

Probing realistic CME model with the Icarus heliosphere model and PUNCH observations

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The Sun-Earth system

**Solar
Corona**

**Inner
Heliosphere**

Dynamic solar wind interacts
with the CMEs

WSA semi-
empirical coronal
model
(Arge et al 2004)

Icarus
(Baratashvili et al. 2025)



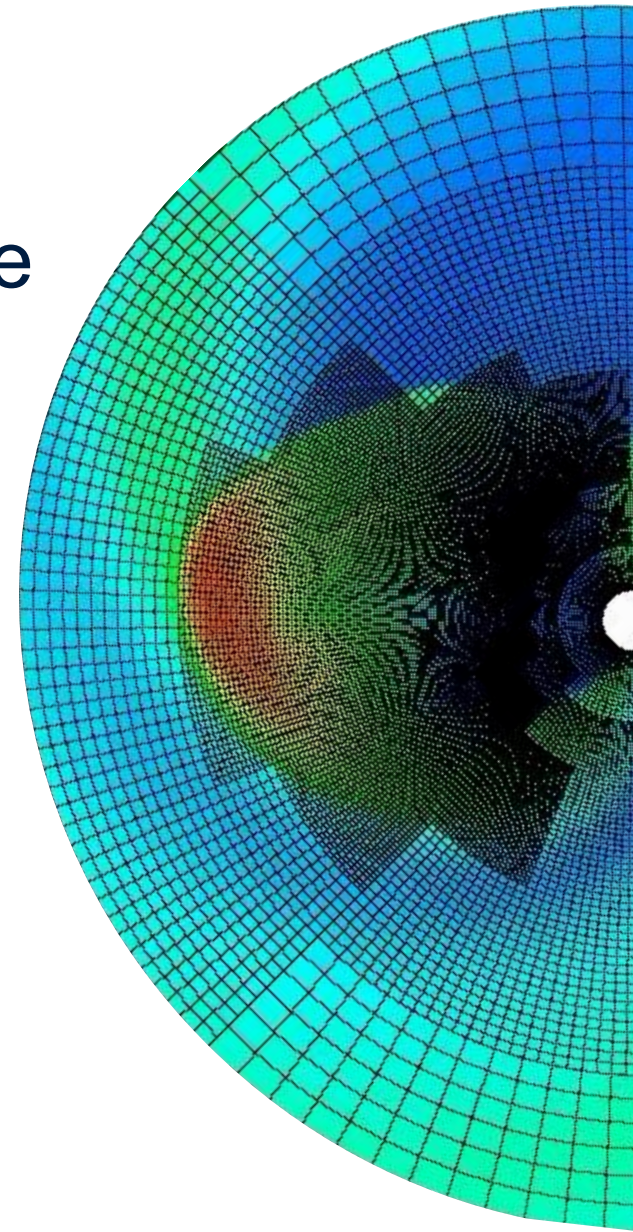
0.1 AU

1 AU

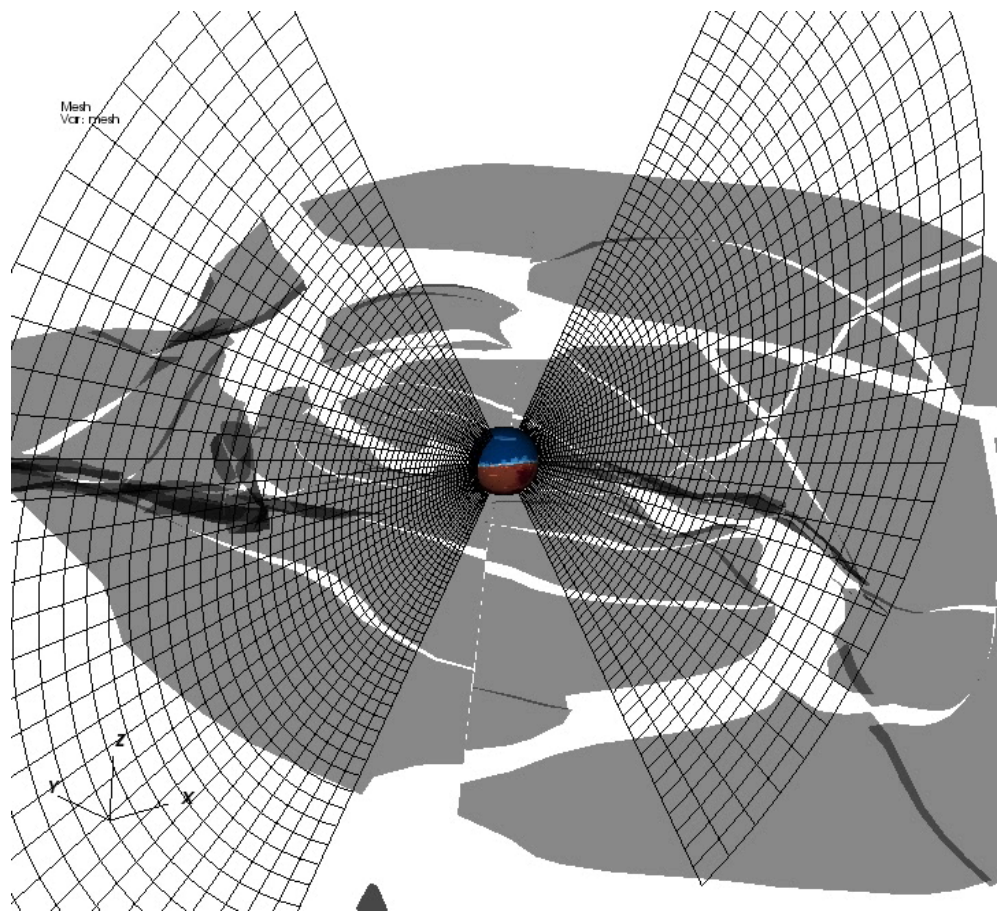


Heliospheric modelling with Icarus

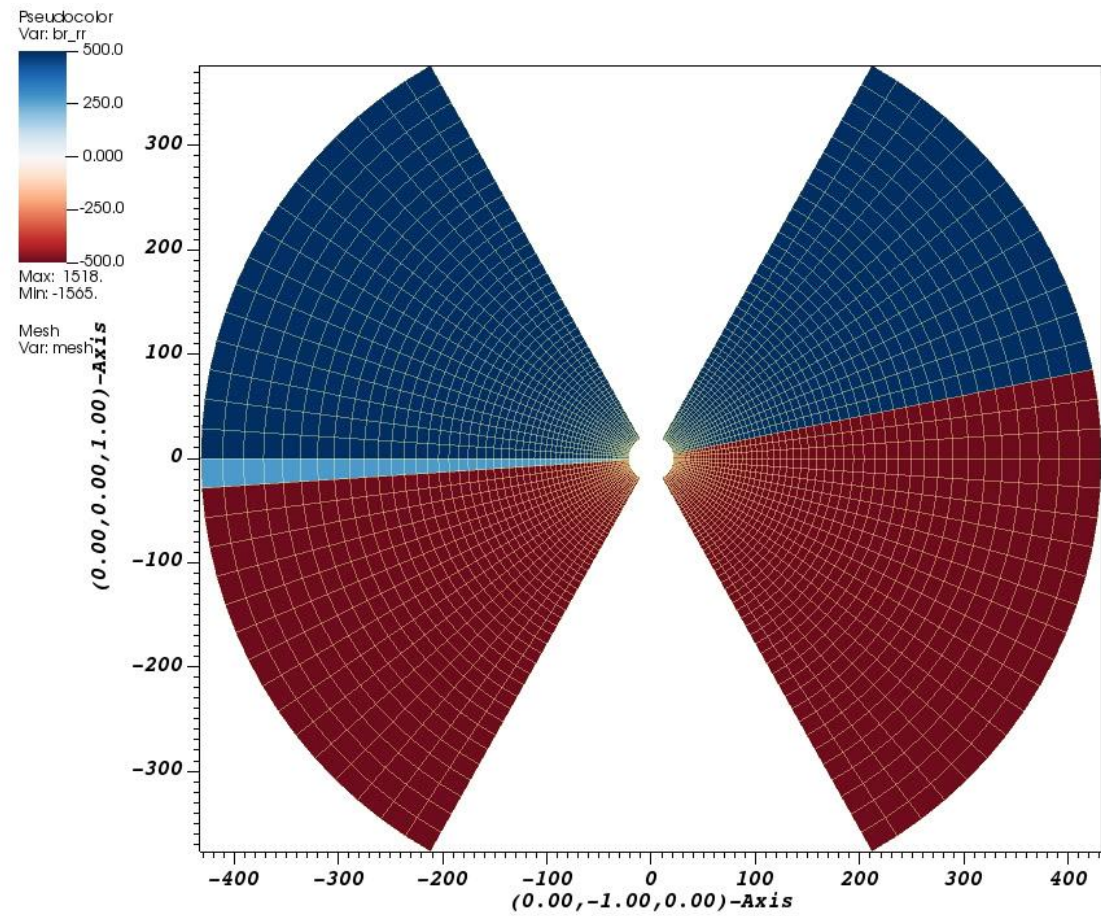
- **Icarus** (Baratashvili et al. 2025) is implemented within **MPI-AMRVAC** (Keppens et al. 2023) framework with MHD module
- **The domain:**
 - Radial - $(21.5, 432) R_{\odot}$
 - Latitudinal - $(-60^{\circ}, 60^{\circ})$ (poles can also be included)
 - Longitudinal - $(0^{\circ}, 360^{\circ})$
- **Co-rotating frame with the Sun**
- **Optimized numerical methods** (Baratashvili et al. 2023)
- **Non-magnetized and magnetized CME models**
- **Time dependent boundary driving**
- **Supports advanced techniques**
 - Radial grid stretching
 - Adaptive mesh refinement (AMR)



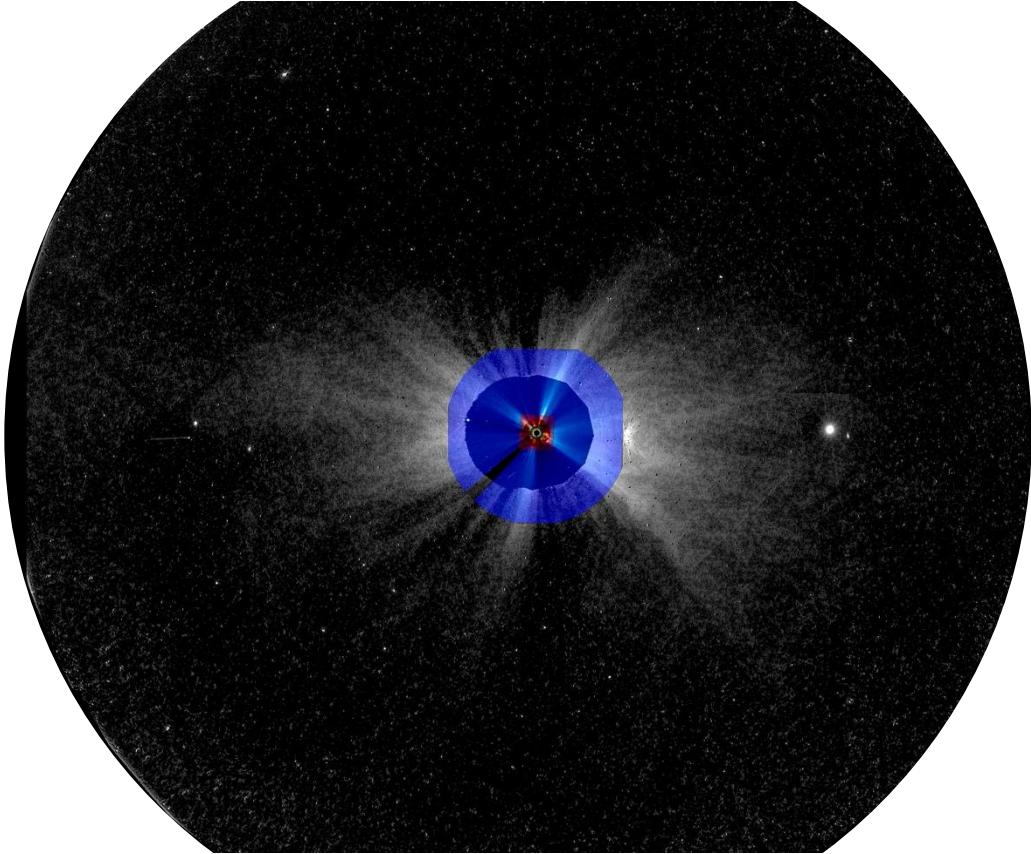
Advanced Techniques: Grid Stretching and Adaptive Mesh Refinement



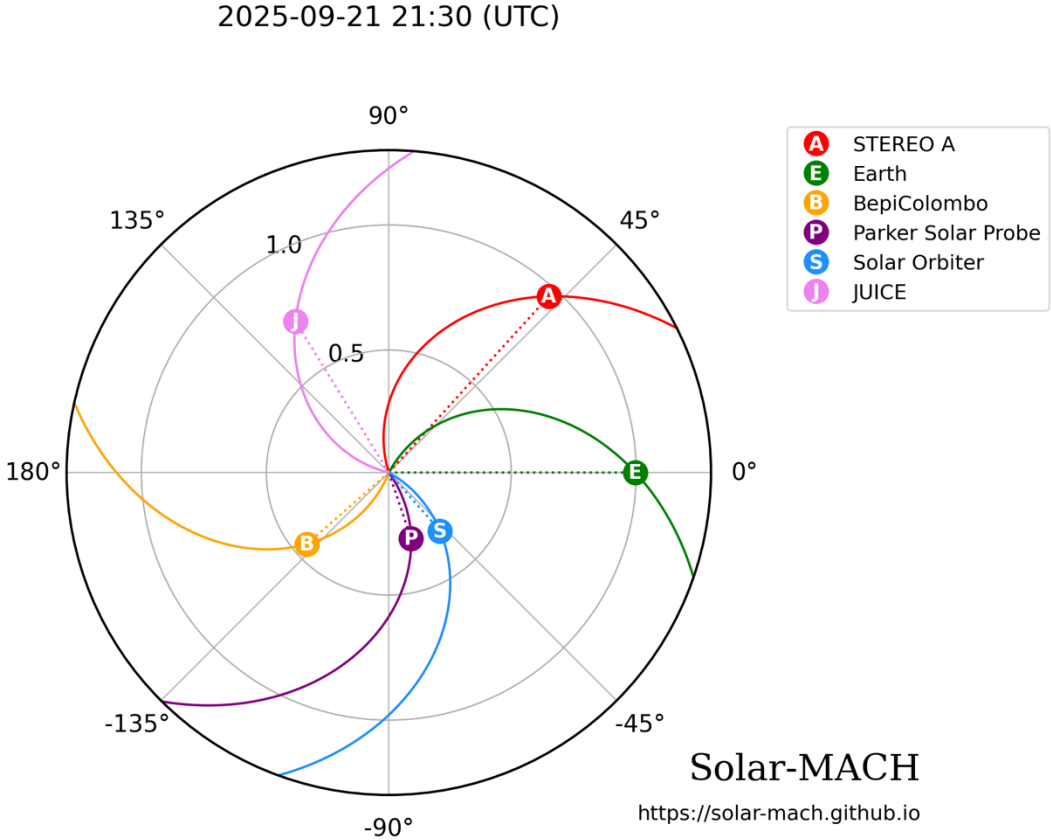
AMR enhanced on the left



September 21, 2025

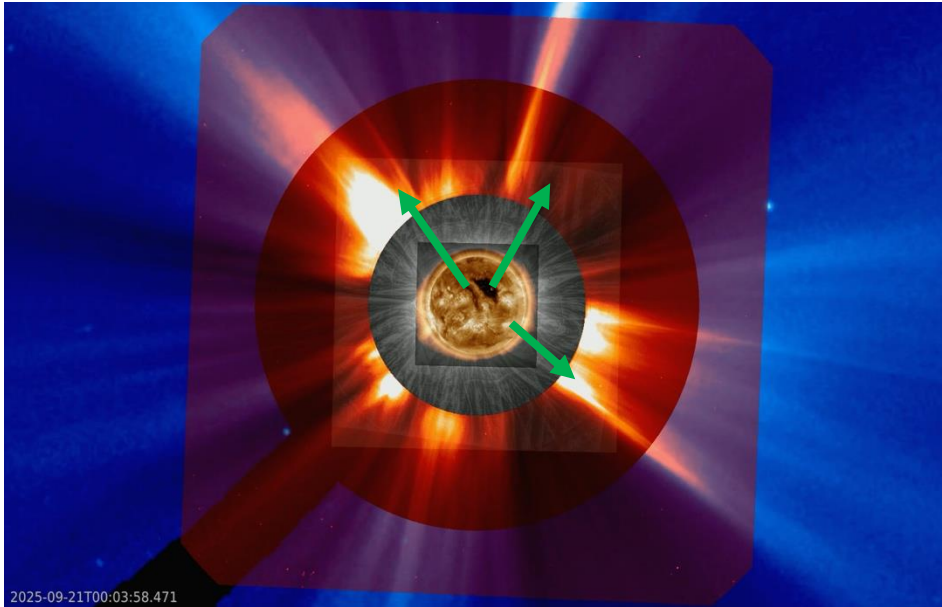


SOHO/LASCO C2+C3 + PUNCH



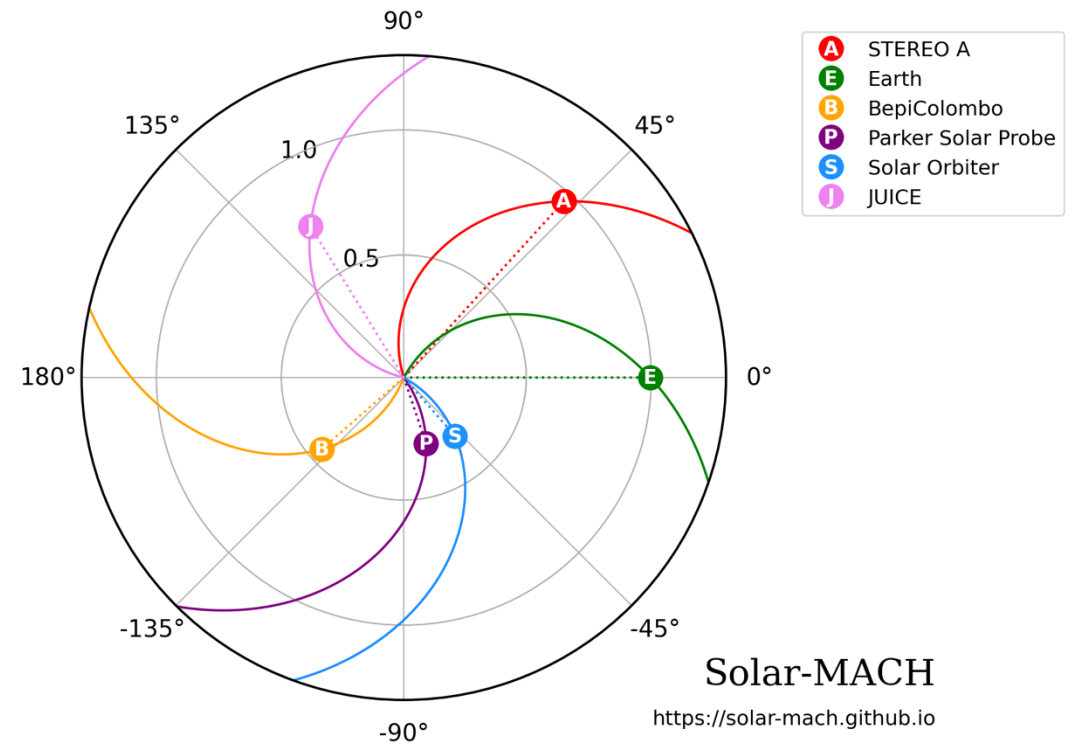
L1 Field of view

Three CMEs erupted almost simultaneously on September 21, 2025



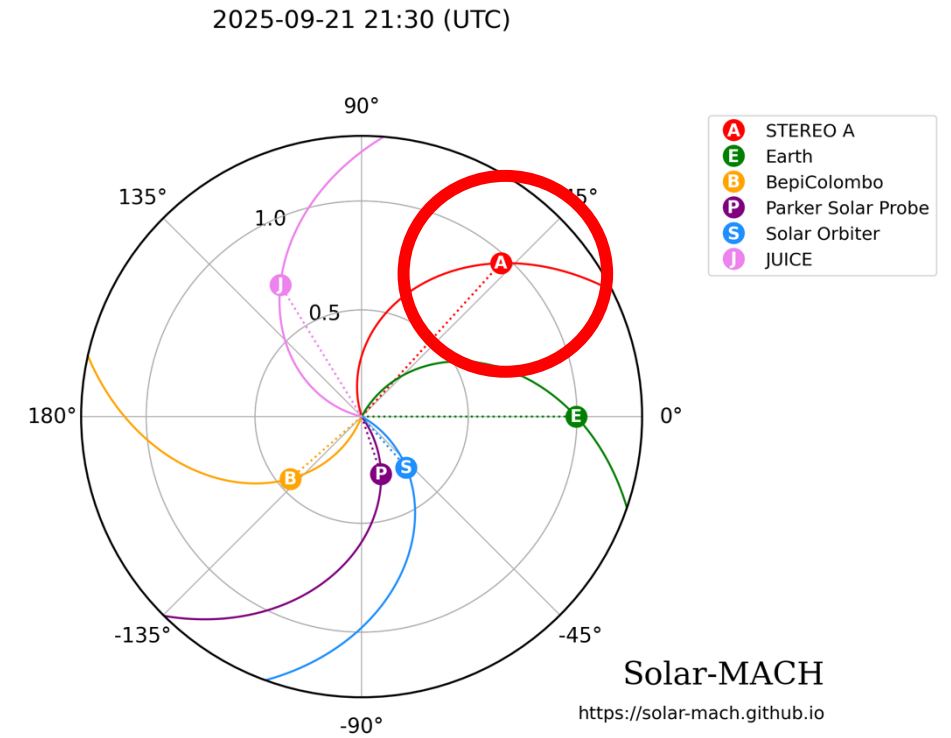
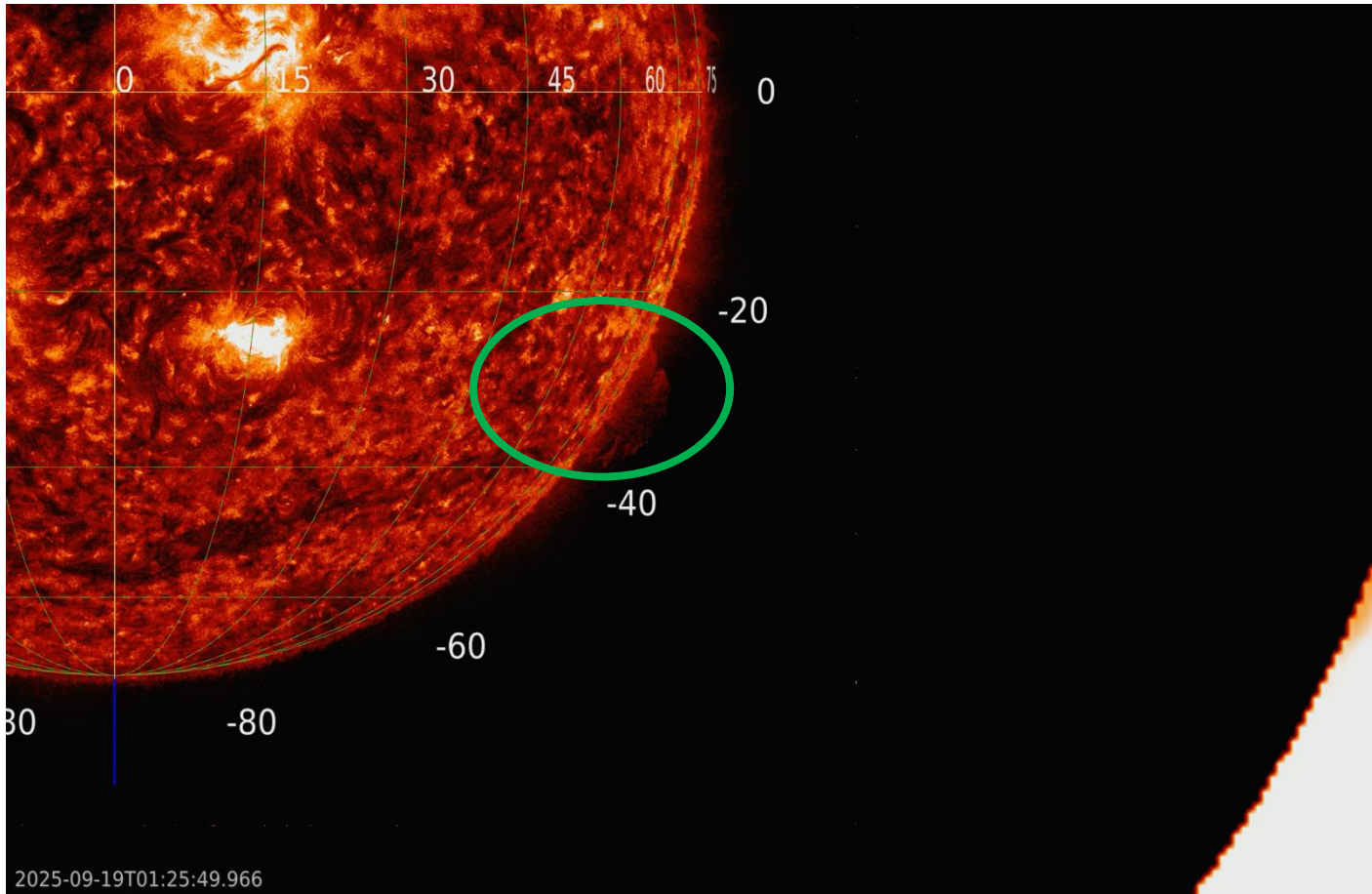
AIA + ASPIICS + SOHO/LASCO C2 + C3

2025-09-21 21:30 (UTC)



- A STEREO A
- E Earth
- B BepiColombo
- P Parker Solar Probe
- S Solar Orbiter
- J JUICE

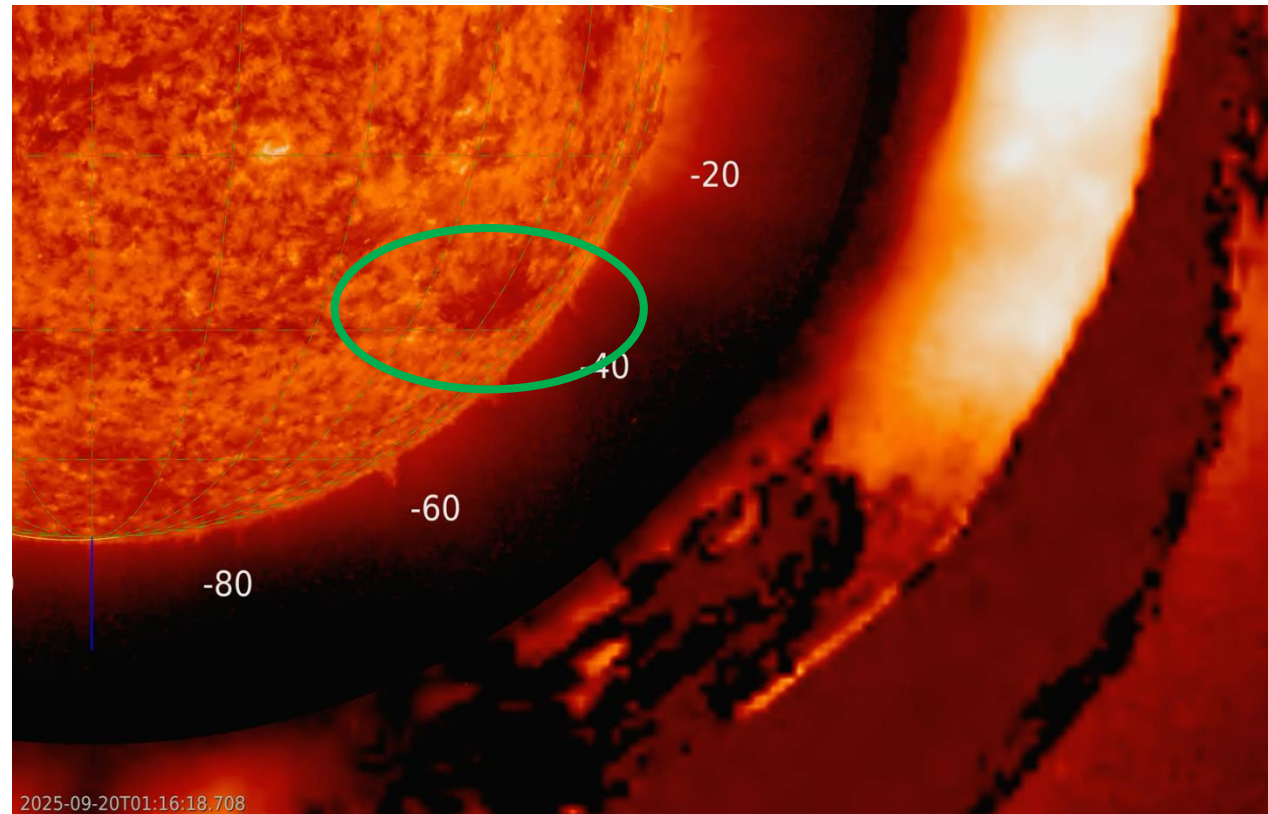
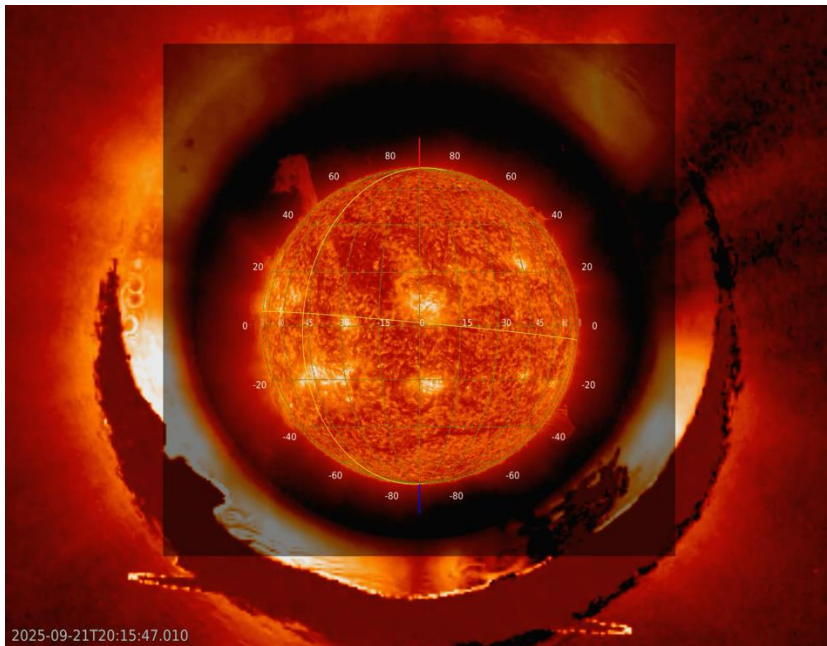
Eruption structure in the south



We observe the structure a few days before, also seen in STA

Stereo-A

EUVI 304 + COR1

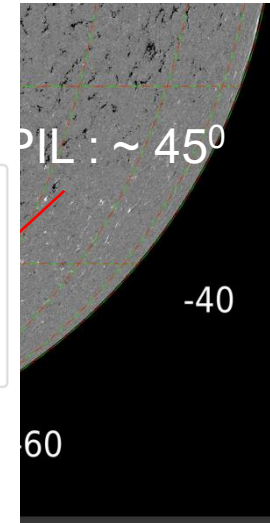
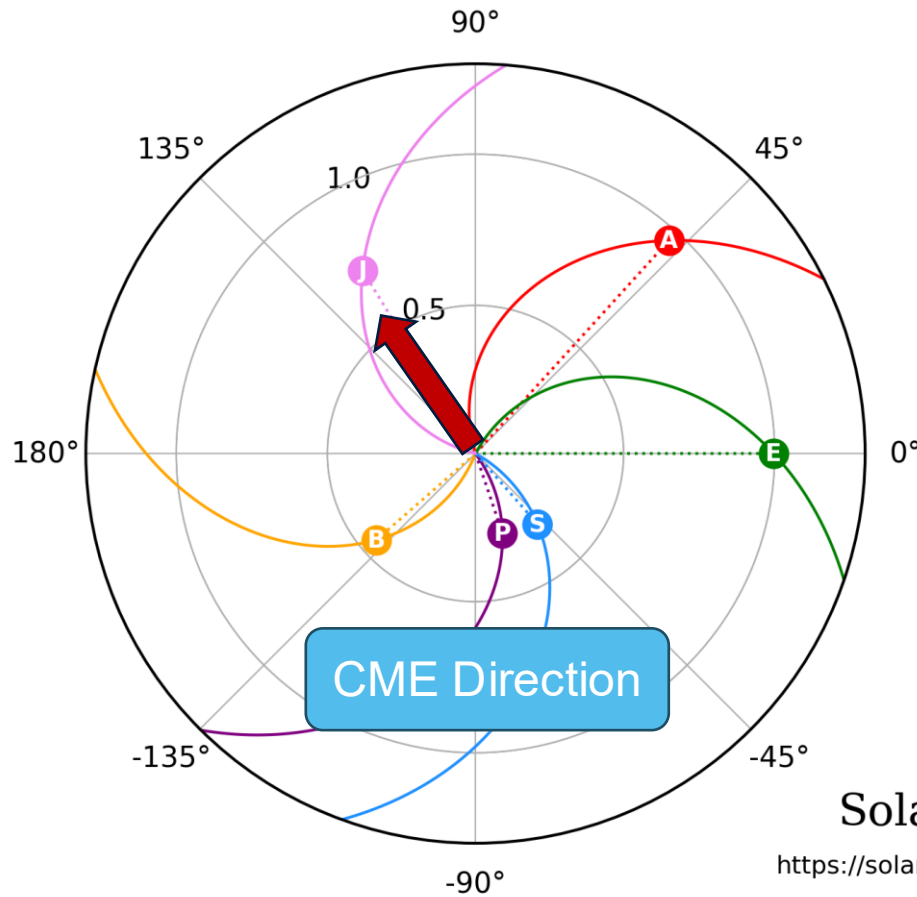
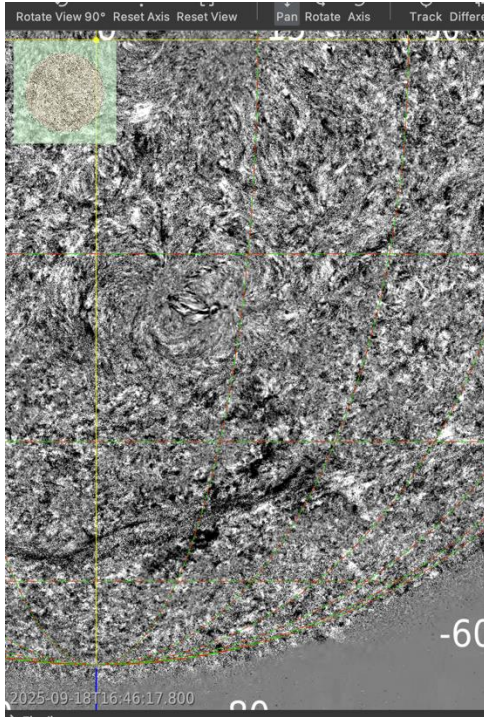


Base difference in EUVI 304

We can look at the magnetic field configuration of the structure a few days before the eruption while we can still see it

Parameter estimation

2025-09-21 21:30 (UTC)



We can assume positive helicity since it is in the south

On 16.09 14:31 it is at
On 21.09 21:31 it must

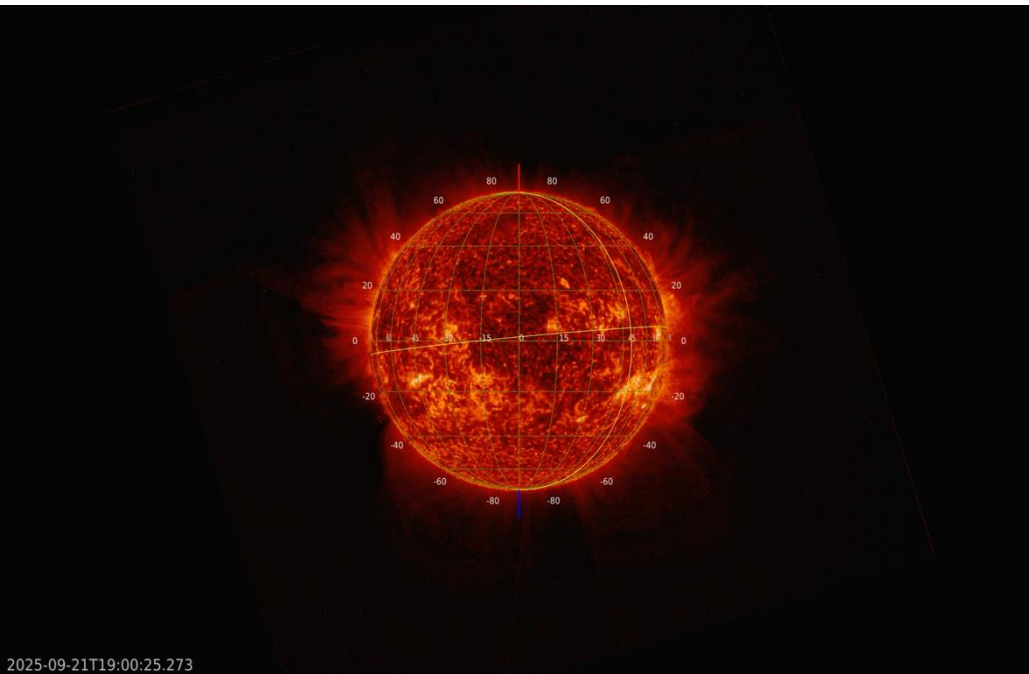
September 21	
hi, theta)	(130, -32)
	NA
	ESW/NES
	45 (-135 spheromak)
	600 km/s
At 21.5	22.09.2025 05:19

21.5RS Boundary Time: 2025-09-21 05:19:49Z

Keep Kept: 0 Start Ensemble

Speed At 21.5

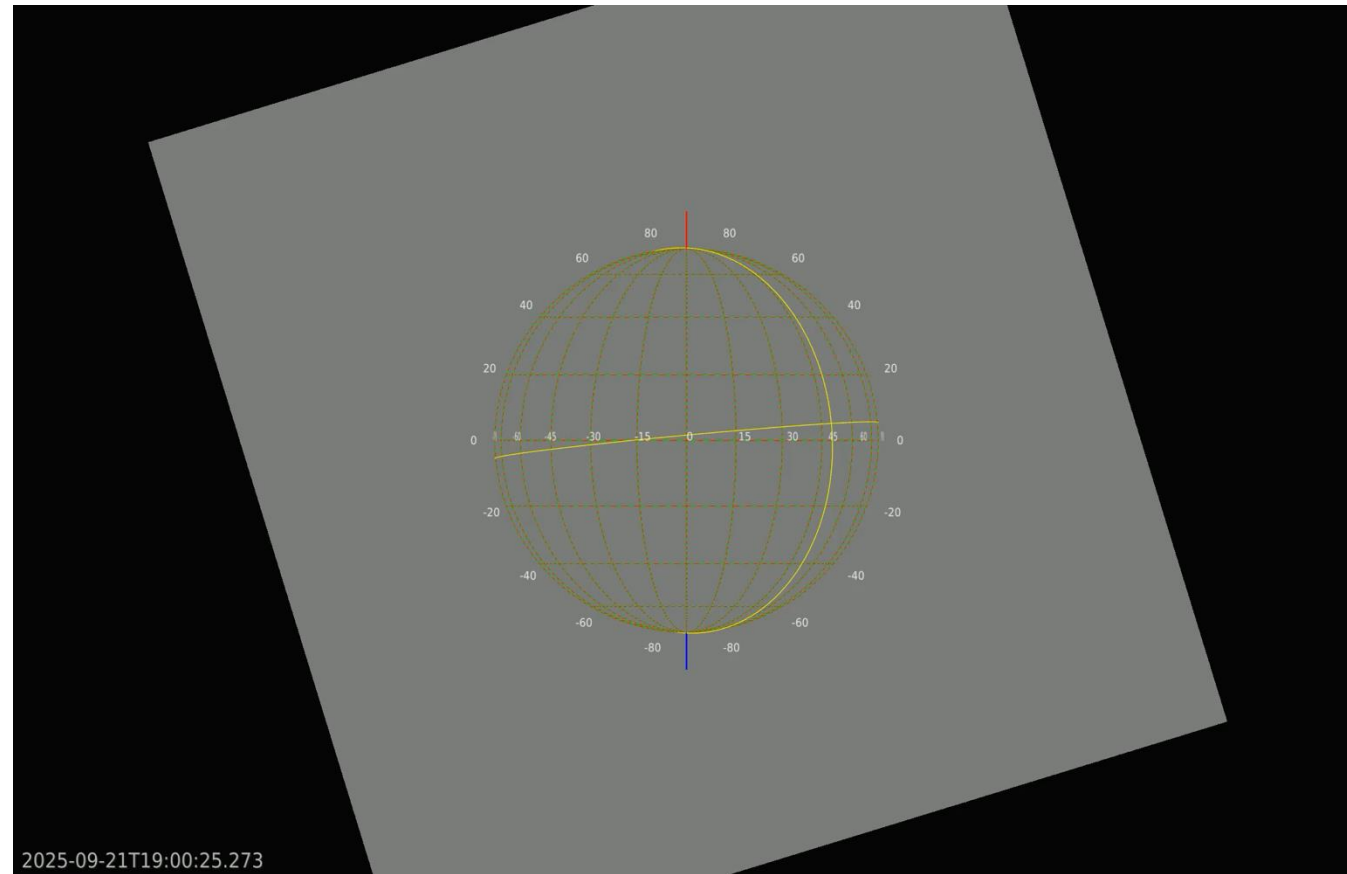
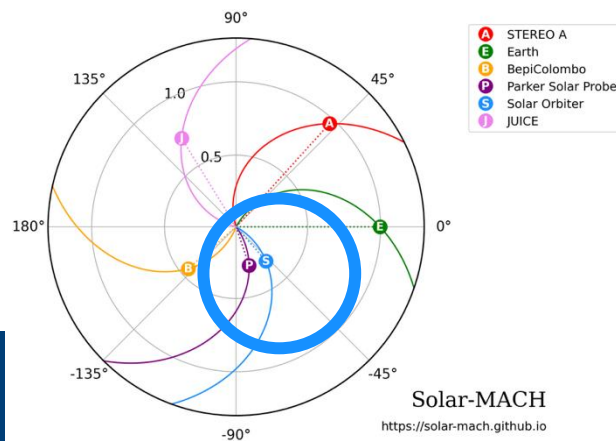
Solar Orbiter



2025-09-21T19:00:25.273

2025-09-21 21:30 (UTC)

EUI 304



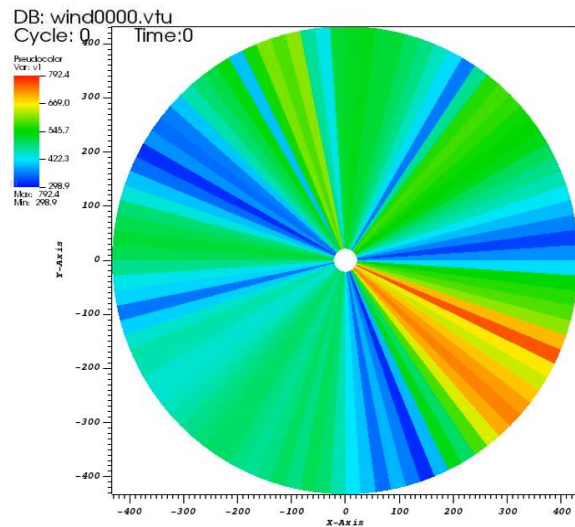
2025-09-21T19:00:25.273

Difference in EUI 304

Icarus simulations

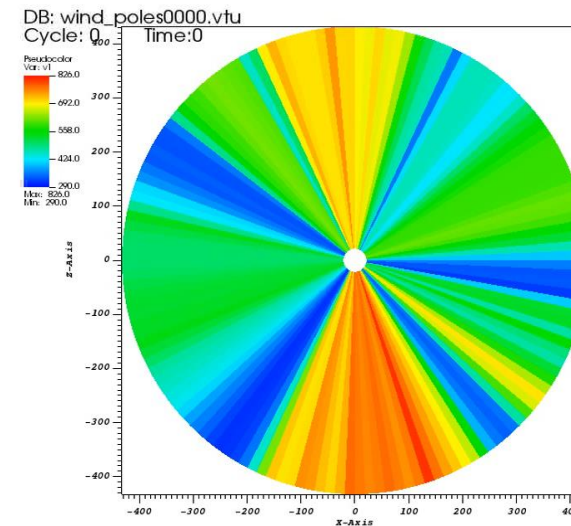
- First, time-dependent background solar wind modelled
 - From 10.09.2025 – 30.10.2025 – 1h cadence
- We included the poles since the CMEs have high altitudes

Equatorial plane



user: baratash
Fri Apr 10 12:08:25 2026

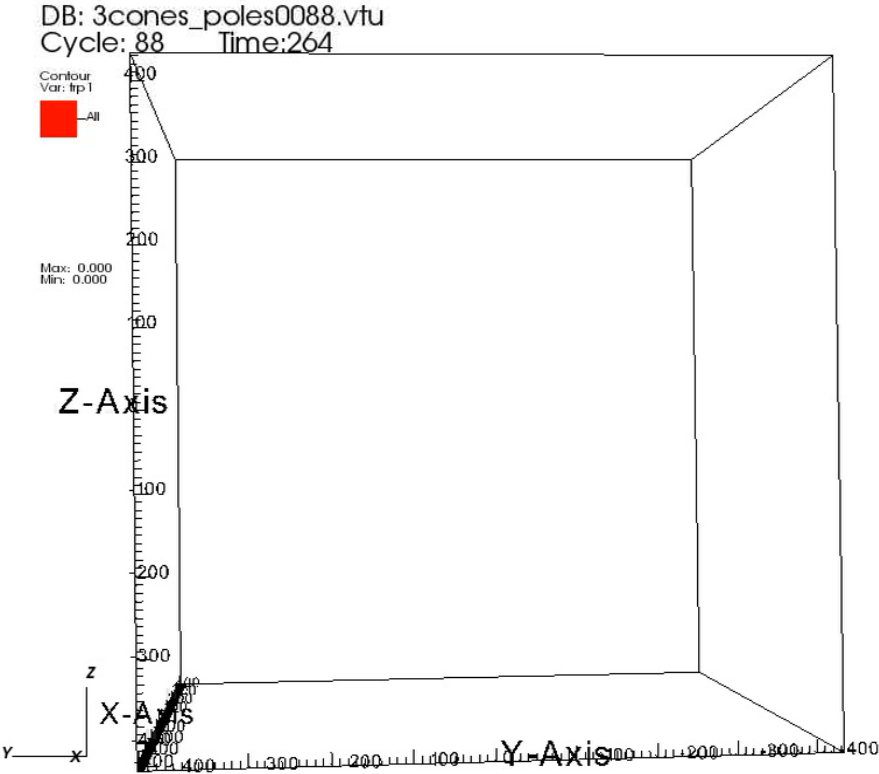
Meridional plane



user: baratash
Mon Apr 13 11:22:46 2026

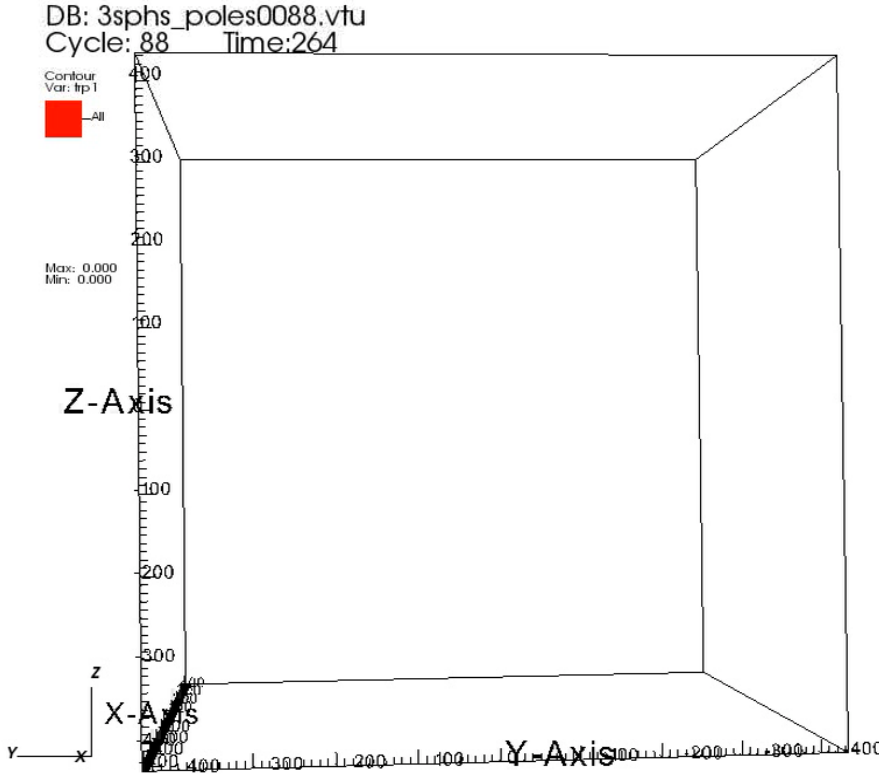
Cone and Spheromak CMEs injected

Cone: CME density



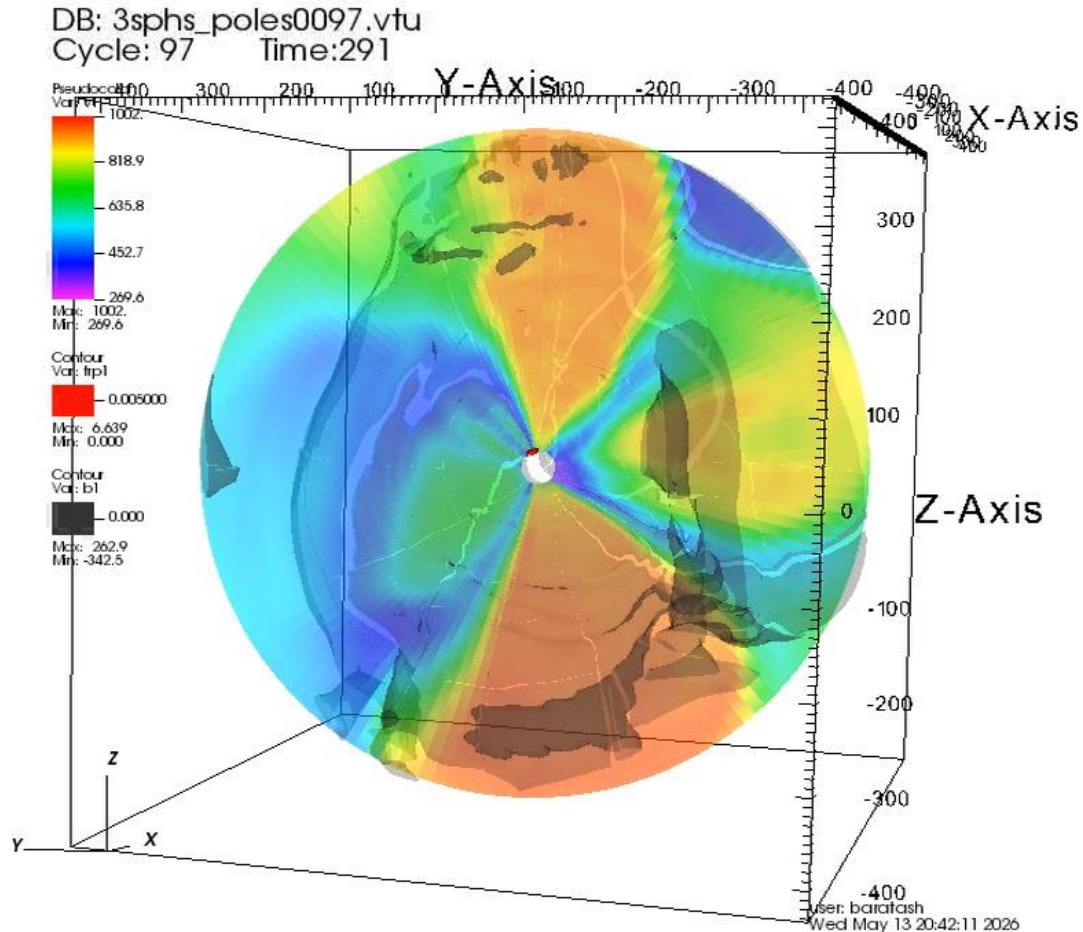
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Wed May 13 17:41:10 2026

Spheromak: CME density



user: baratash
Wed May 13 17:33:07 2026

Spheromak CME in the meridional plane

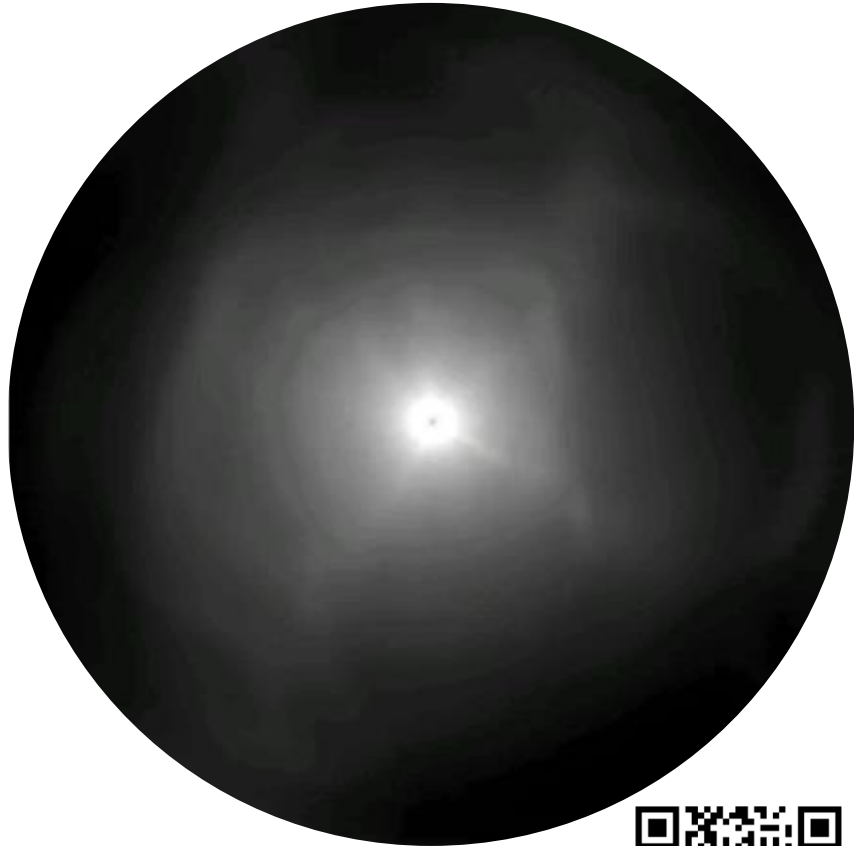


- Background – radial velocity
- Heliospheric current sheet – shaded grey
- Contours of the CME density - red

Note* the movie is in the co-rotating frame fixed to earth location when CMEs erupt

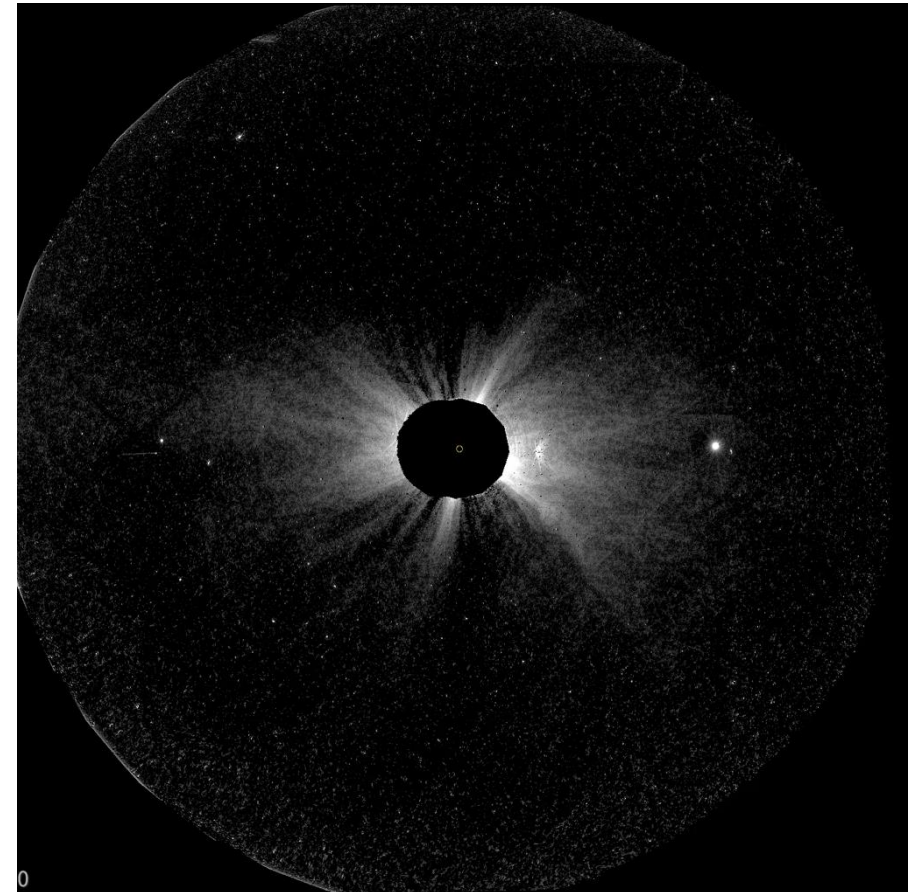
Synthetic white-light image – cone CME

PUNCH – PAM – V0K



FoMo model

[Van Doorselaere et al. 2016](#)



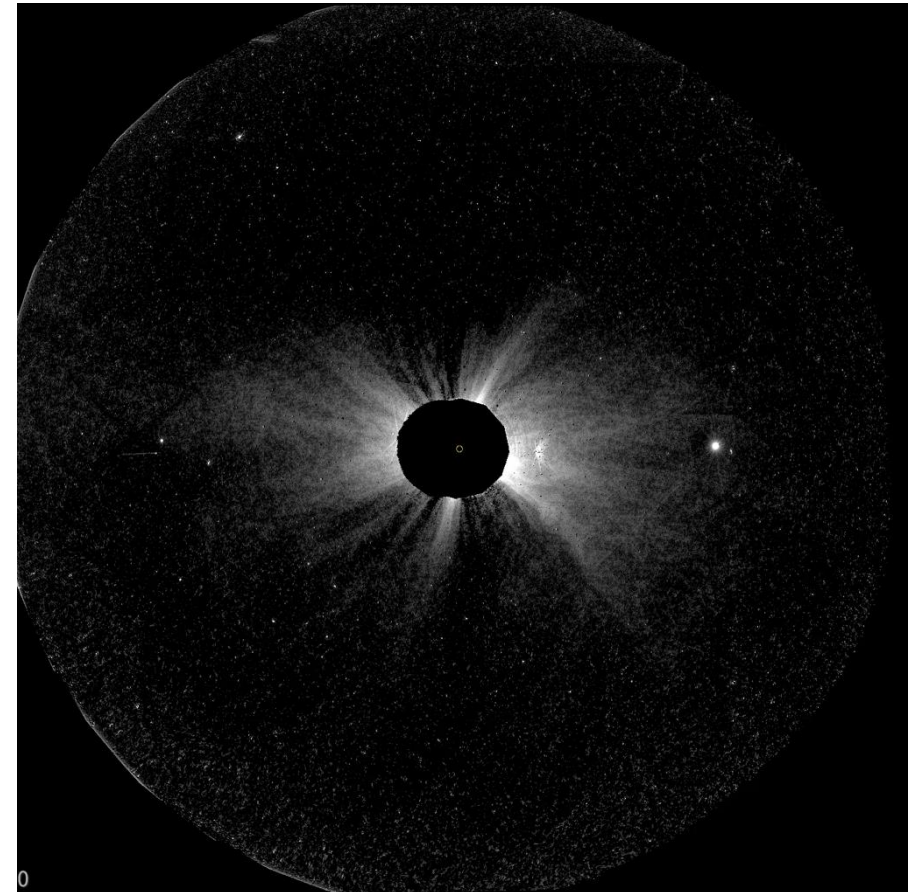
Synthetic white-light image – spheromak CME

PUNCH – PAM – VOK



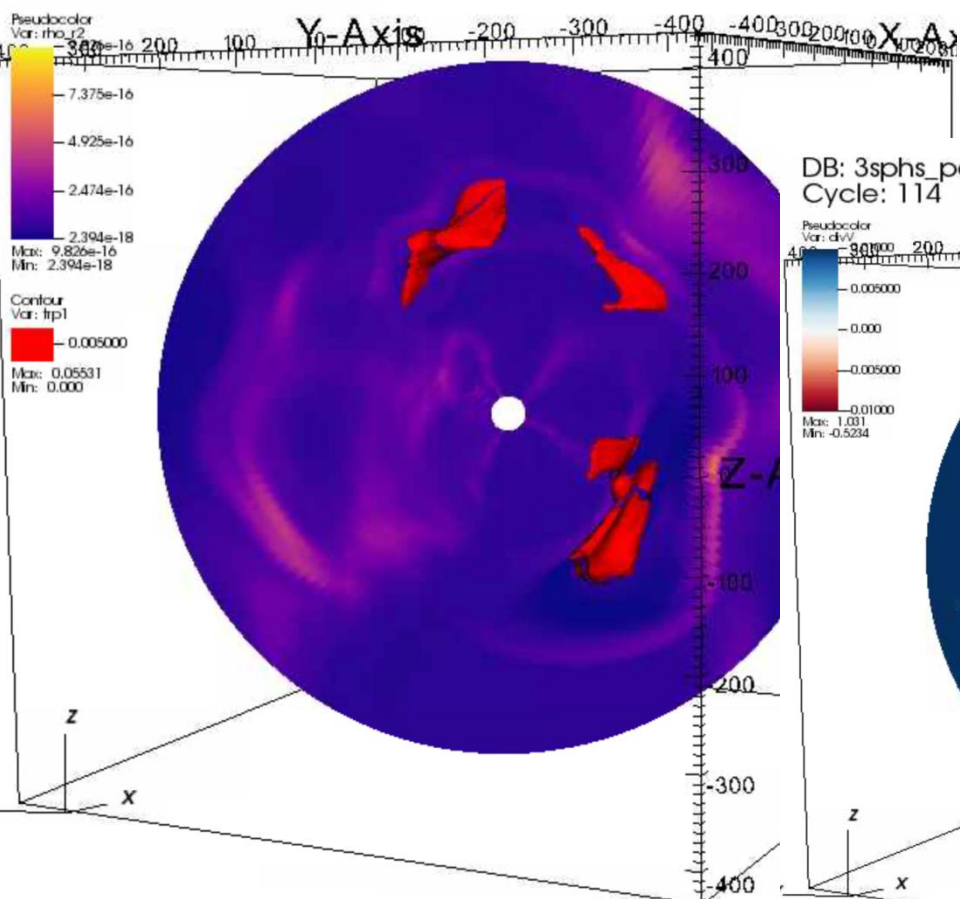
FoMo model

[Van Doorselaere et al. 2016](#)

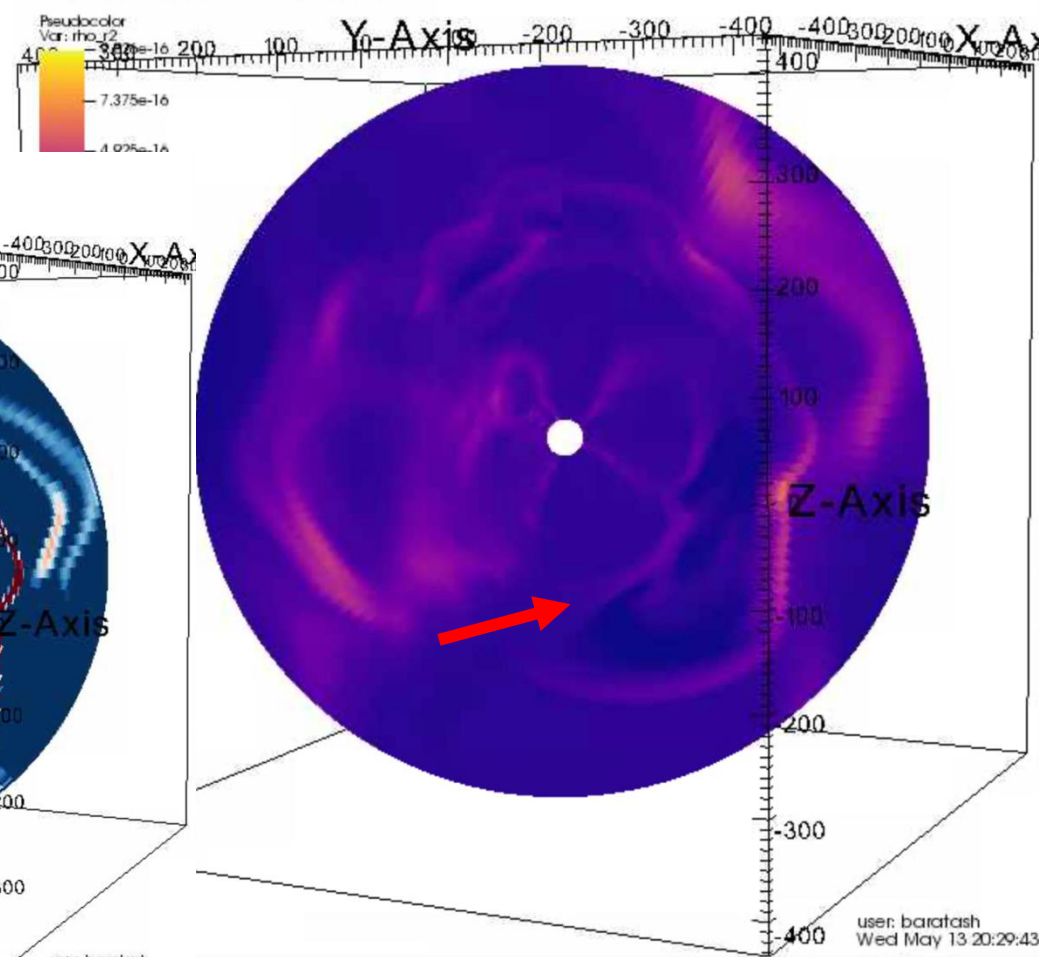


Enhanced density behind the CME

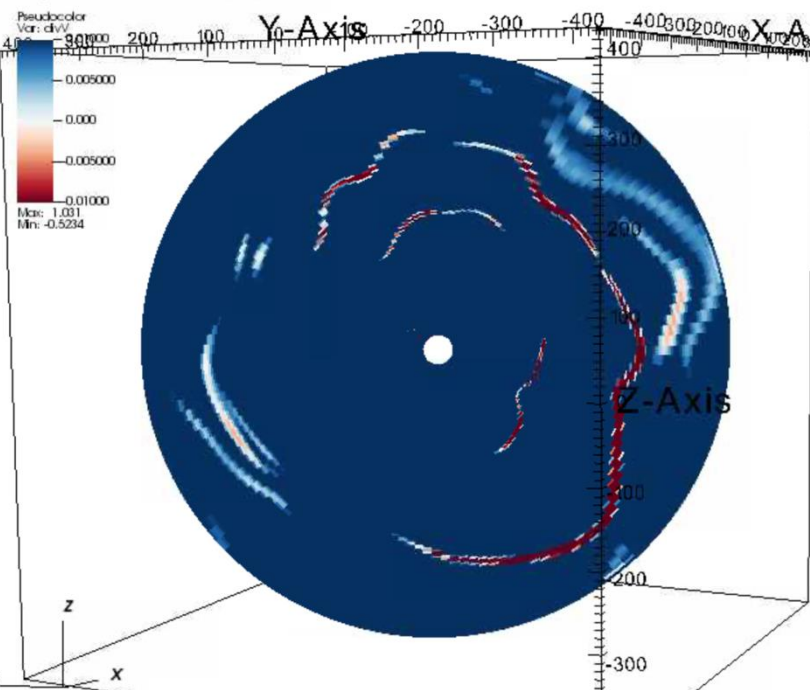
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Cycle: 114 Time:342



DB: 3sphs_poles0114.vtu
Cycle: 114 Time:342



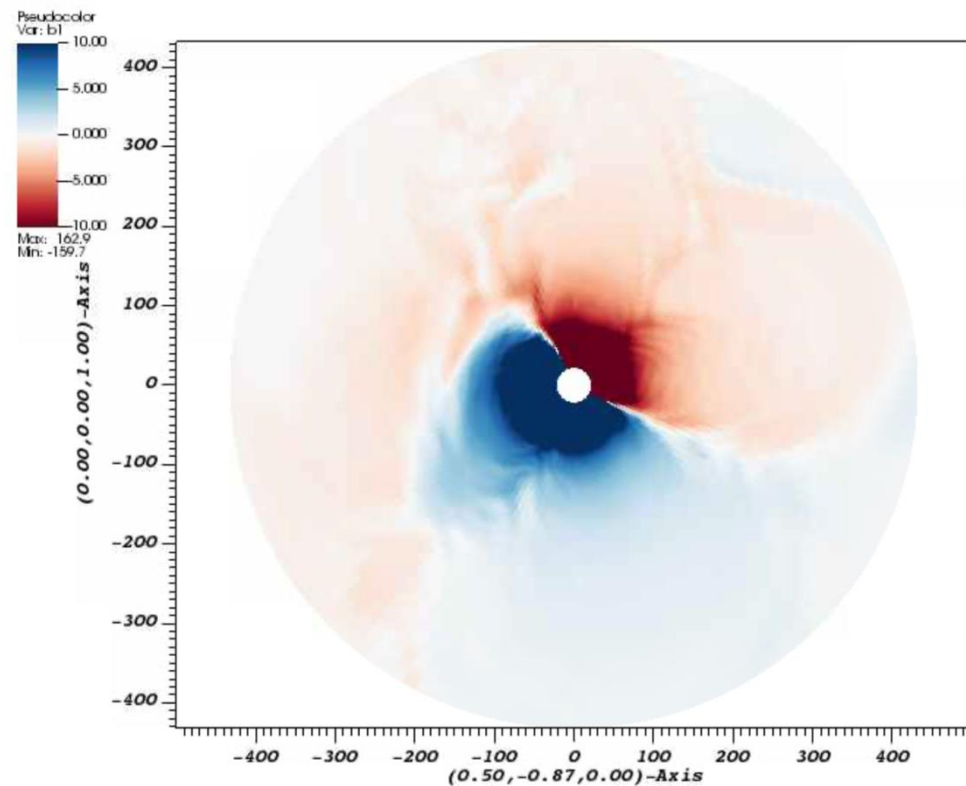
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Wed May 13 20:24:15

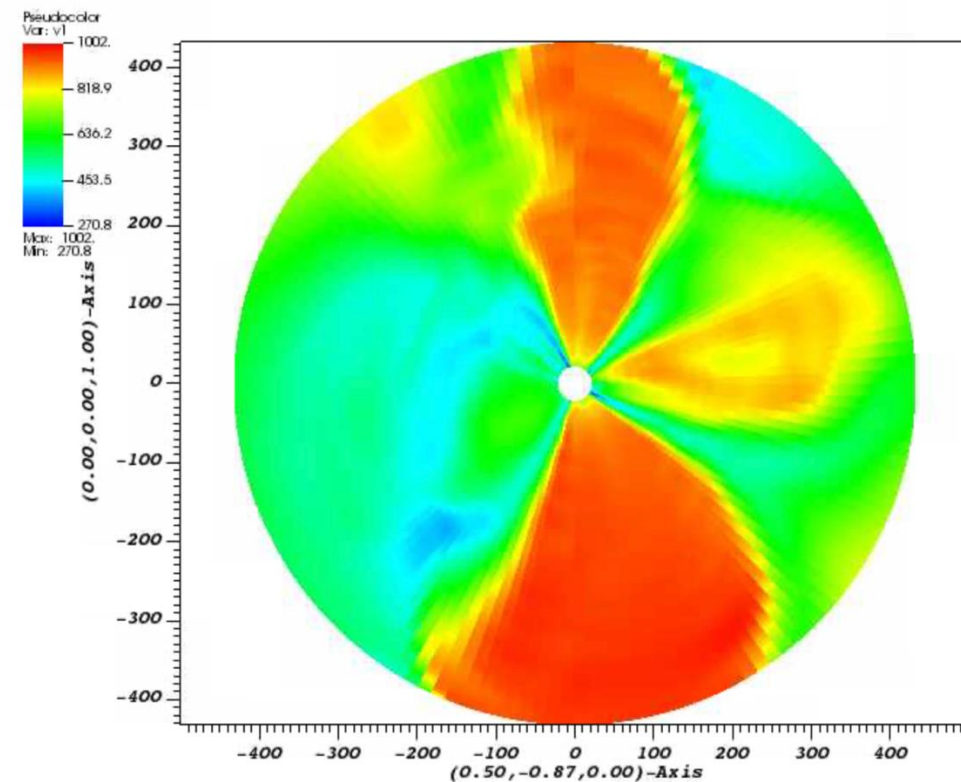
user: baratash
Wed May 13 20:29:43

DB: 3sphs_poles0095.vtu
Cycle: 95 Time:285



user: baratash
Wed May 13 18:49:41 2026

DB: 3sphs_poles0095.vtu
Cycle: 95 Time:285



user: baratash
Wed May 13 18:47:52 2026

Summary and future prospects

Thank you!

- We are injecting FRI3D CME model which has a more complex and realistic structure
- We are connecting Icarus to FORWARD WL tool
- We are generating movies in inertial frame to compare with PUNCH
- Gradually add CMEs in September – October 2025

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