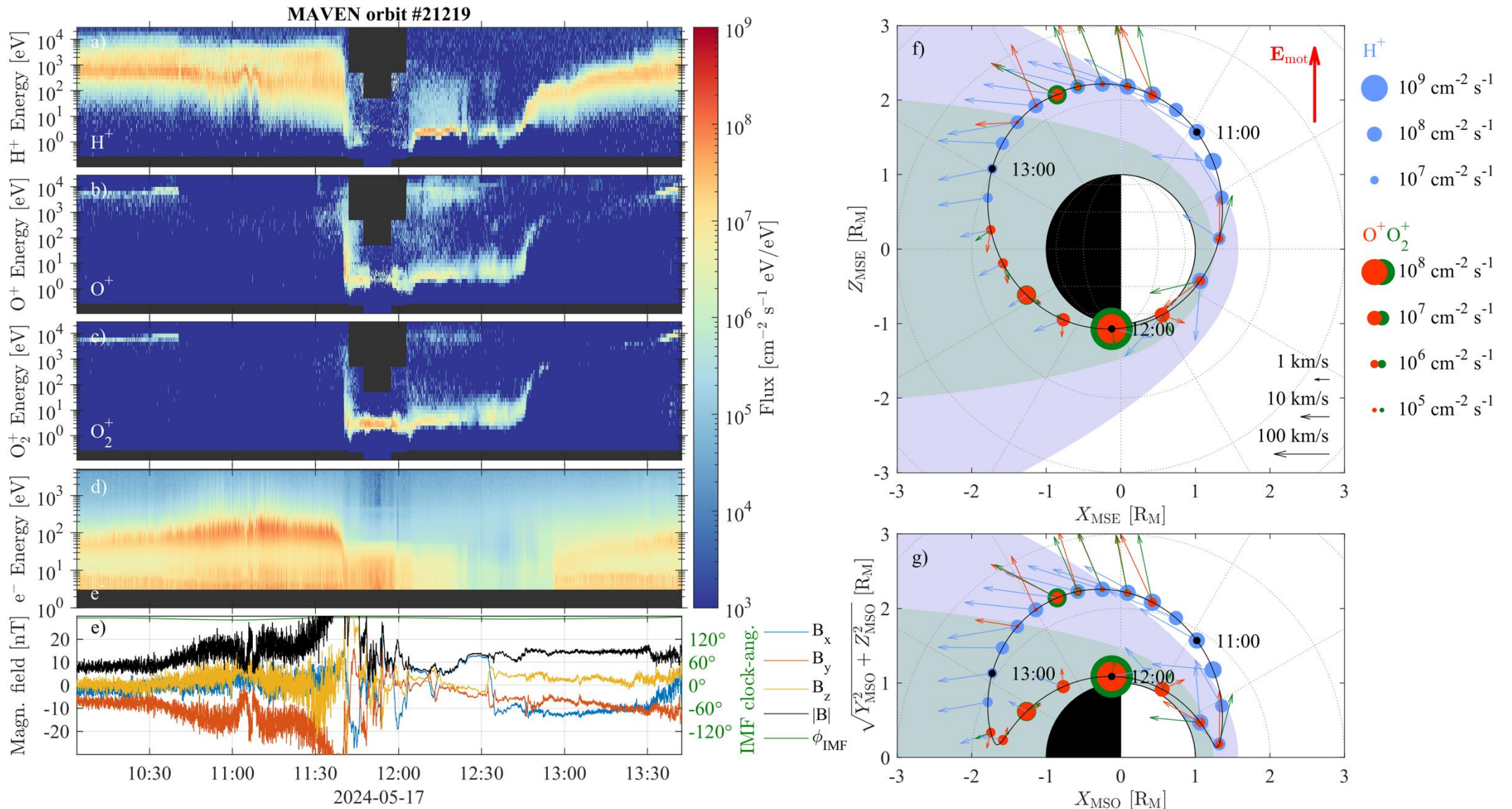


- At Mars, ~80% of atmospheric ions escape through the tail
- MAVEN's orbit was ideal for measuring ions escaping through the deep tail
- Mars Express' orbit was ideal for simultaneously measuring the upstream solar wind

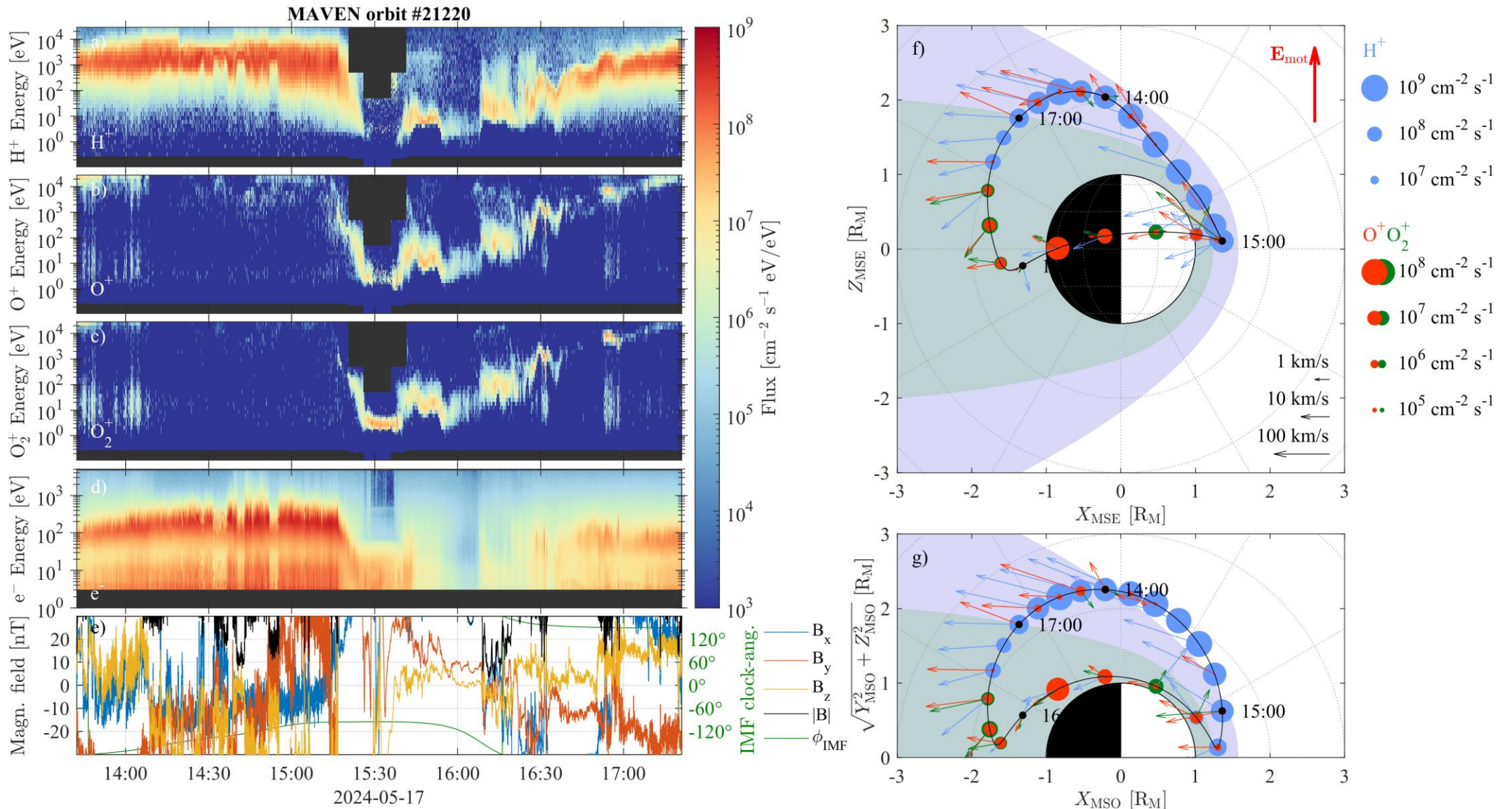
# Before impact



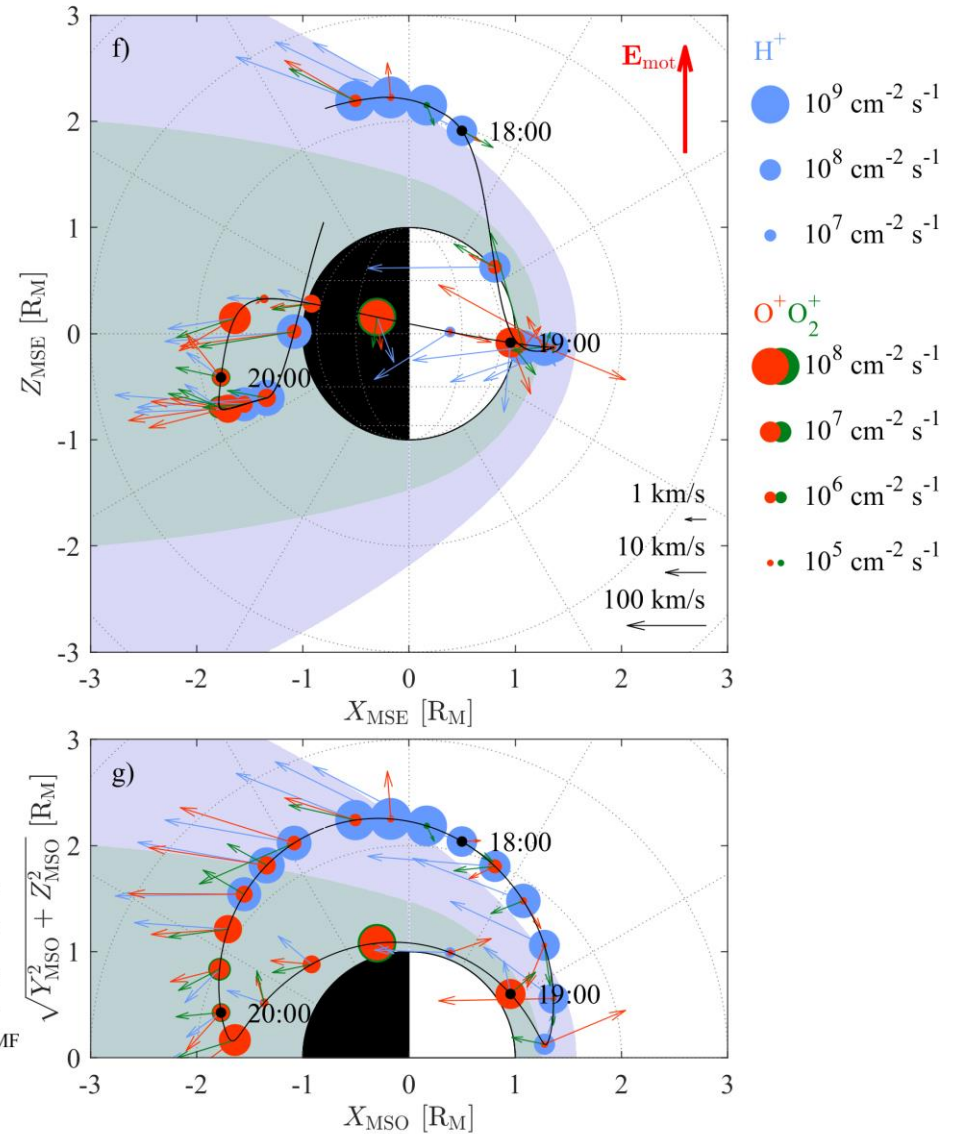
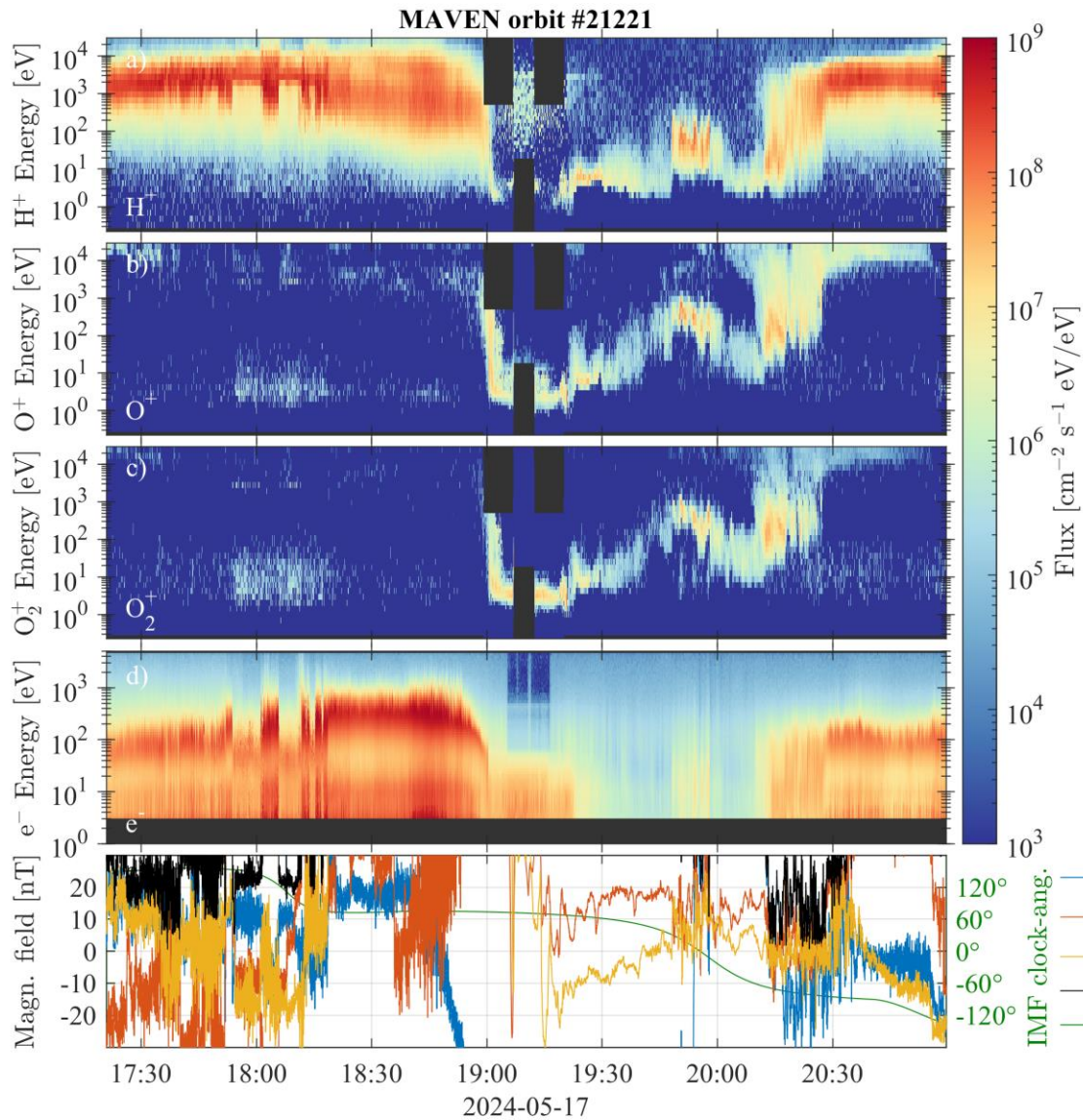
# Before impact



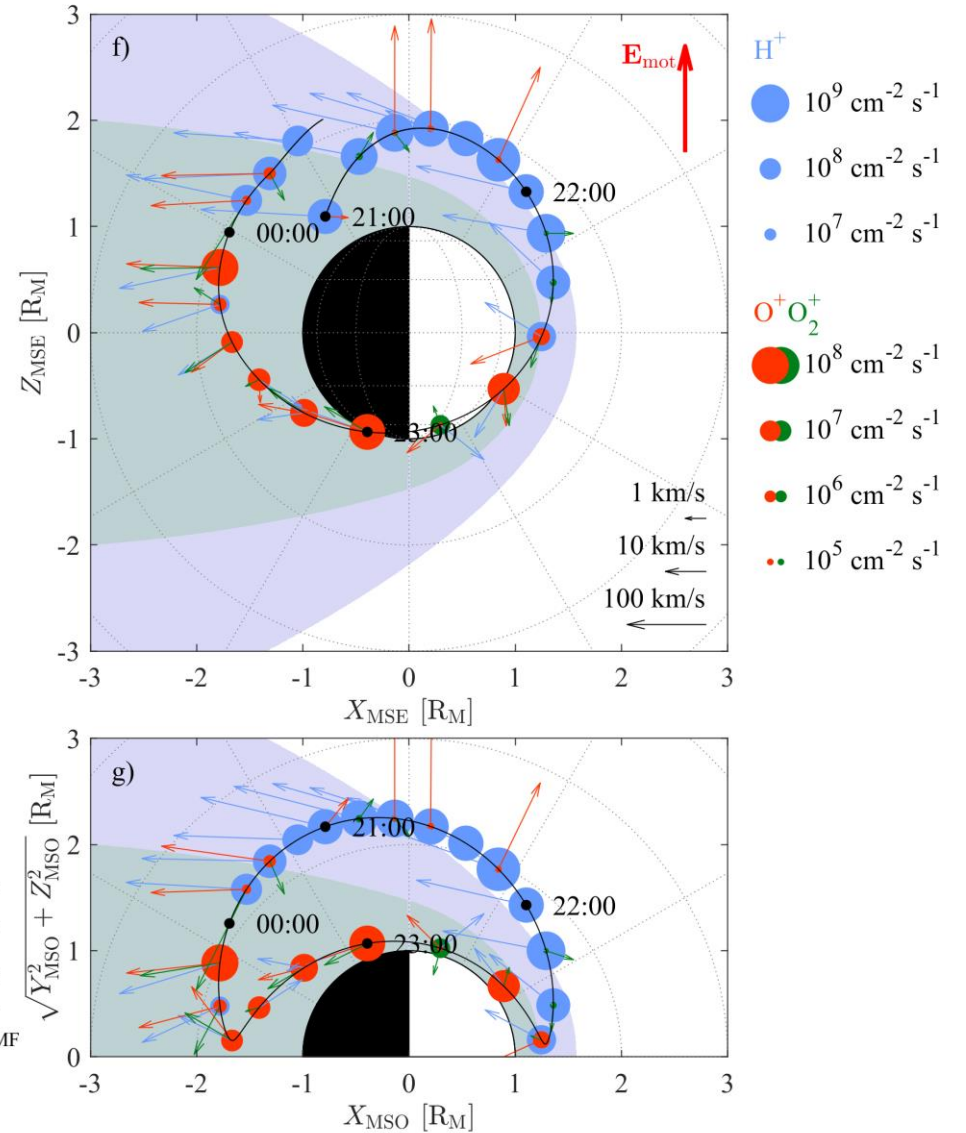
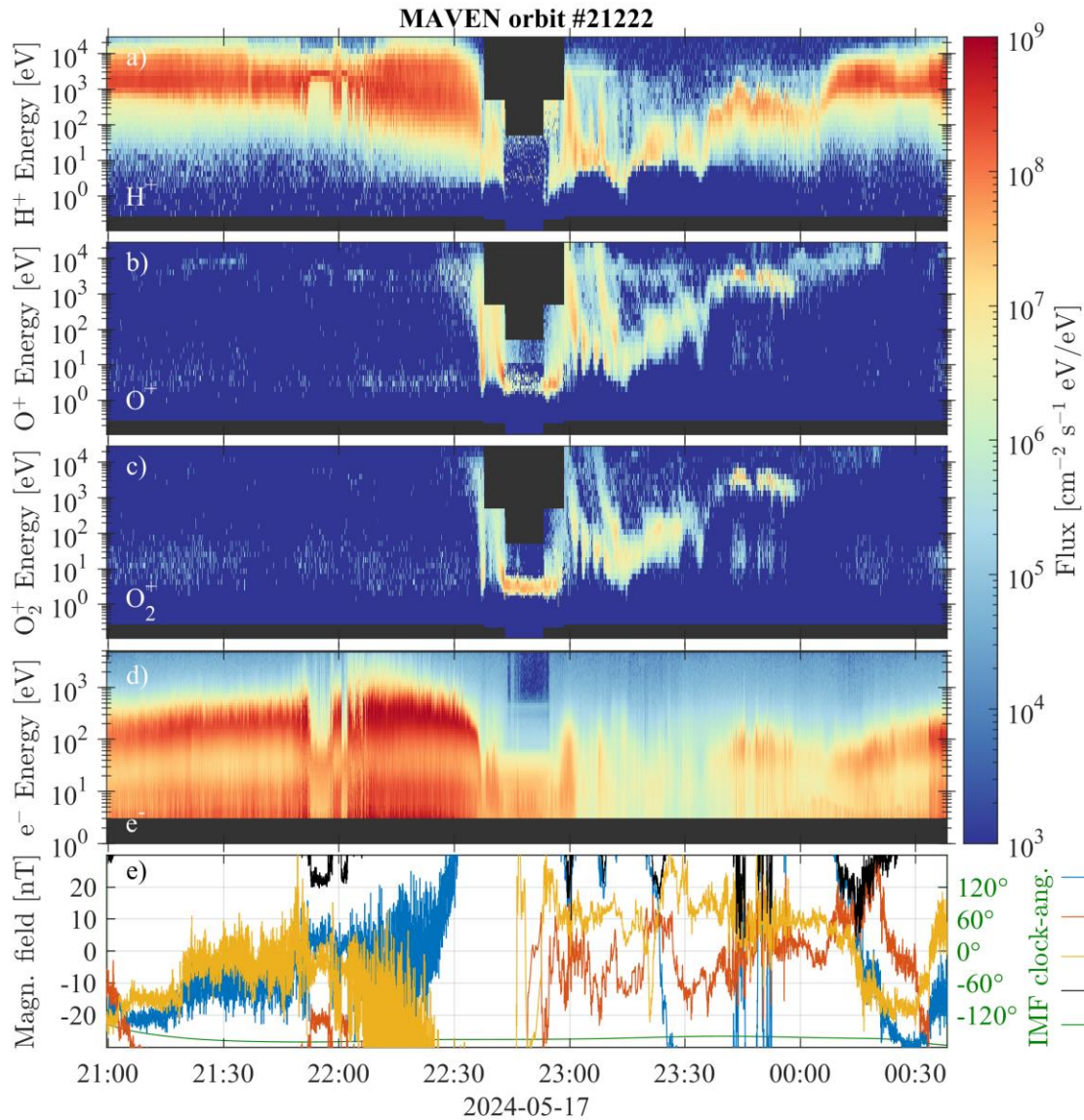
# Impact



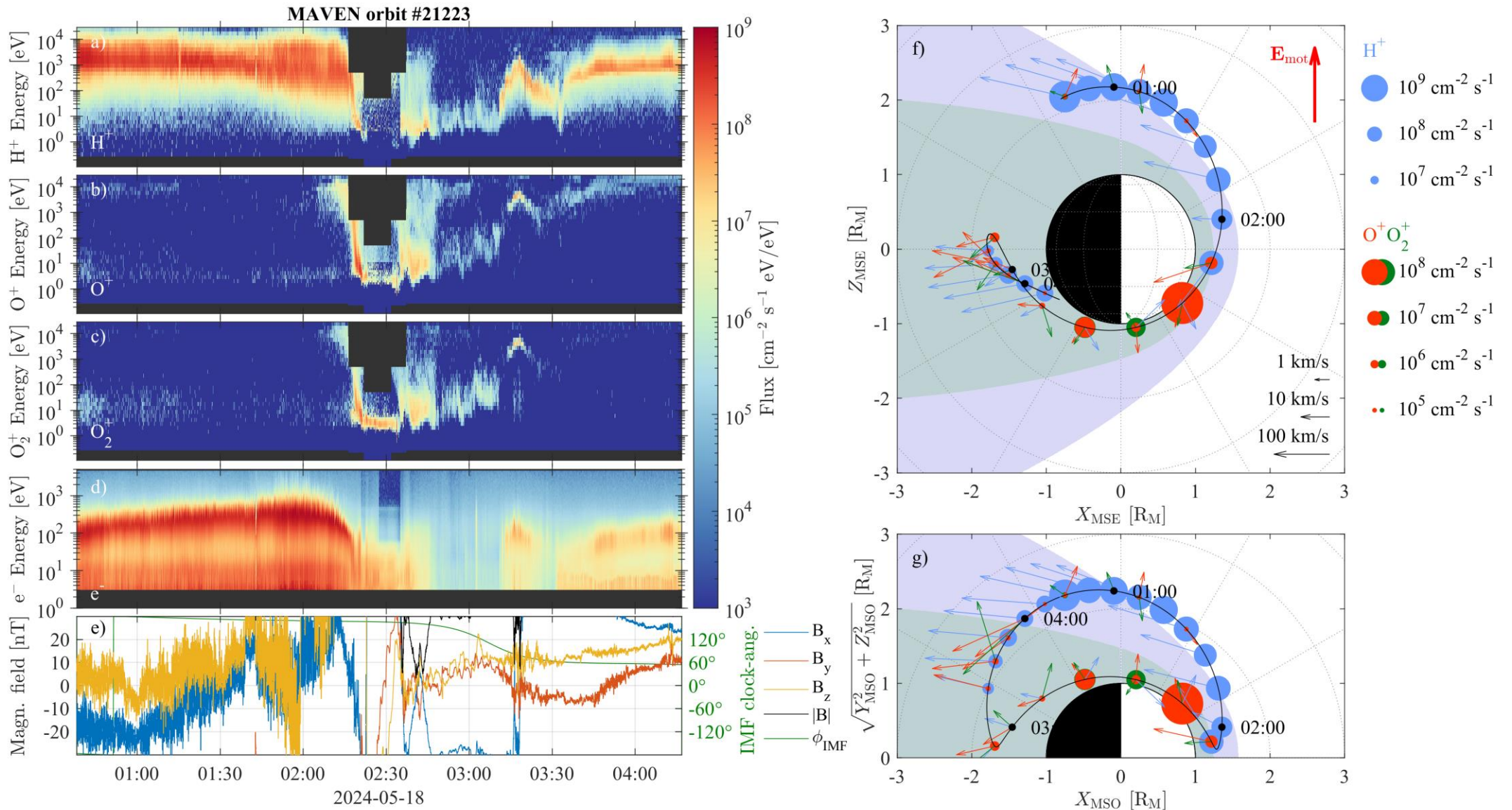
# ICME



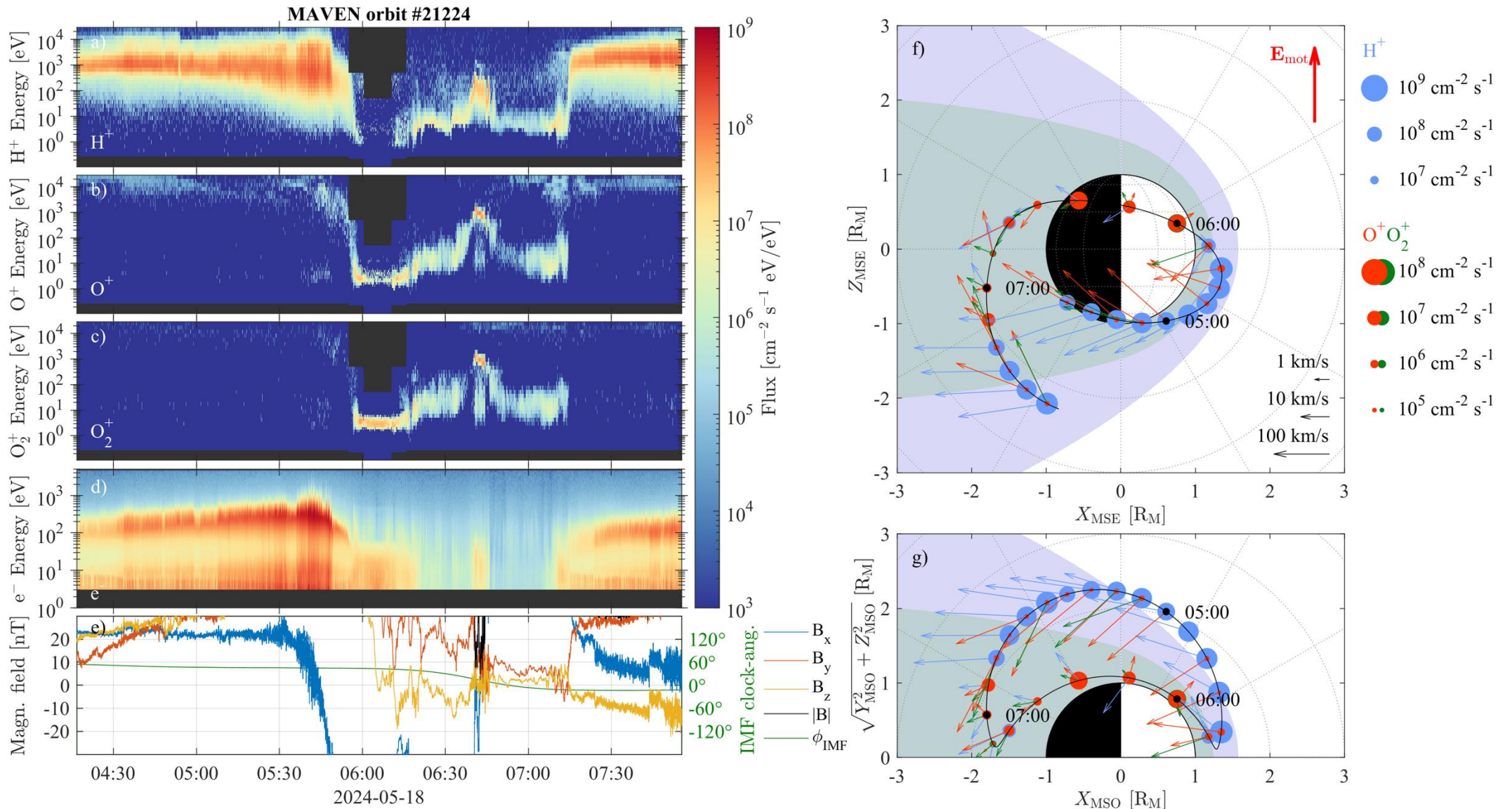
# ICME



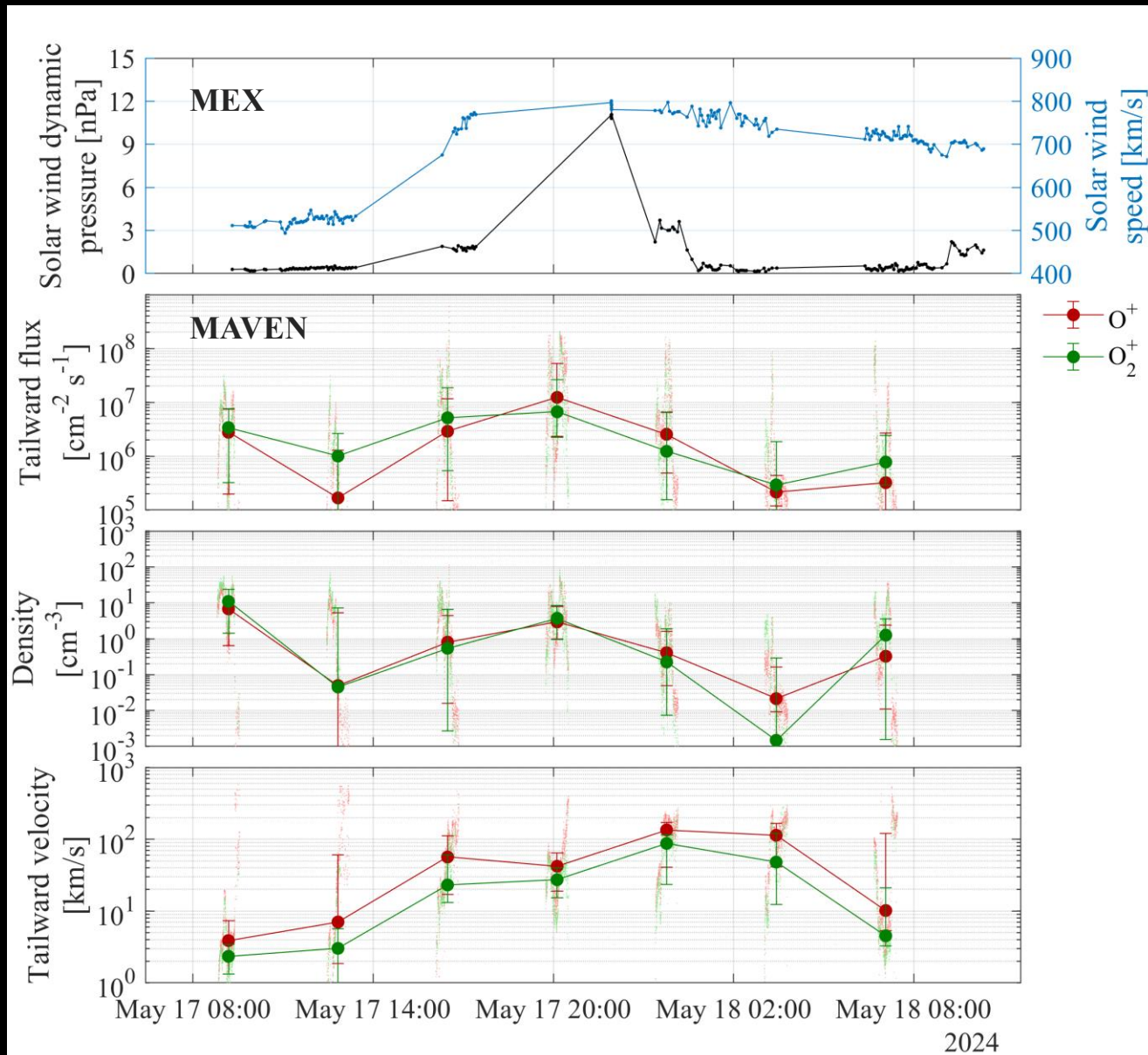
# ICME



# After impact

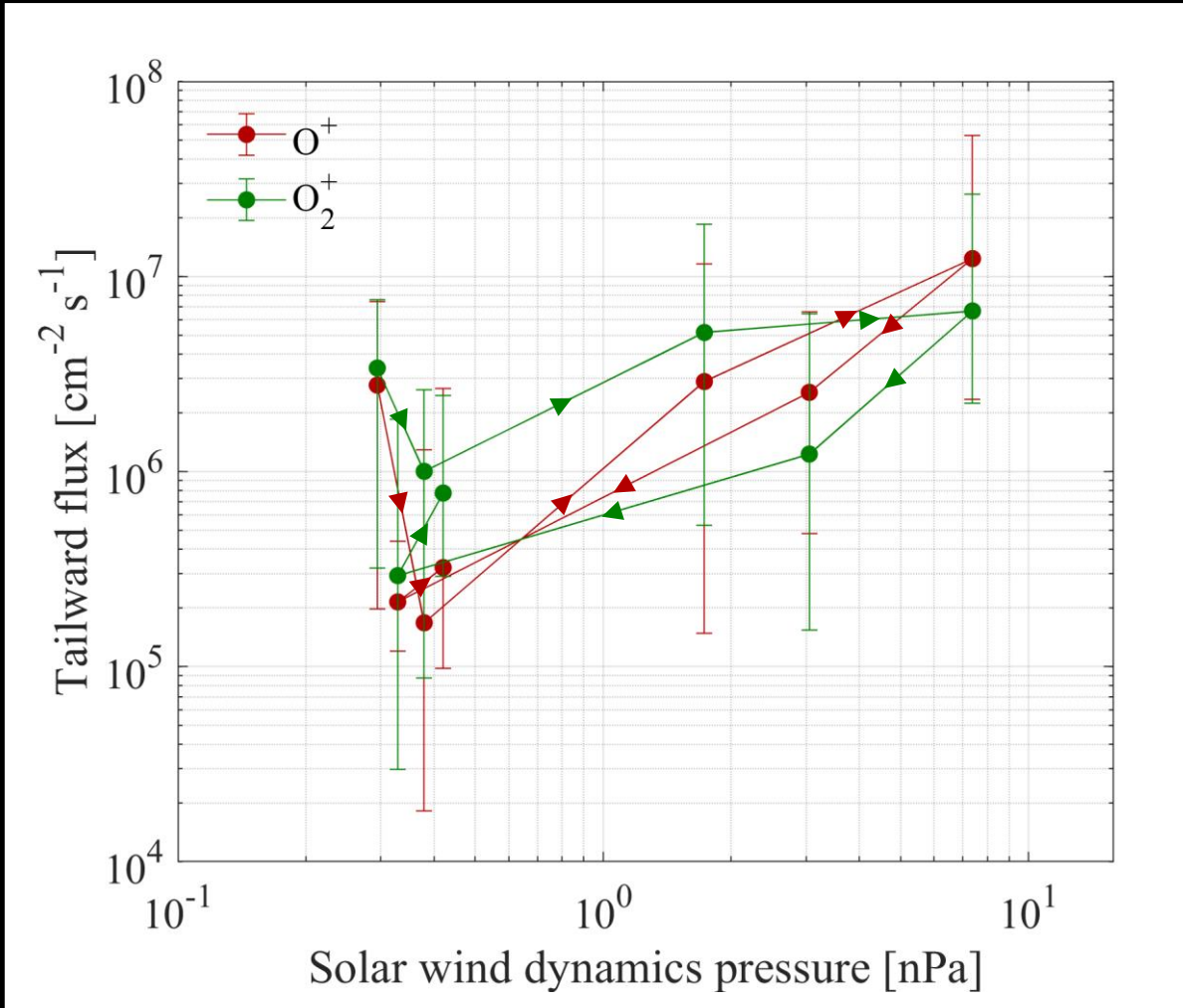


# Timeline



- Tailward O<sup>+</sup>, O<sub>2</sub><sup>+</sup> fluxes increase at ICME onset, reduce to below pre-ICME levels after impact
- O<sup>+</sup>, O<sub>2</sub><sup>+</sup> velocities increase steadily
- The reduction in flux appears due to evacuation of plasma in the tail

# Solar Wind as a Driver of Ion Escape



- The intense ICME solar wind drove a momentary increase in  $O^+$ ,  $O_2^+$  fluxes
- $O_2^+$  flux saturates and is lower for the descending phase of the event
- Ions cannot be resupplied at the rate they are energized to escape

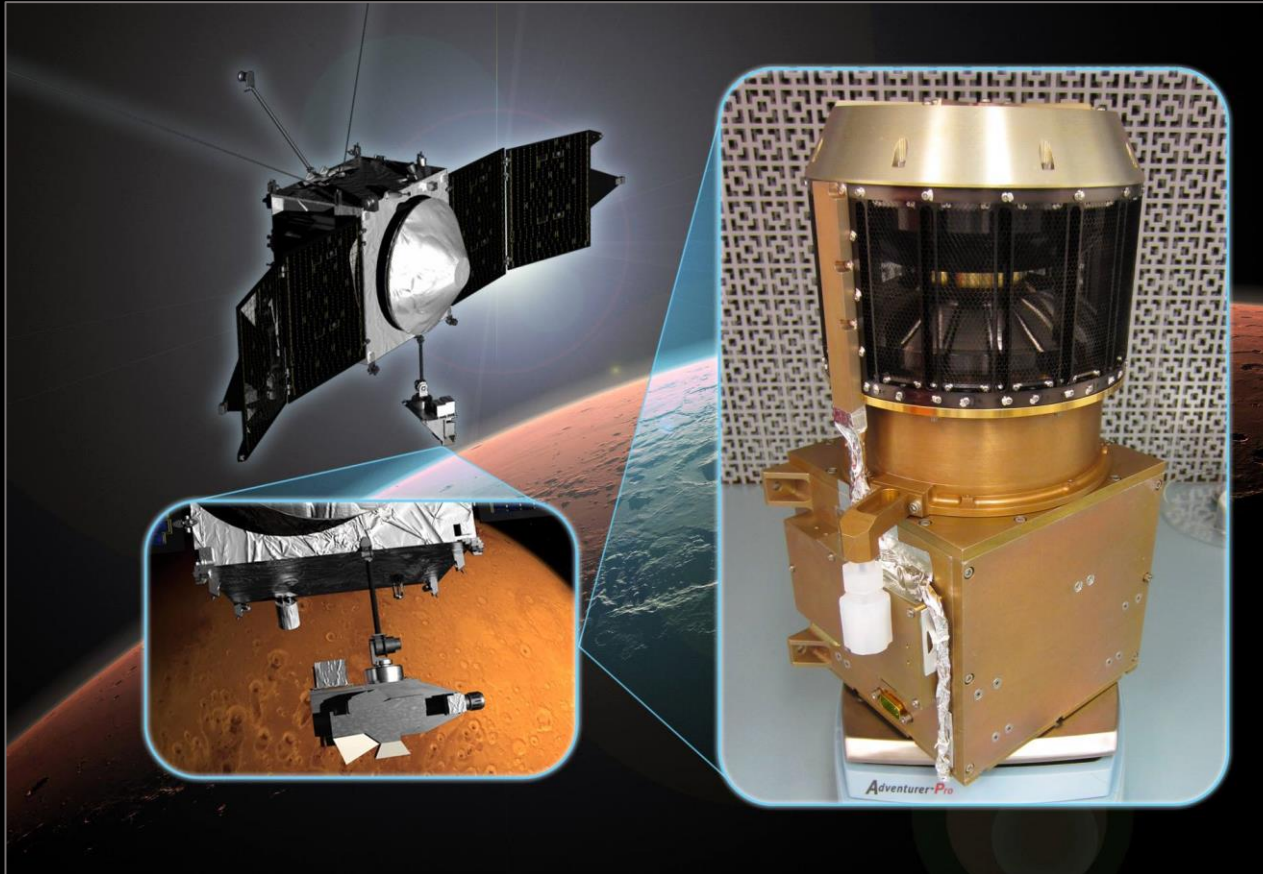
# Summary

- An ICME impacted Mars on May 17, 2024, local effects observed by MAVEN and Mars Express
- $O_2^+$  escaping fluxes show a hysteresis in response to the ICME conditions with fluxes lower for the latter part of the impact, despite similar dynamic pressure
- Tailward velocities of escaping ions increase with time, reduction in flux appears due to evacuation of plasma in the tail and insufficient resupply feeding the process

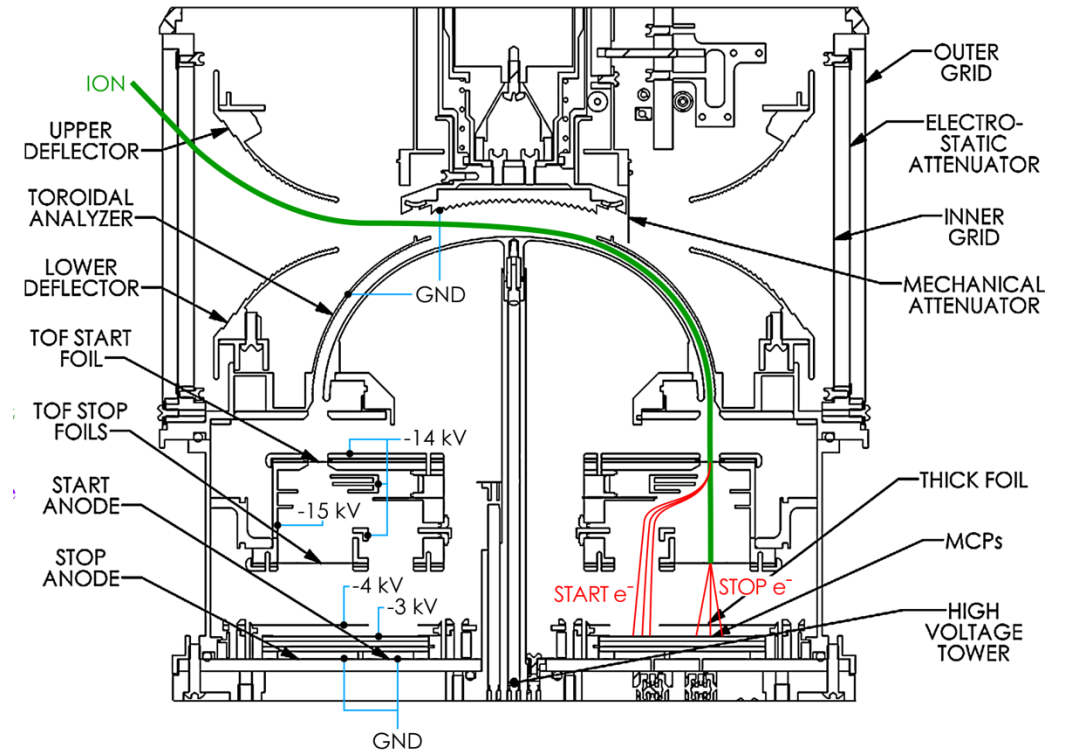
Extra



# MAVEN/STATIC



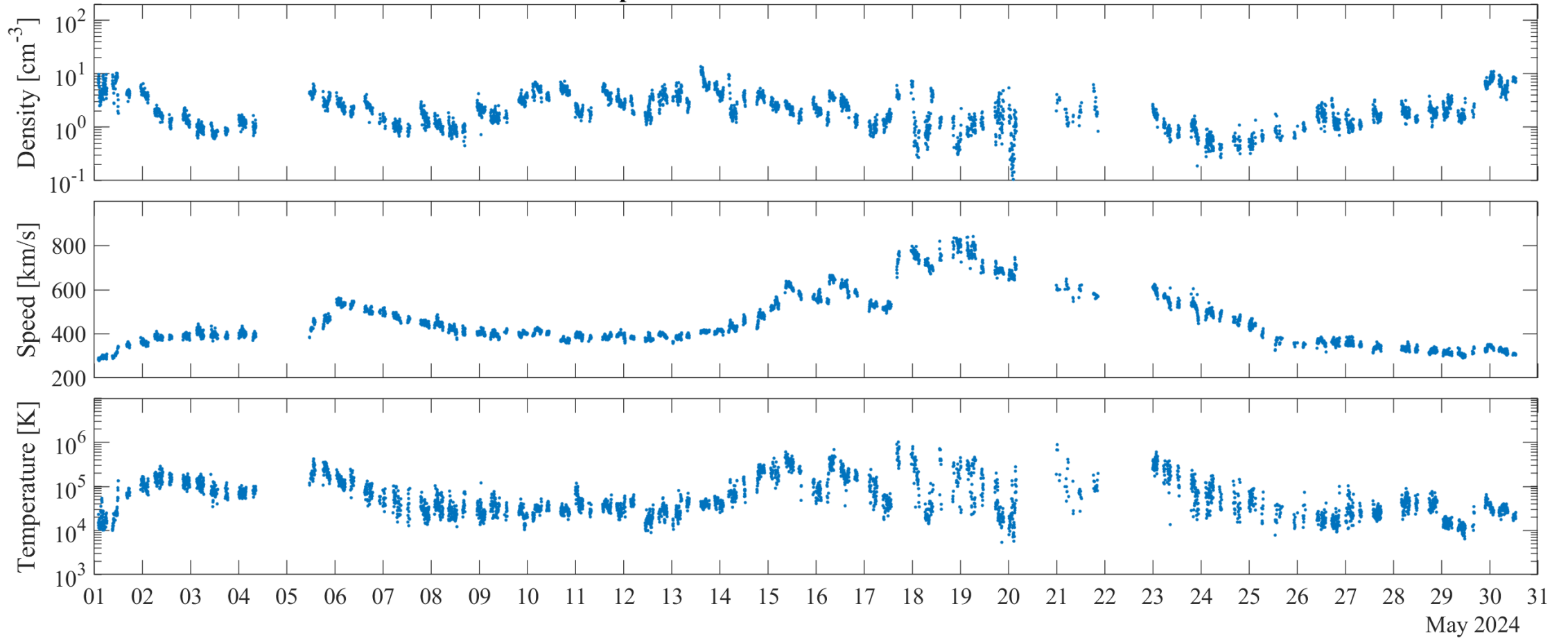
MAVEN STATIC



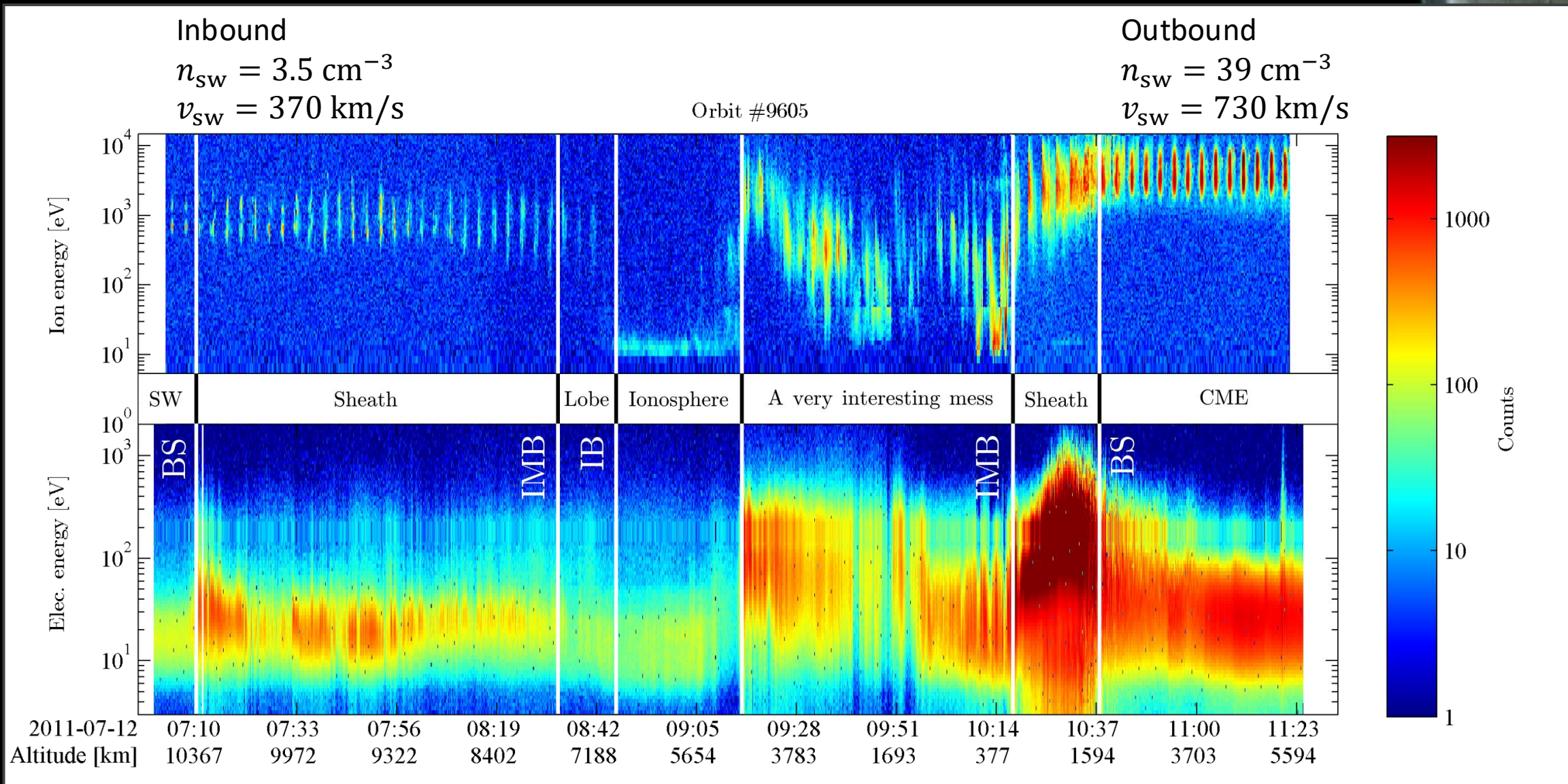
Measurement:  
 Differential flux  $j(\theta, \phi, E, m) [\text{cm}^{-2} \text{s}^{-1} \text{sr}^{-1} \text{eV}^{-1}]$

# May 2024 events

Mars Express ASPERA-3/IMA Solar Wind Moments



# July 2011 Primordial Solar Wind Event



# July 2011 Primordial Solar Wind Event

Orbit #9606

