

## Planets — problems (F. Bagenal)

### 1. Minimum Mass Solar Nebular

- (a) Estimate the amount of material needed to make a solar system.
- (b) Mass is not everything — check you have enough:
  - Rock/metal to make the terrestrial planets
  - Hydrogen to make Jupiter
  - water/ammonia/methane to make Uranus & Neptune
- (c) Estimate the mass density of the nebular

### 2. Collapse time

Make a rough estimate of the minimum time to collapse the solar nebula.

*Hint: Think of a parcel of gas as an object orbiting the Sun on a very eccentric orbit. Think Kepler's 3<sup>rd</sup> Law.*

### 3. Mass accretion

- (a) Think of a sphere orbiting the Sun and accreting solar nebular mass. Show how the mass accretion could be approximated

$$\frac{dM}{dt} \sim \pi R^2 v_{\text{orbit}} \rho_{\text{nebula}} .$$

- (b) Integrate to derive the time to accrete to radius  $R$

$$T \sim \frac{4R}{v_{\text{orbit}}} \left( \frac{\rho_{\text{planet}}}{\rho_{\text{nebula}}} \right) .$$

- (c) Use the following to compute the time required to accrete a 1000 km object.
  - $v_{\text{orbit}} \sim 30 \text{ km/s}$  at 1 AU
  - $\rho_{\text{nebula}} \sim 10^{-7} \text{ kg m}^{-3}$ .
  - $\rho_{\text{planet}} \sim 3000 \text{ kg m}^{-3}$ .

### 4. Heating & Cooling Planets

sketch profiles of interior temperature of a terrestrial planet that

- (a) accreted quickly with no radioactivity.
- (b) only had radioactive heating
- (c) accreted very slowly.

*Hint: think of conducting heat from the interior, and cooling from the surface (via radiation).*