

Whole Atmosphere Community Climate Model with Thermosphere/Ionosphere Extension (WACCM-X v.2): Model Capabilities and Validation

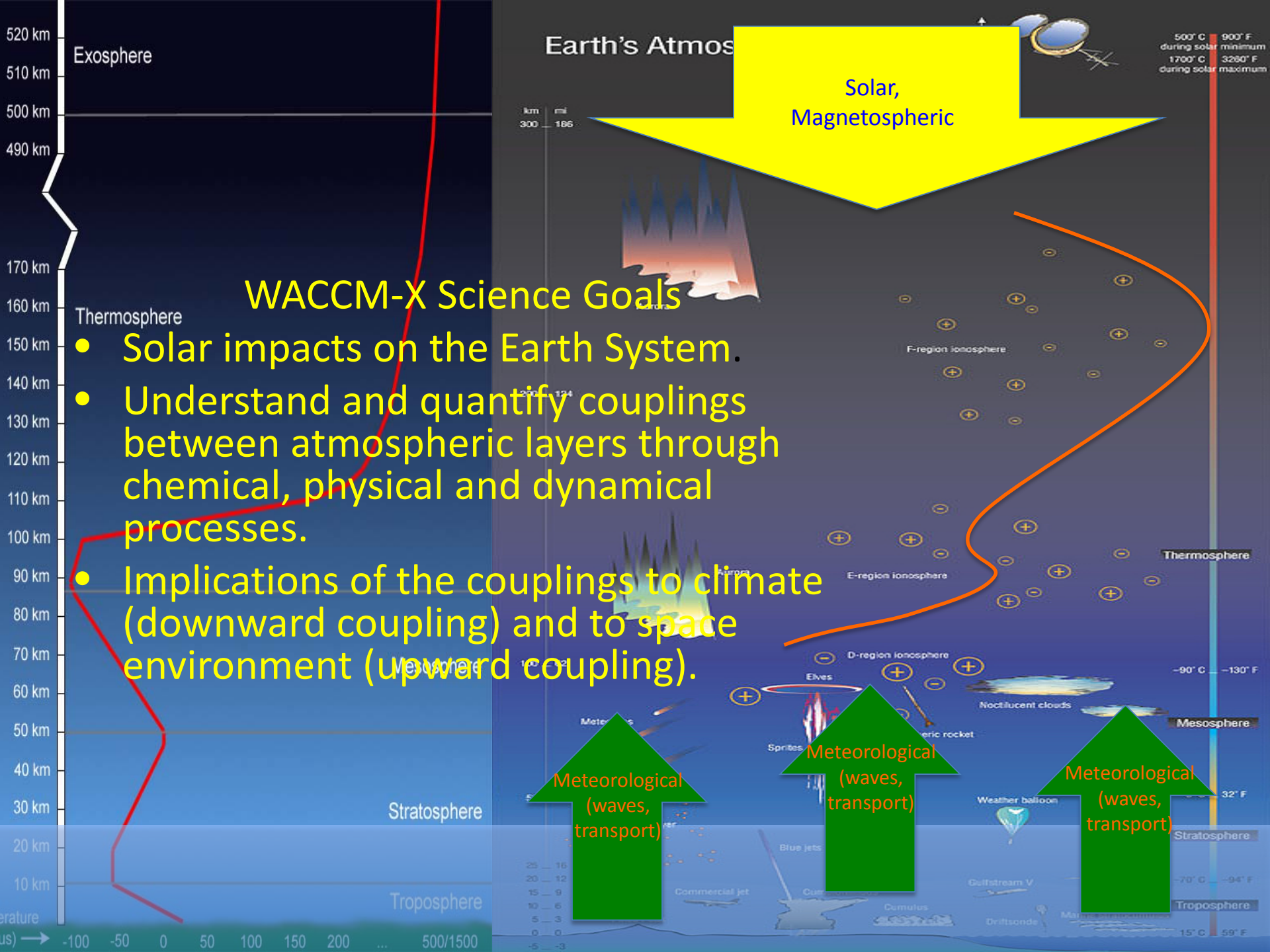
Han-Li Liu and WACCM-X Team

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Major CESM WACCM/WACCM-X Components

Model Framework	Chemistry	Neutral Atmos. Physics	Ionospheric Physics	Resolution
<p>Atmosphere component of NCAR Community Earth System Model (CESM)</p> <p>Extension of the NCAR Community Atmosphere Model (CAM)</p> <p>Finite Volume Dynamical Core (modified to consider species dependent Cp, R, m)</p> <p>Spectral Element Dynamical Core</p>	<p>MOZART+ Ion Chemistry (~60+ species)</p> <p>Fully-interactive with dynamics.</p>	<p>Long wave/short wave/EUV</p> <p>RRTMG</p> <p>IR cooling (LTE/non-LTE)</p> <p>Modal Aerosol/CARMA</p> <p>Convection, precip., and cloud param.</p> <p>Parameterized GW</p> <p>Major/minor species diffusion (+UBC)</p> <p>Molecular viscosity and thermal conductivity (+UBC)</p> <p>Species dependent Cp, R, m.</p>	<p>Parameterized electric field at high, mid, low latitudes. IGRF geomagnetic field.</p> <p>Auroral processes, ion drag and Joule heating</p> <p>Ion/electron energy equations</p> <p>Ambipolar diffusion</p> <p>Ion/electron transport</p> <p>Ionospheric dynamo</p> <p>Coupling with plasmasphere/magnetosphere</p>	<p>Horizontal: 1.9° x 2.5° (lat x lon configurable as needed)</p> <p>Vertical: 66 levels (0-140km) 81/126 levels 0--600km</p> <p>Mesoscale-resolving version: 0.25 deg/0.1 scale height.</p>

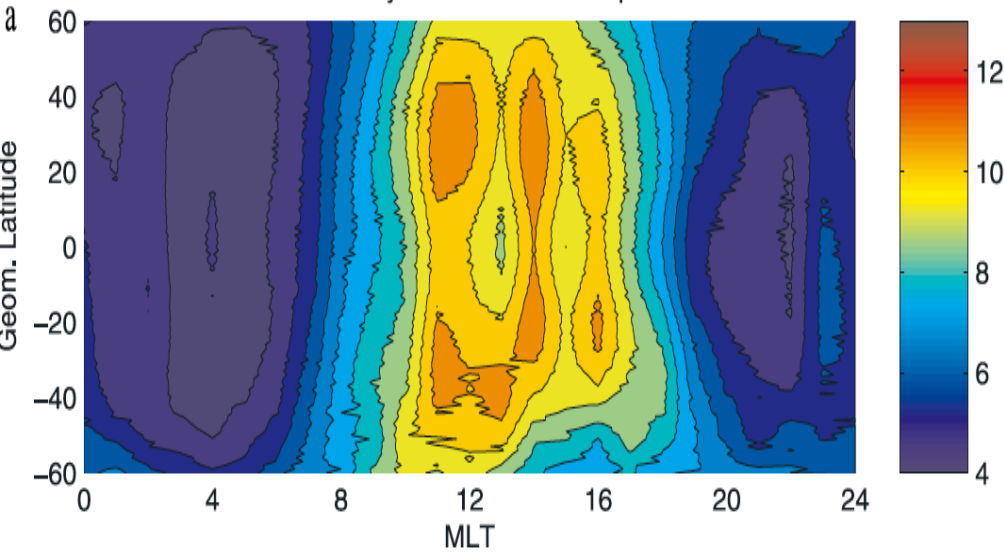
Key WACCM-X Capabilities

- Physics-based whole atmosphere general circulation model (0-700km)
- Solves dynamics, radiative transfer, photolysis and energetics
- Fully interactive chemistry, including ion chemistry.
- Ionospheric electrodynamics using fully interactive dynamo
- Ion transport in the *F*-region
- Magnetospheric inputs using empirical or specifications, including AMIE
- Coupling with a plasmasphere model (partnership with NRL)
- Whole atmosphere data assimilation for specification and forecast.

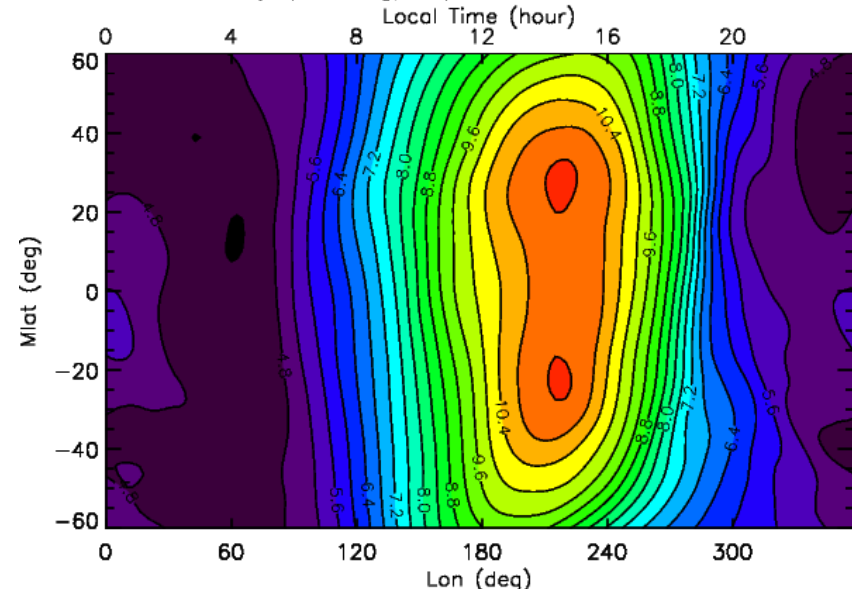


Mass and Electron Density at 400km

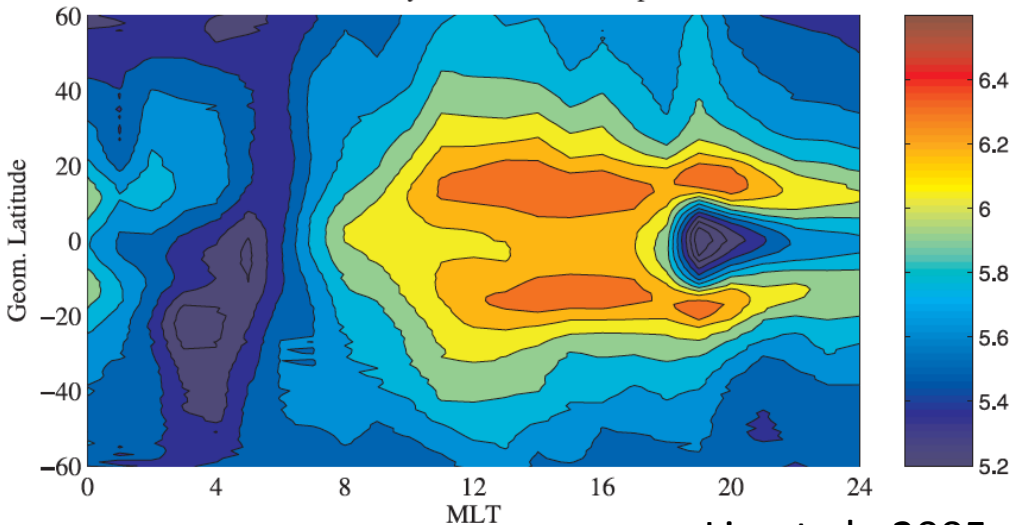
Mass Density from CHAMP for Kp=0...2



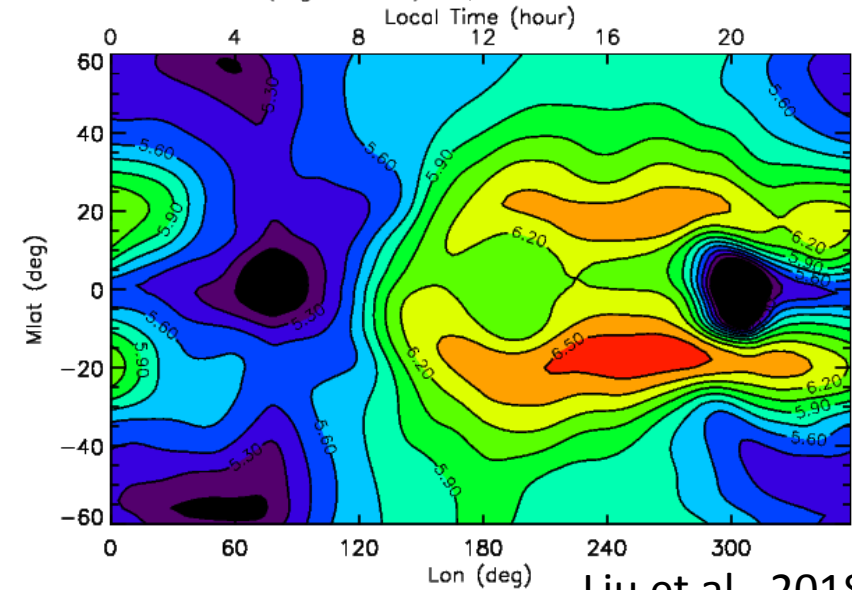
density (10⁻¹² kg/m³) Sep 400km, F10.7=200



Electron Density from CHAMP for Kp=0..2



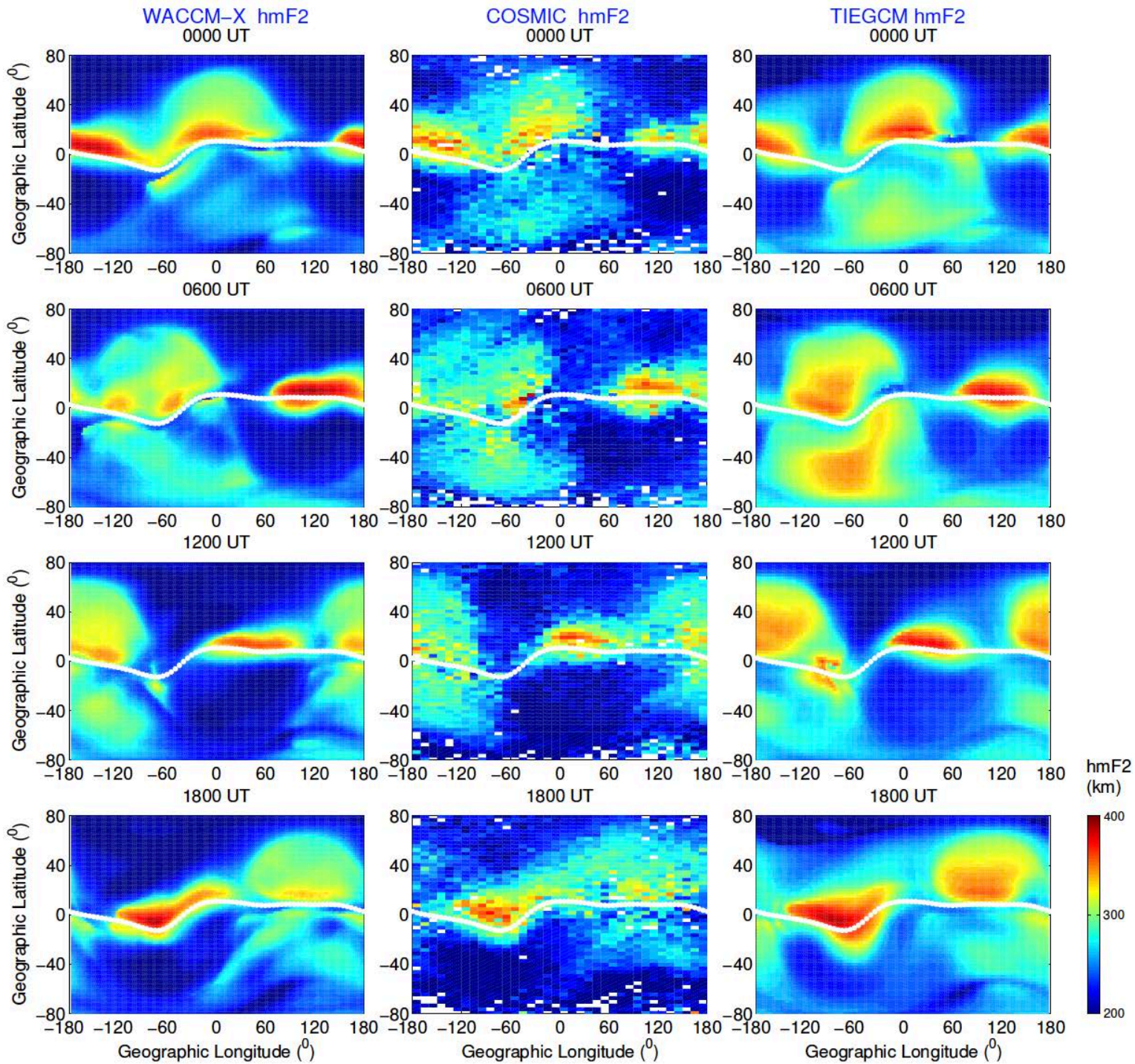
Ne (log10 cm⁻³) Sep 400km, F10.7=200



Liu et al., 2005

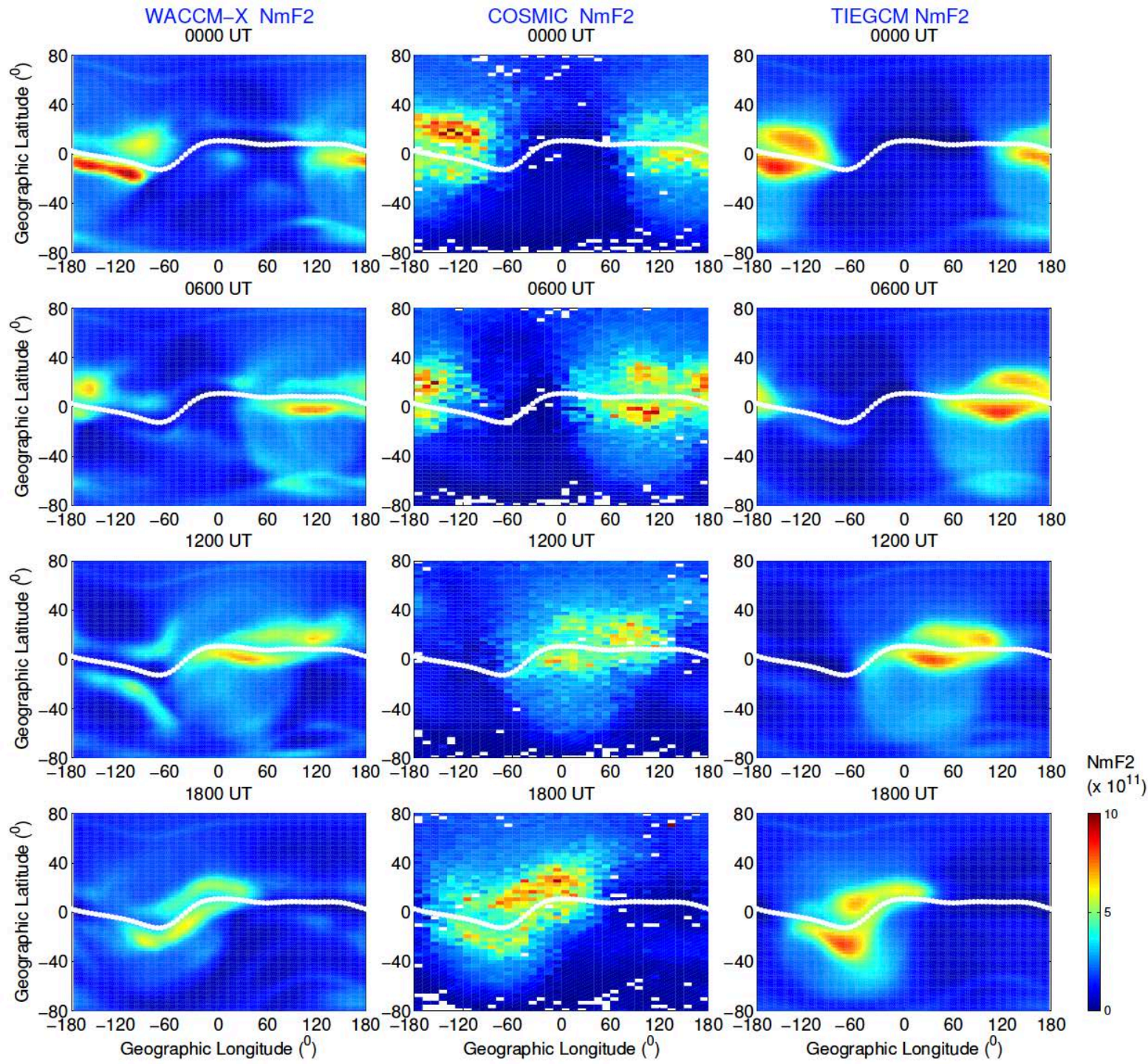
Liu et al., 2018

Comparison with COSMIC 2008 June

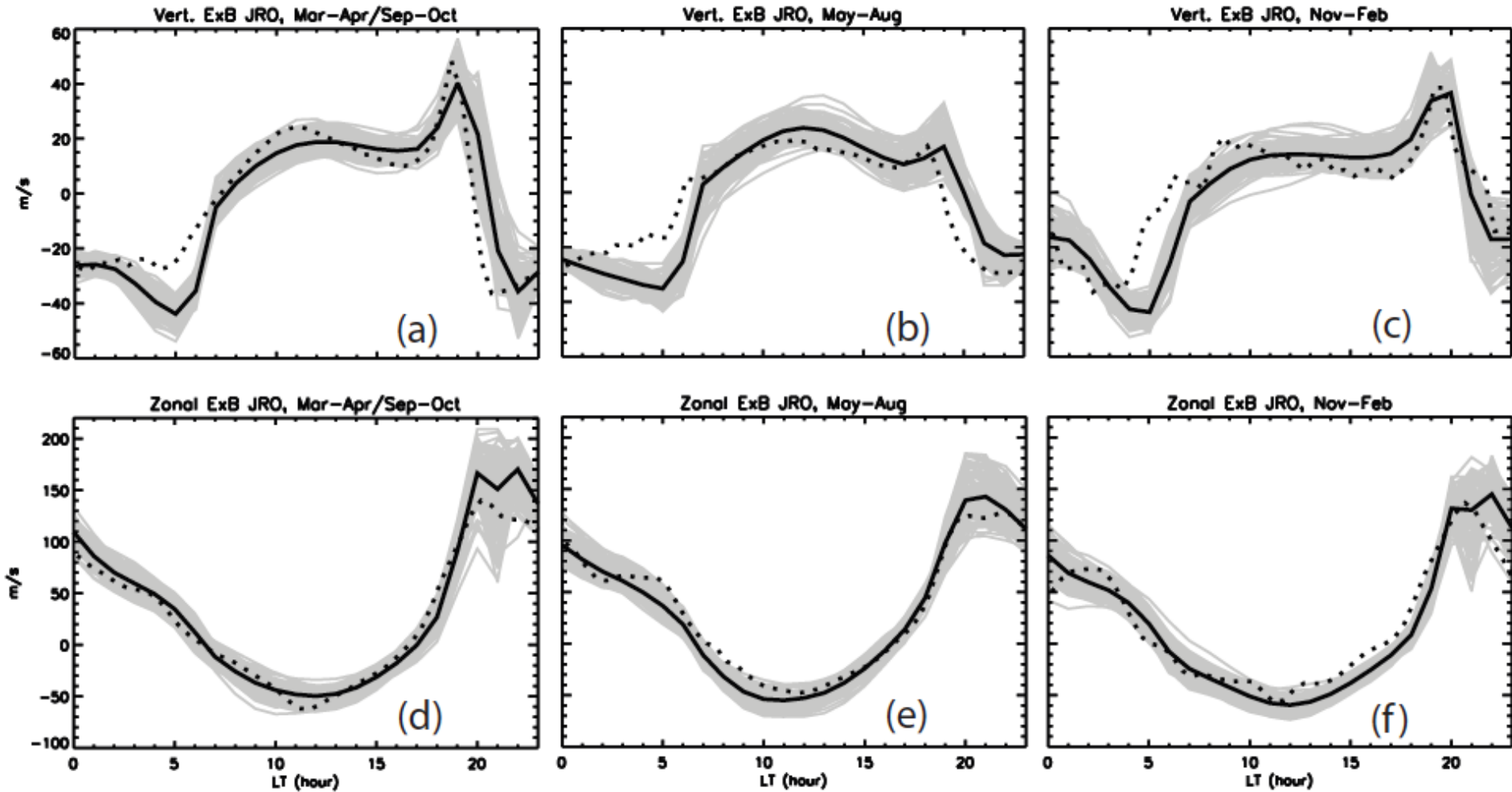


J. Liu et al. 2018

Comparison with COSMIC 2008 June



ExB Drifts: WACCM-X vs Climatology

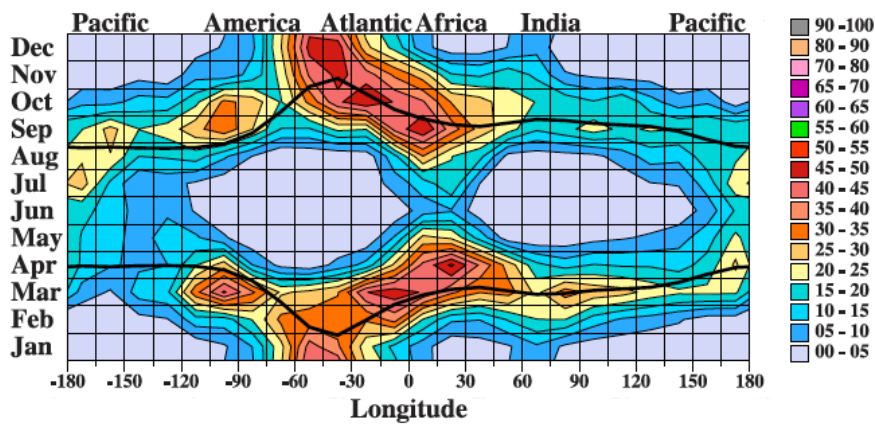


Liu et al., 2018

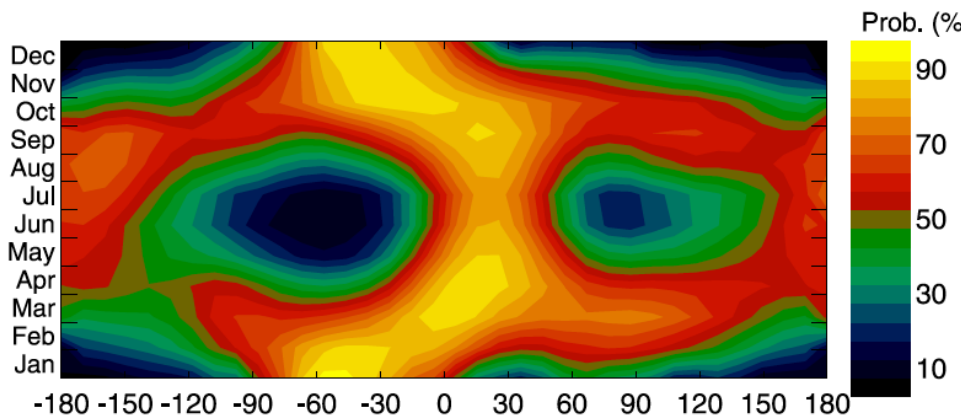
Dotted line: JRO climatology (Fejer et al., 1991)

Variability of PRE and Equatorial Plasma Bubbles (EPB)

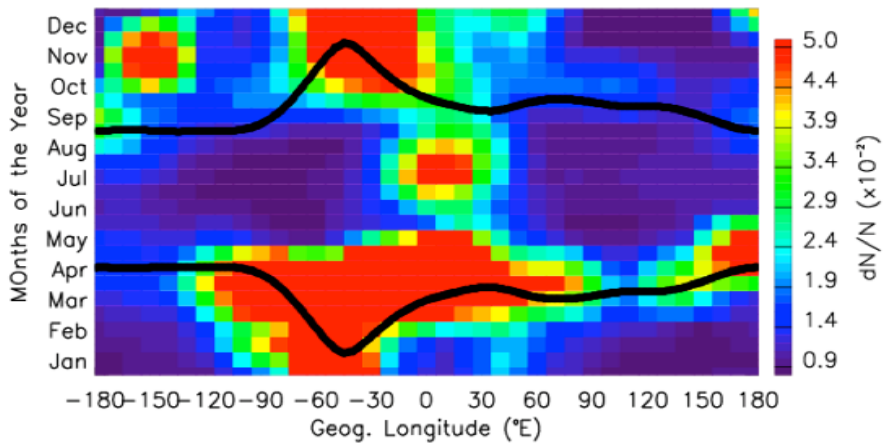
DMSP EPB Rates 1999 - 2002



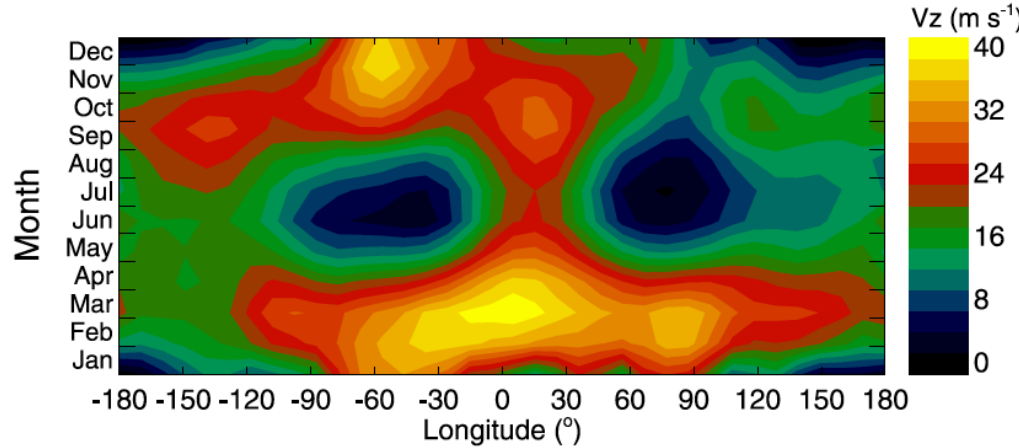
Gentile et al., 2006



19:00 - 24:00 Local Time

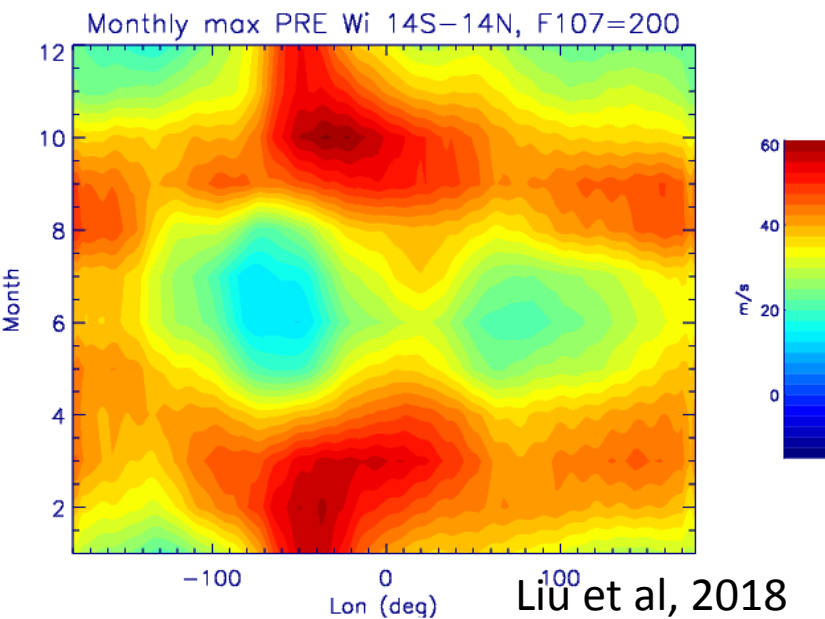
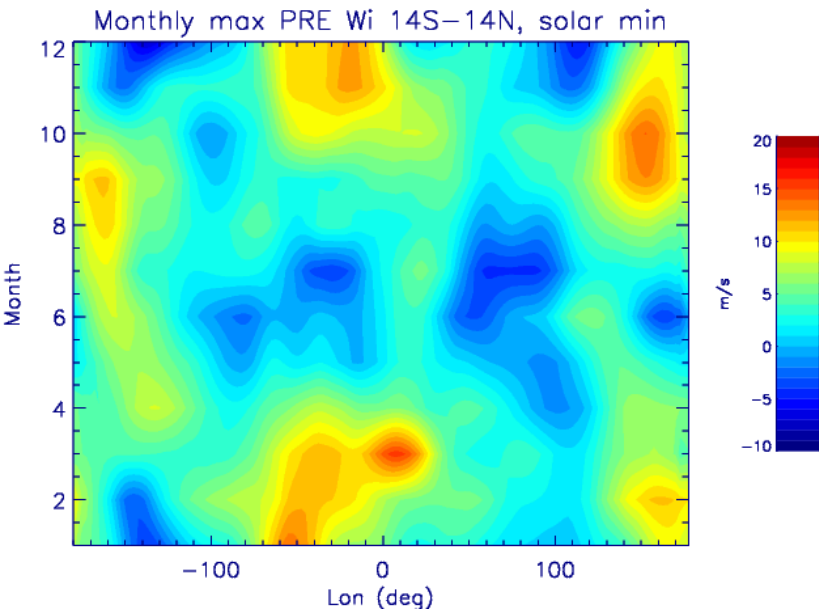


Yizengaw et al., 2014

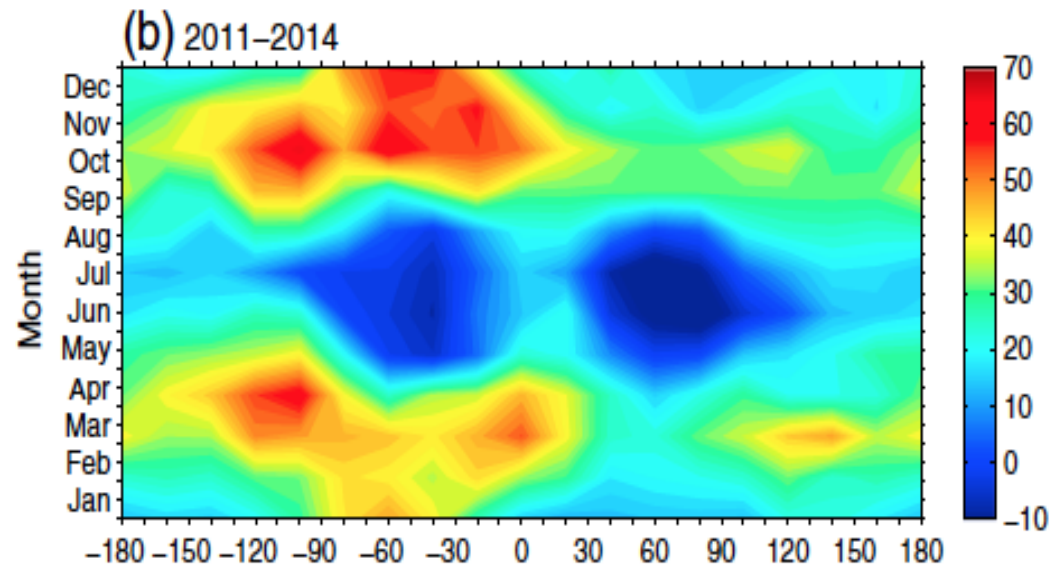
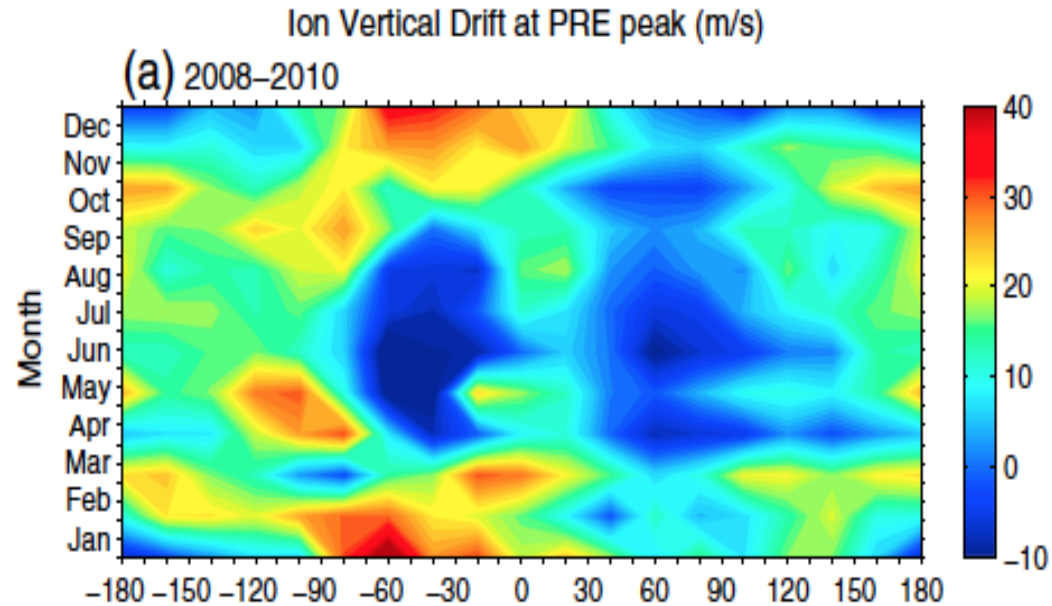


Kil et al., 2009

Monthly Mean PRE Peak: WACCM-X and Obs.

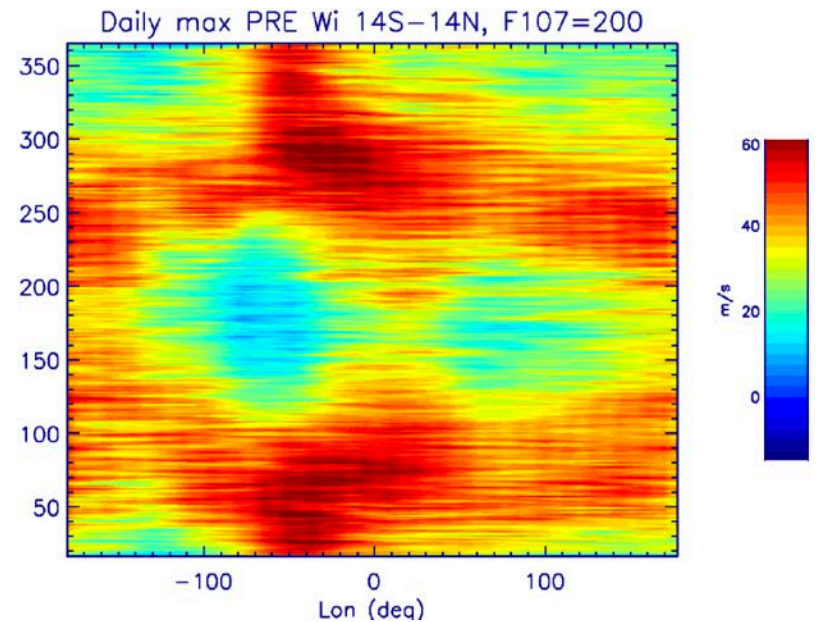
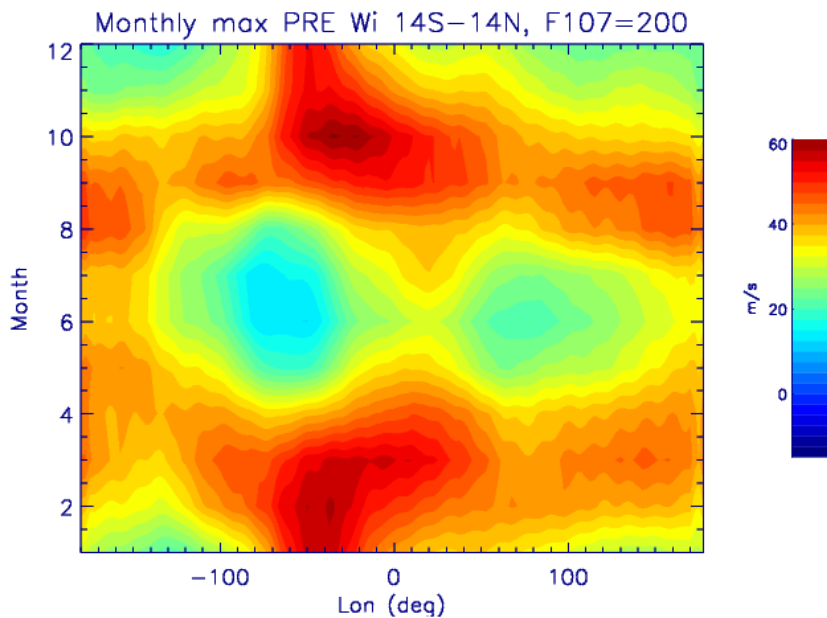
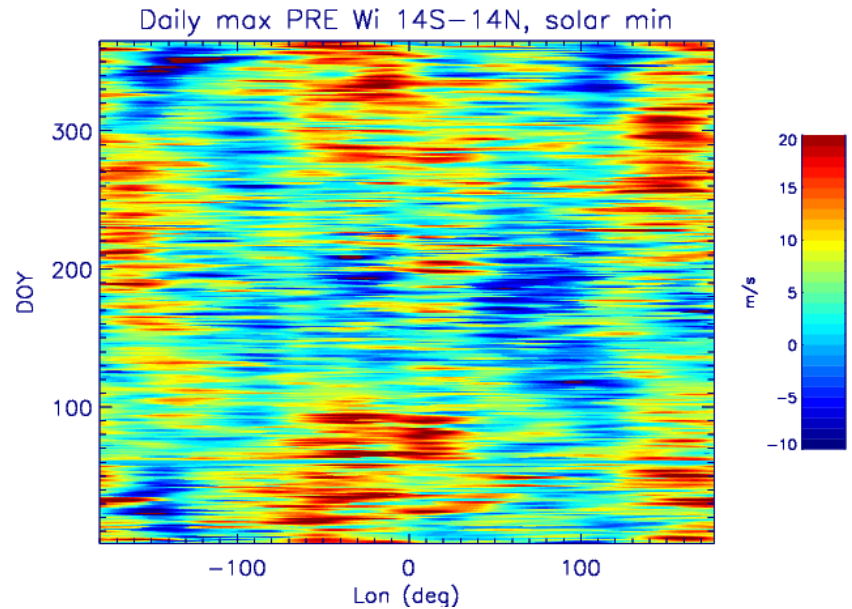
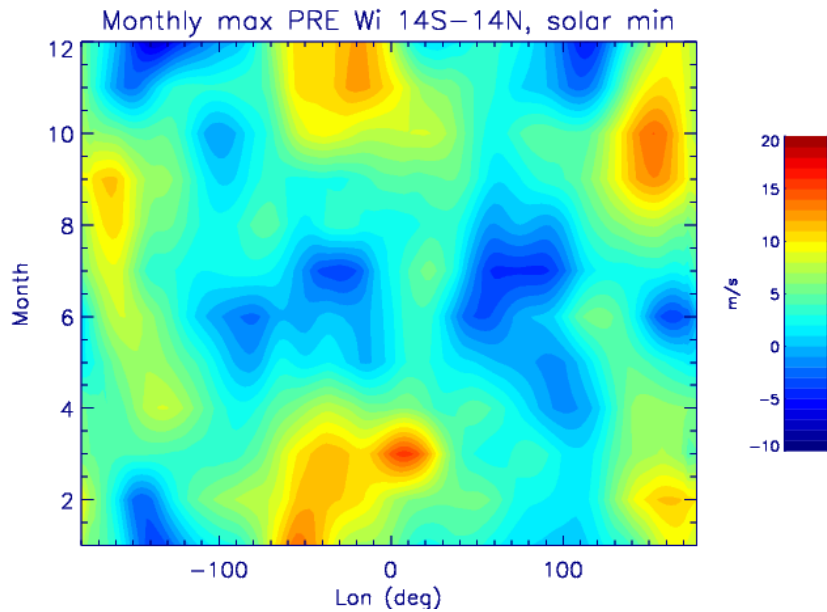


Liu et al, 2018

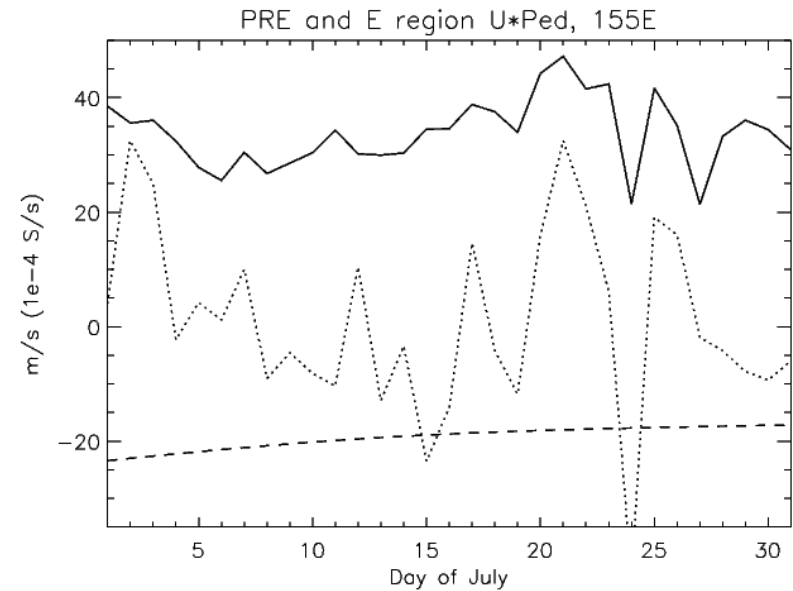
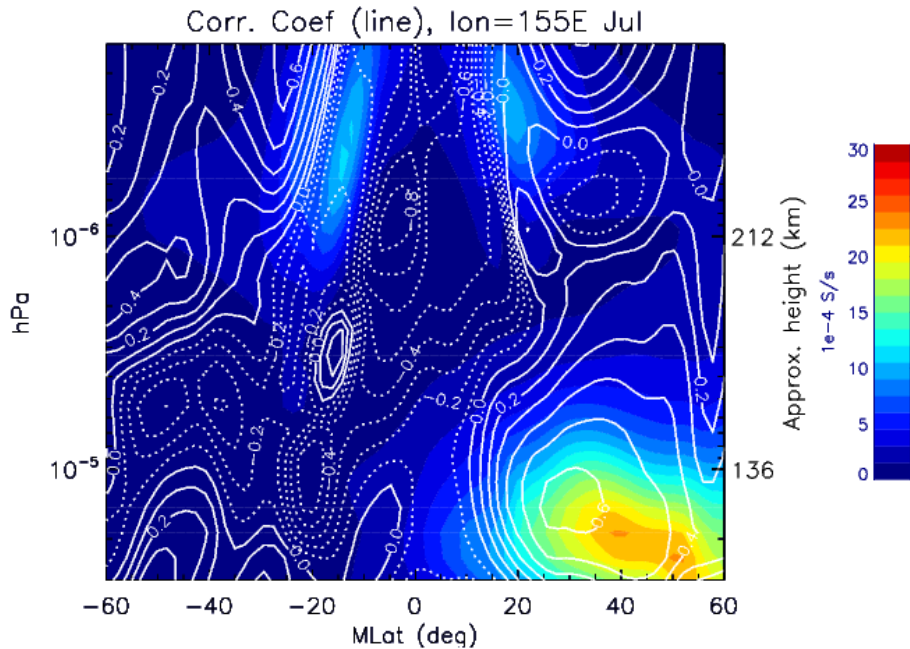


Huang and Hairston, 2015

Monthly vs Daily Variability



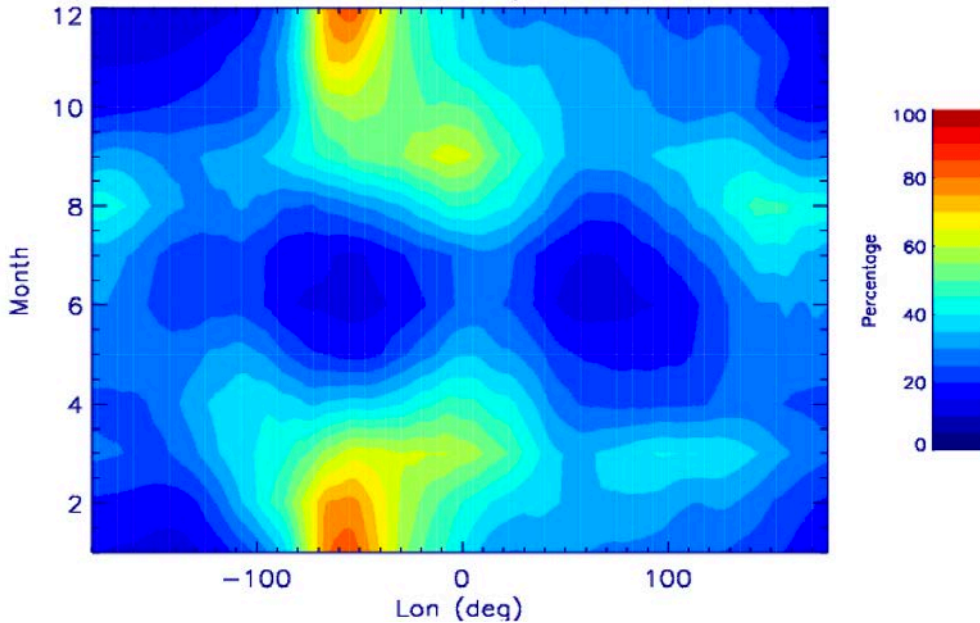
Variability of PRE and E-region Dynamo



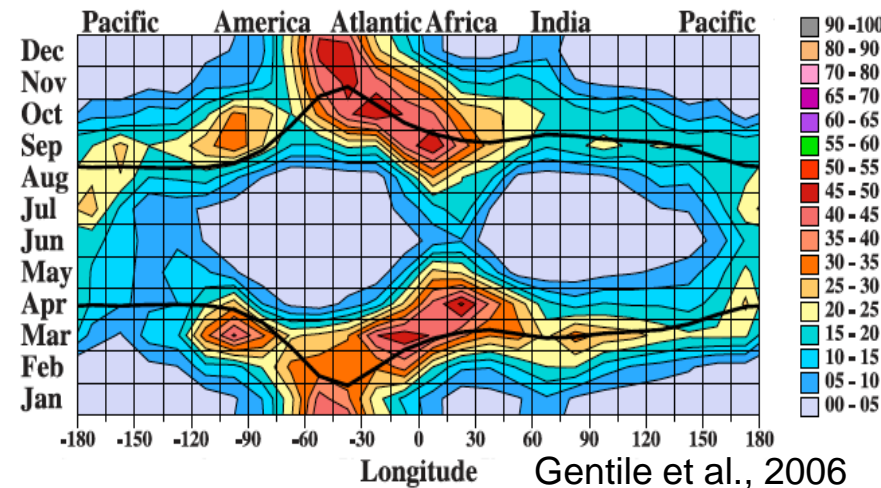
- Summer side E-region neutral wind variability at sunset time strongly affects PRE variability.
- E-region is strongly affected by lower atmospheric waves.

Occurrence Frequency of Equatorial Plasma Bubbles

Deduced EPB Rates, 2000–2002

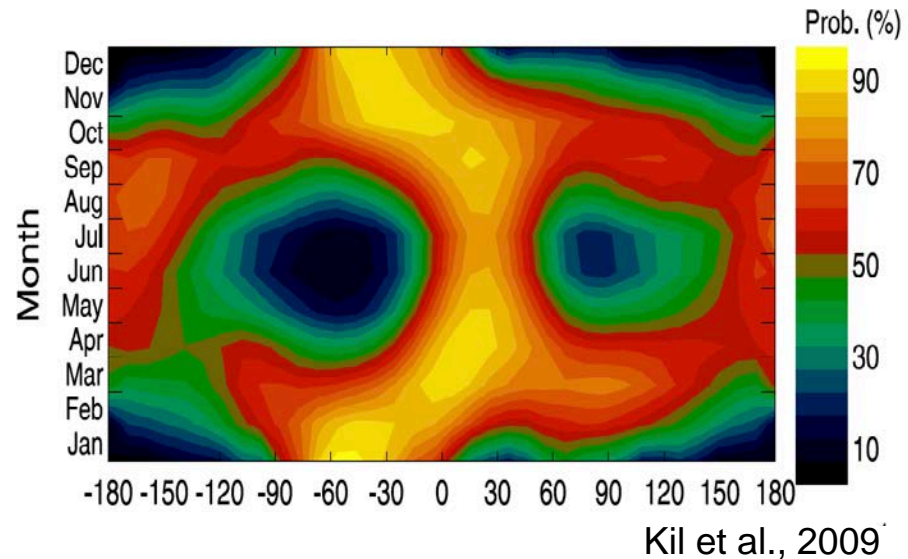


DMSP EPB Rates 1999 - 2002



Gentile et al., 2006

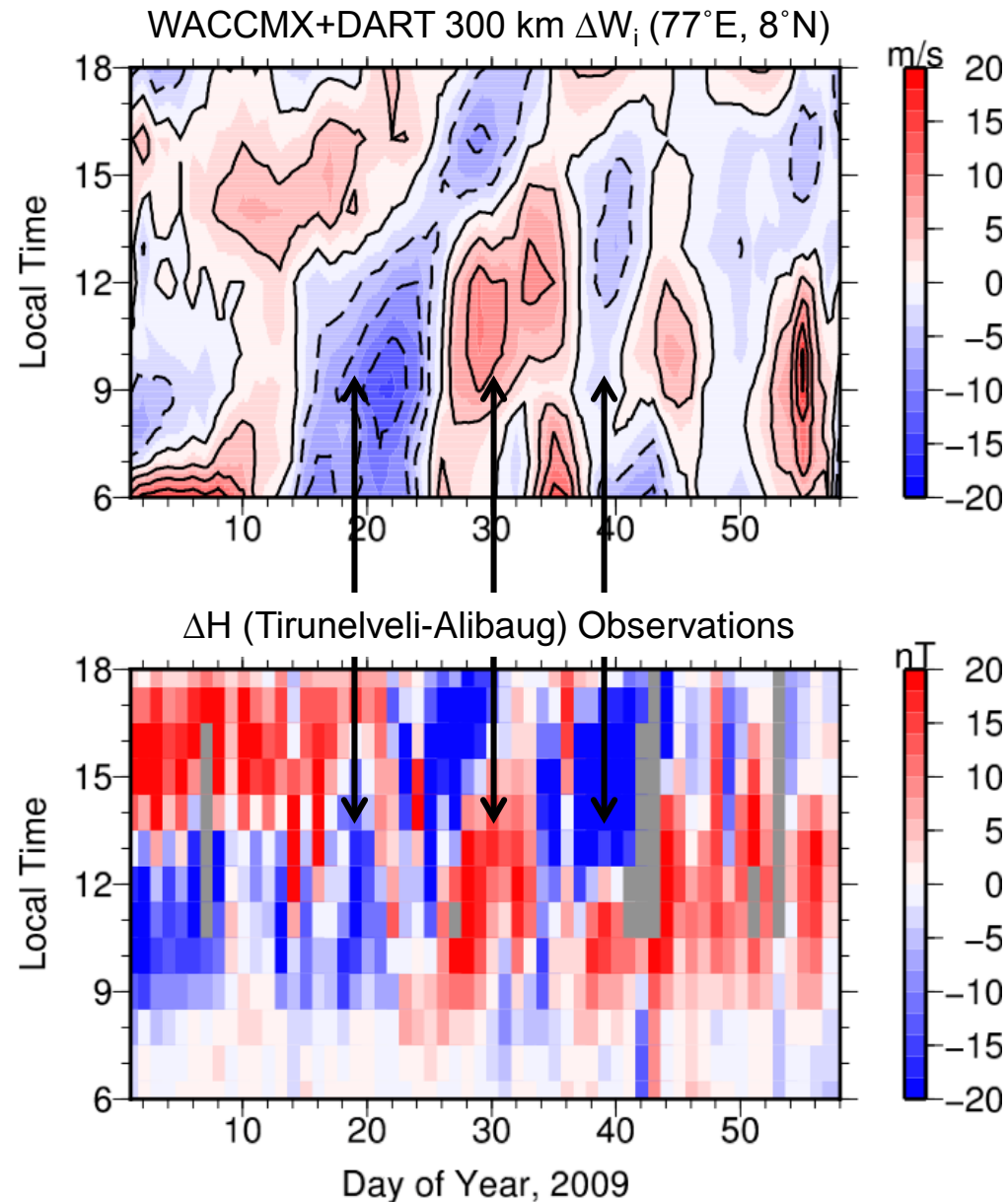
- The agreement between deduced EPB rates and the observed rates suggest
 - Large-scale dynamics and electrodynamics play a key role in preconditioning EPB
 - Feasibility for probabilistic forecast of EPB—an outlook/warning (analogous to tornado forecast).
- Simulating EPB evolution requires high-resolution capability.



Kil et al., 2009

WACCM-X Data Assimilation: WACCM-X+DART

Ionosphere results based on assimilating data only below 100km (meteorological observations, Aura/MLS, and TIMED/SABER): **Dominant ionospheric features can be forecast 10-20 days in advance.**



Summary

- Key WACCM-X capabilities have been developed, and validated against thermospheric and ionospheric observations for both geomagnetically disturbed periods and quiet periods.
- Simulated PRE, an important quantity for the formation of EPB, shows longitudinal and seasonal variation similar to observations.
- Simulated PRE varies significantly from day-to-day. Deduced EPB rate is similar to observations.
 - Large-scale dynamics/electrodynamics important for preconditioning EPB.
 - Feasibility of probabilistic forecast of equatorial space weather events.