

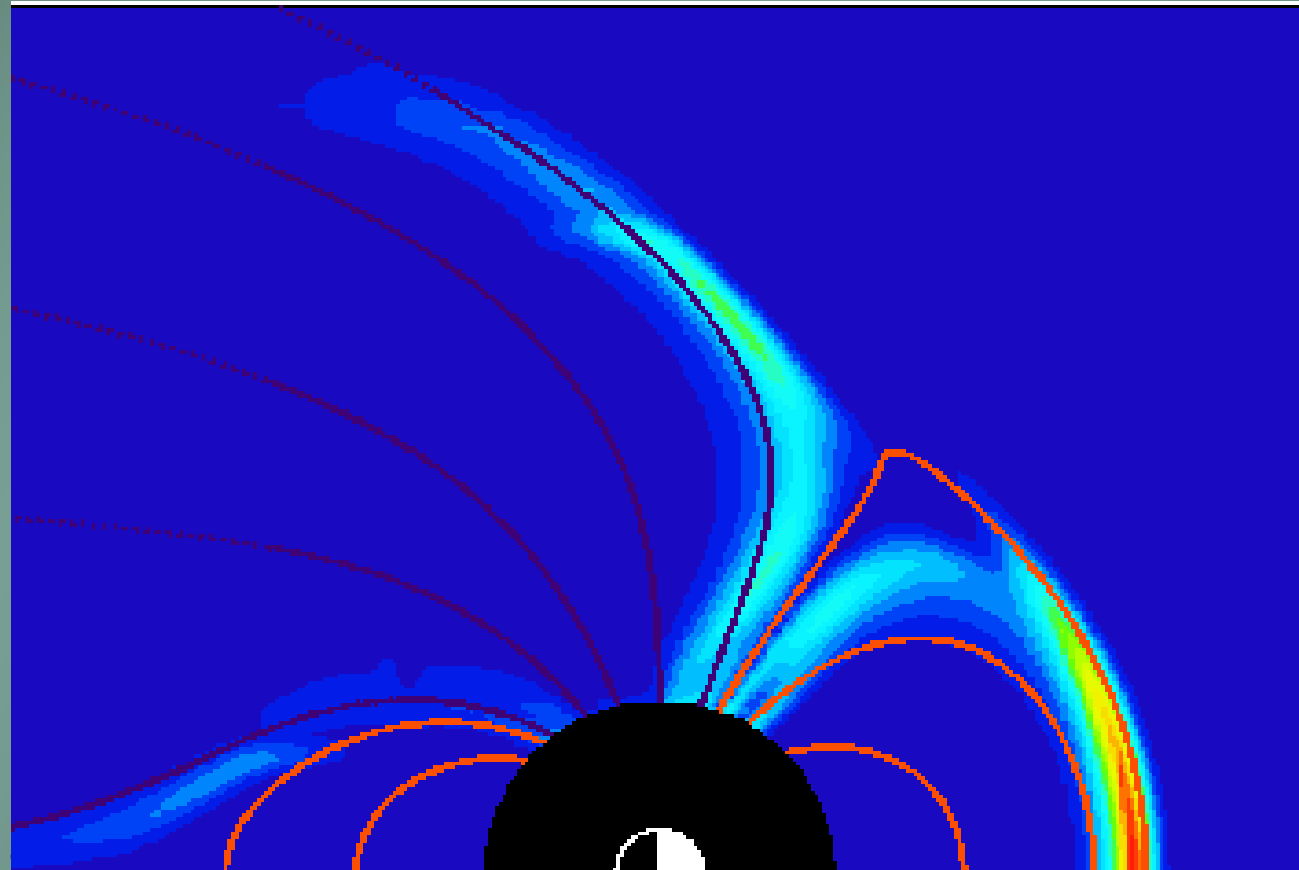


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The Chapman-Ferraro Problem

The ~~CCMC~~ Solution

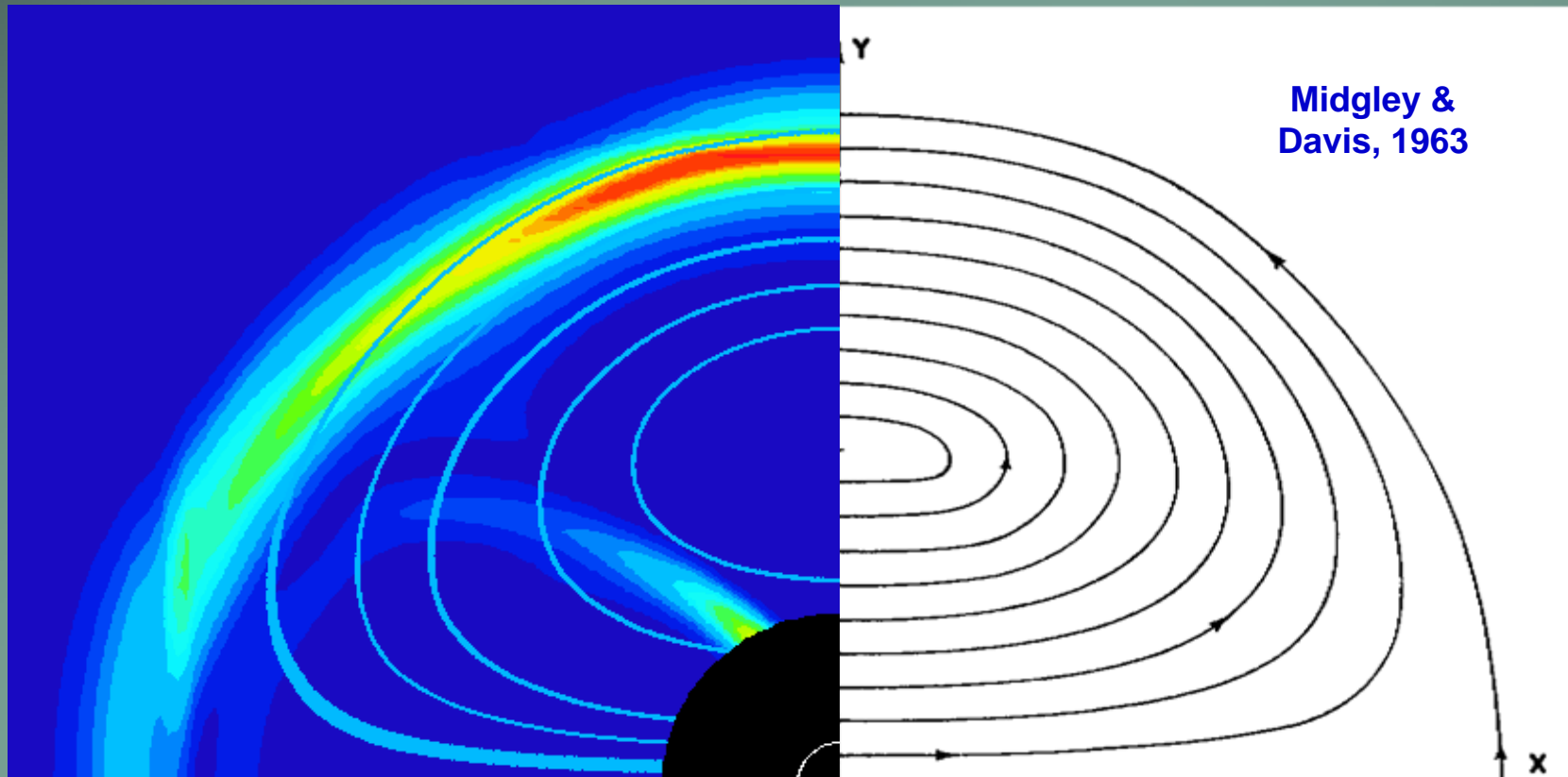
Find the size and shape of the boundary between the solar corpuscular radiation and the earth-enclosing cavity it forms, and determine the magnetic field everywhere within the cavity.



The Chapman-Ferraro Current System



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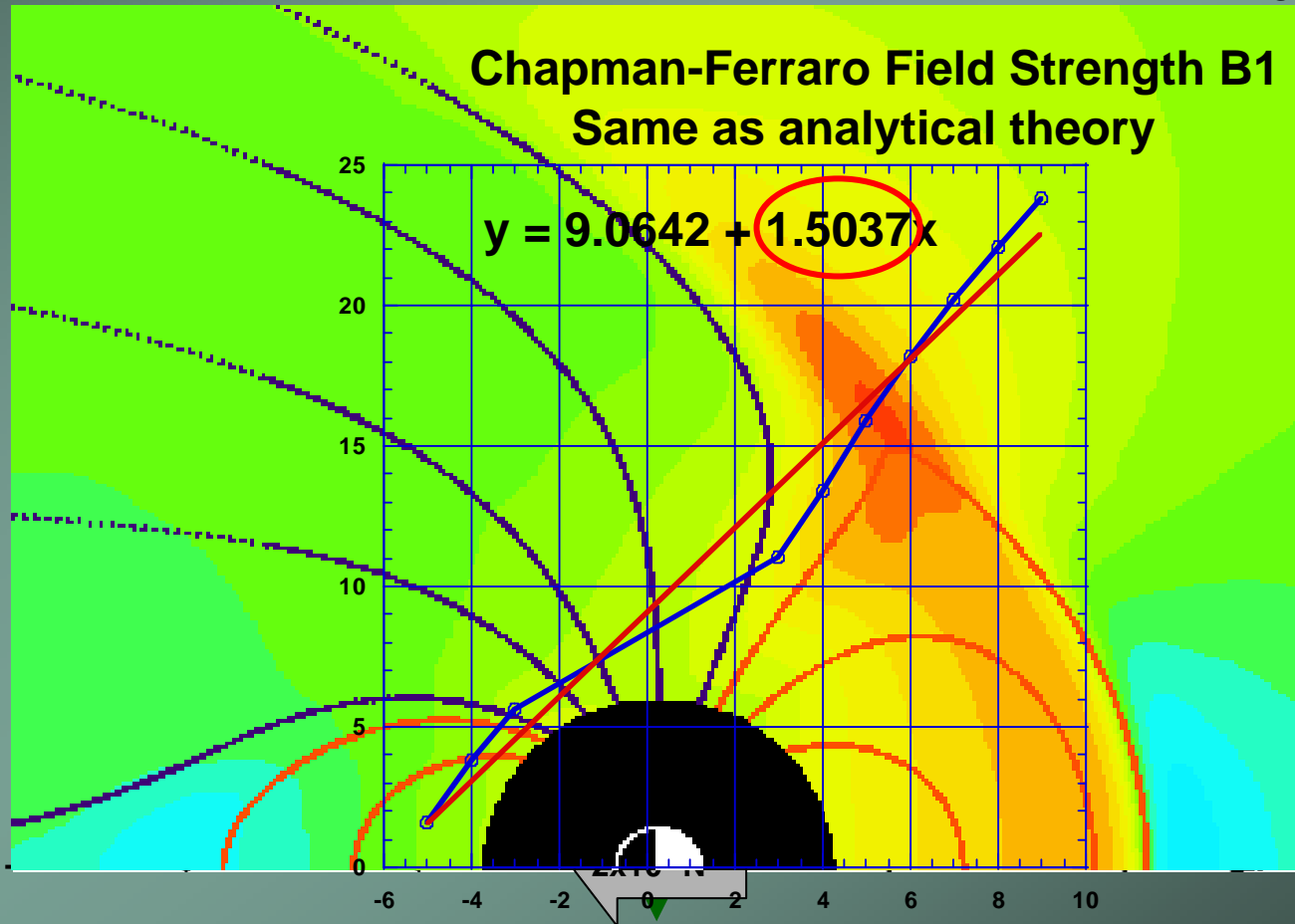


The Chapman-Ferraro Force

The Analytical Solution



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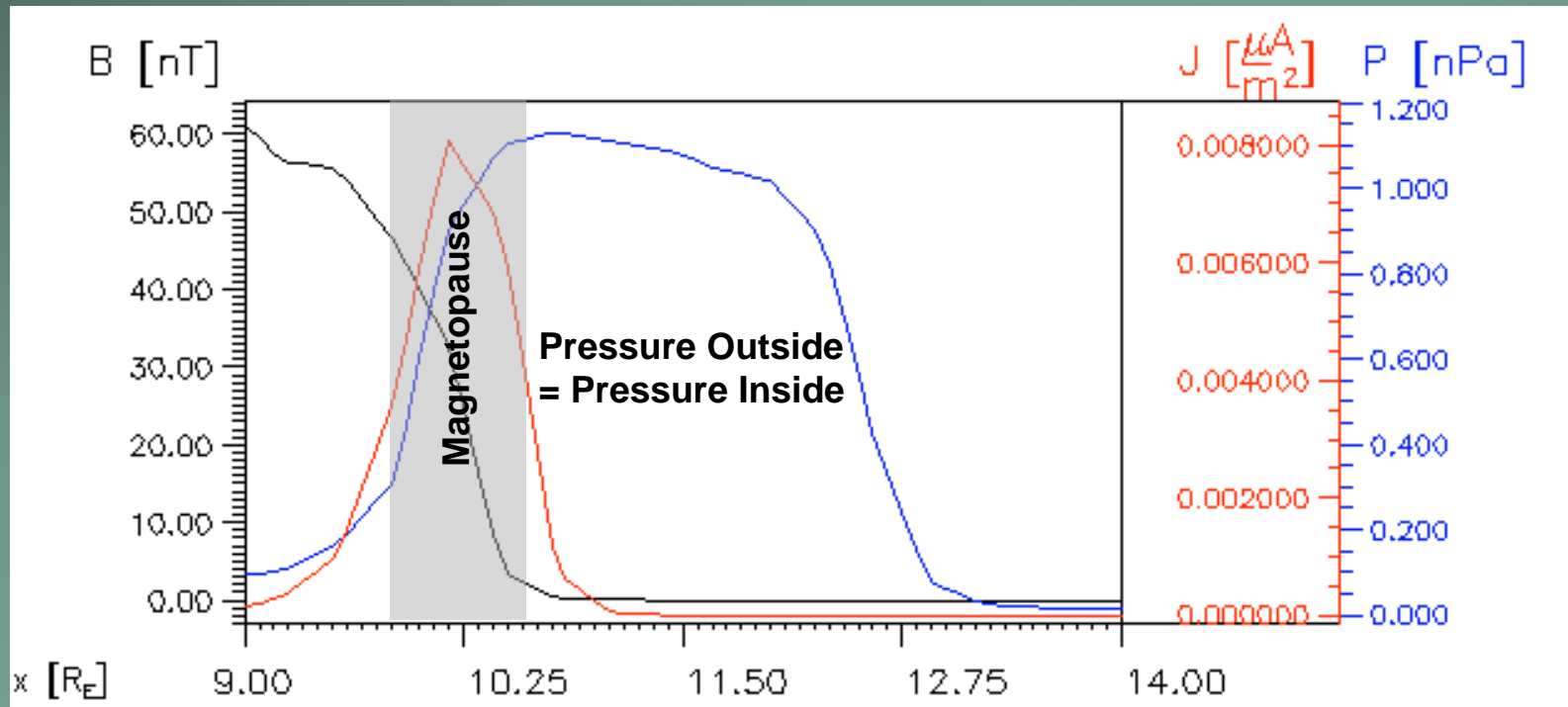
Stagnation Streamline Parameters Hydrodynamic (Zero IMF) Case



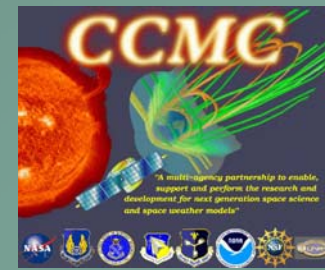
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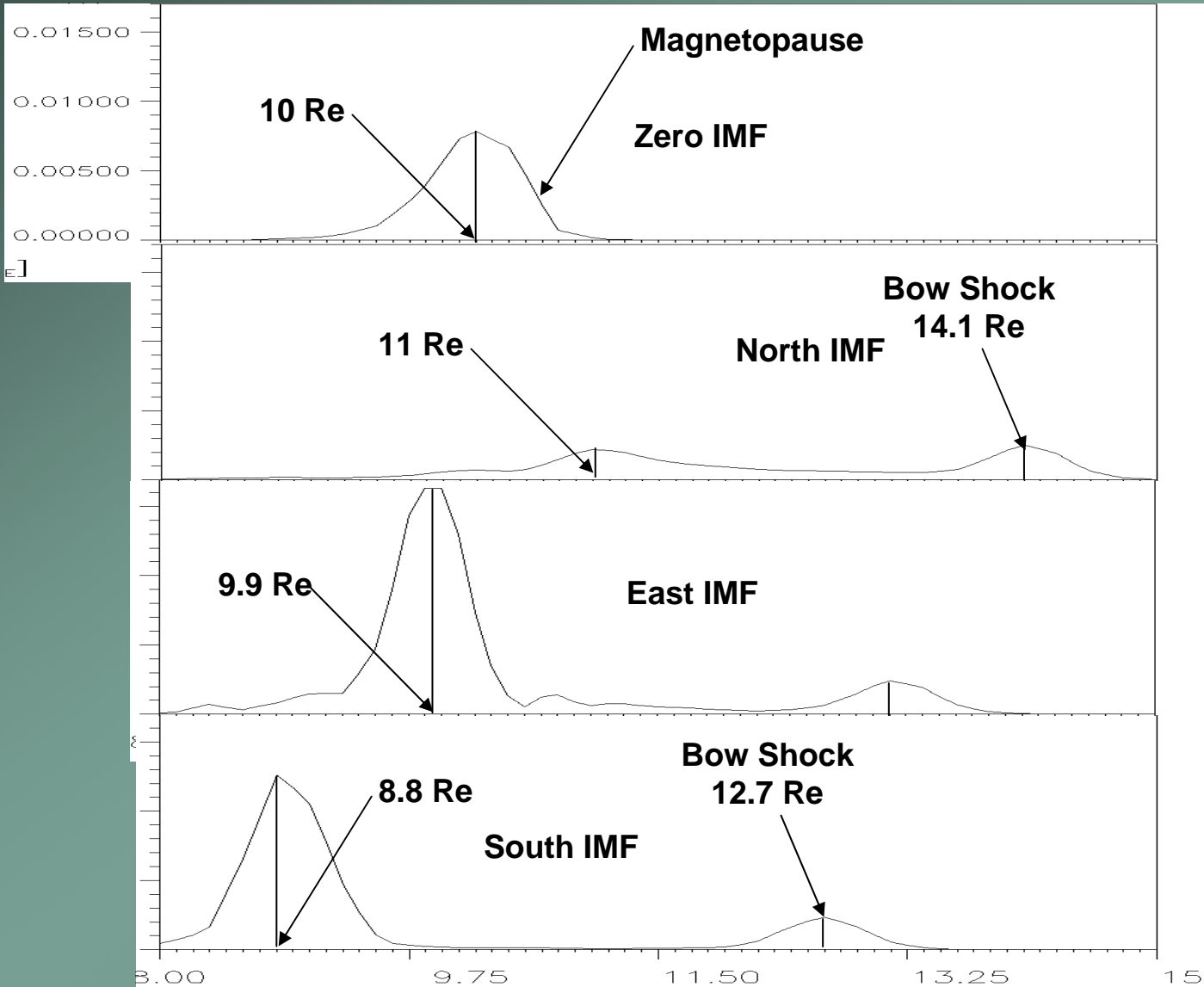


Variation with IMF Orientation Distance to Stagnation Point



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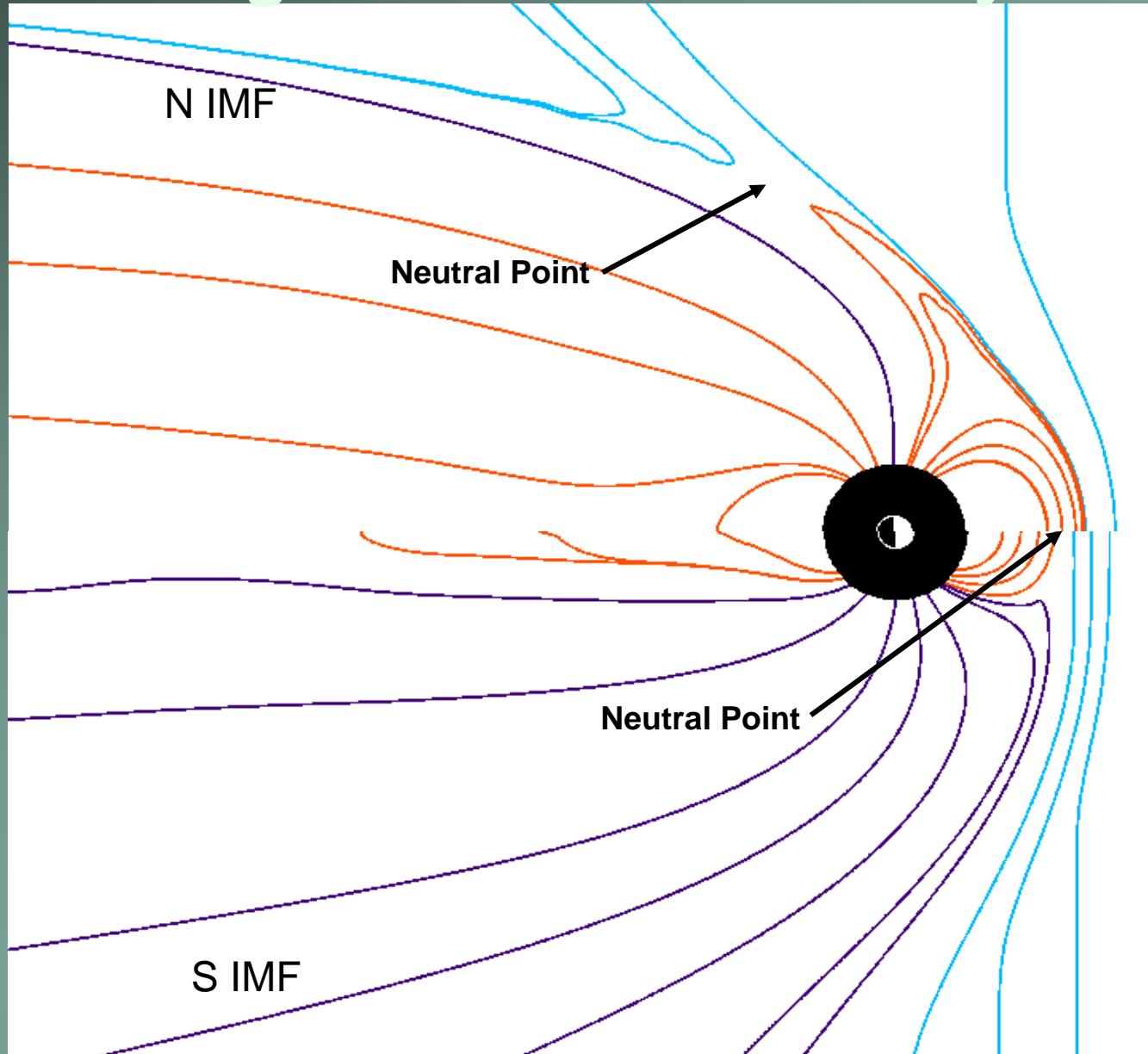
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Standoff Distance =
3.1 Re

Standoff Distance =
3.9 Re

Variation with IMF Orientation Magnetic Field Geometry



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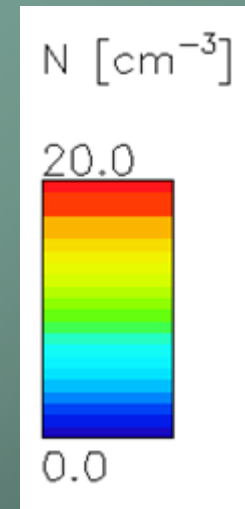
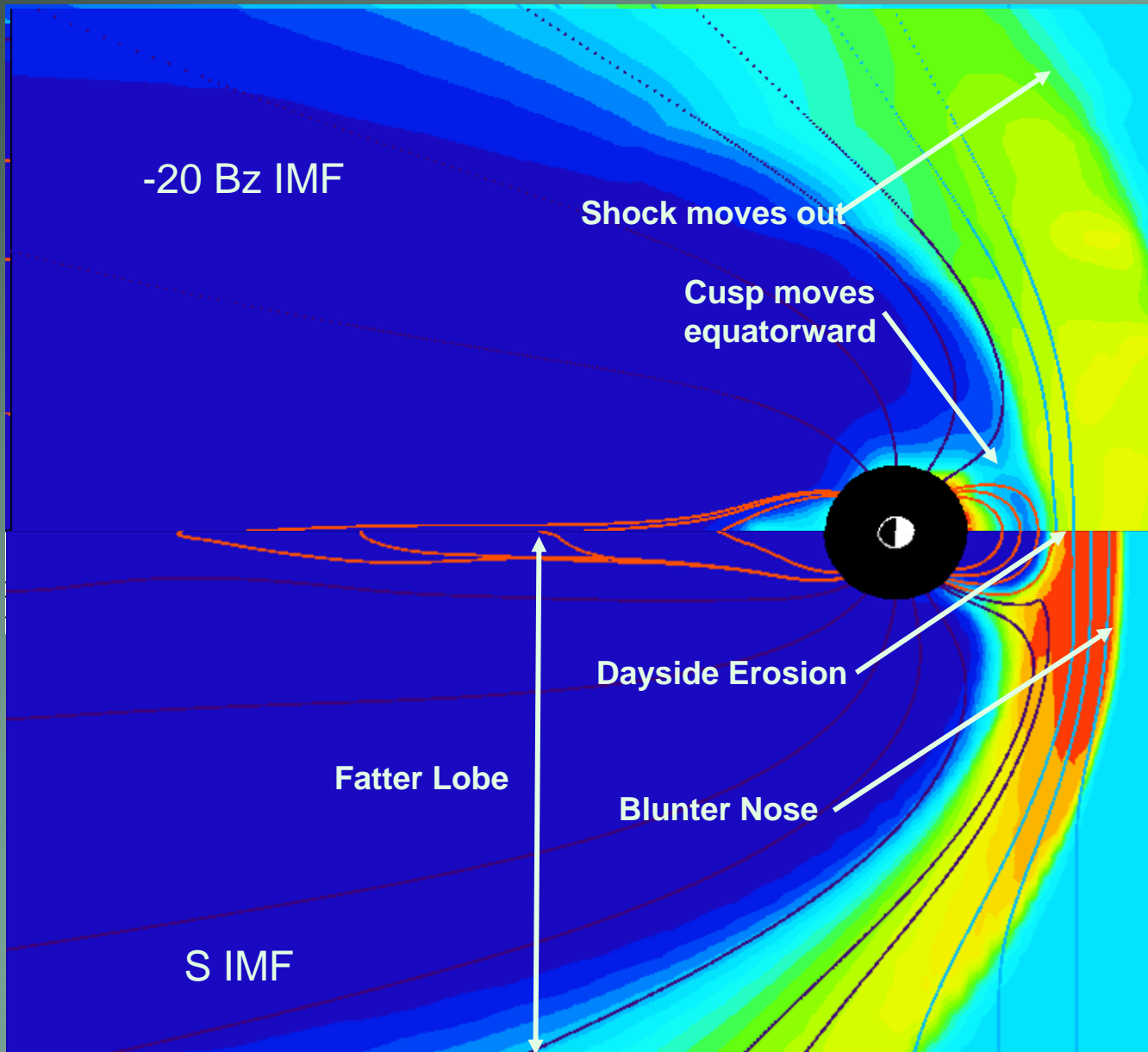
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Magnetosphere Shape

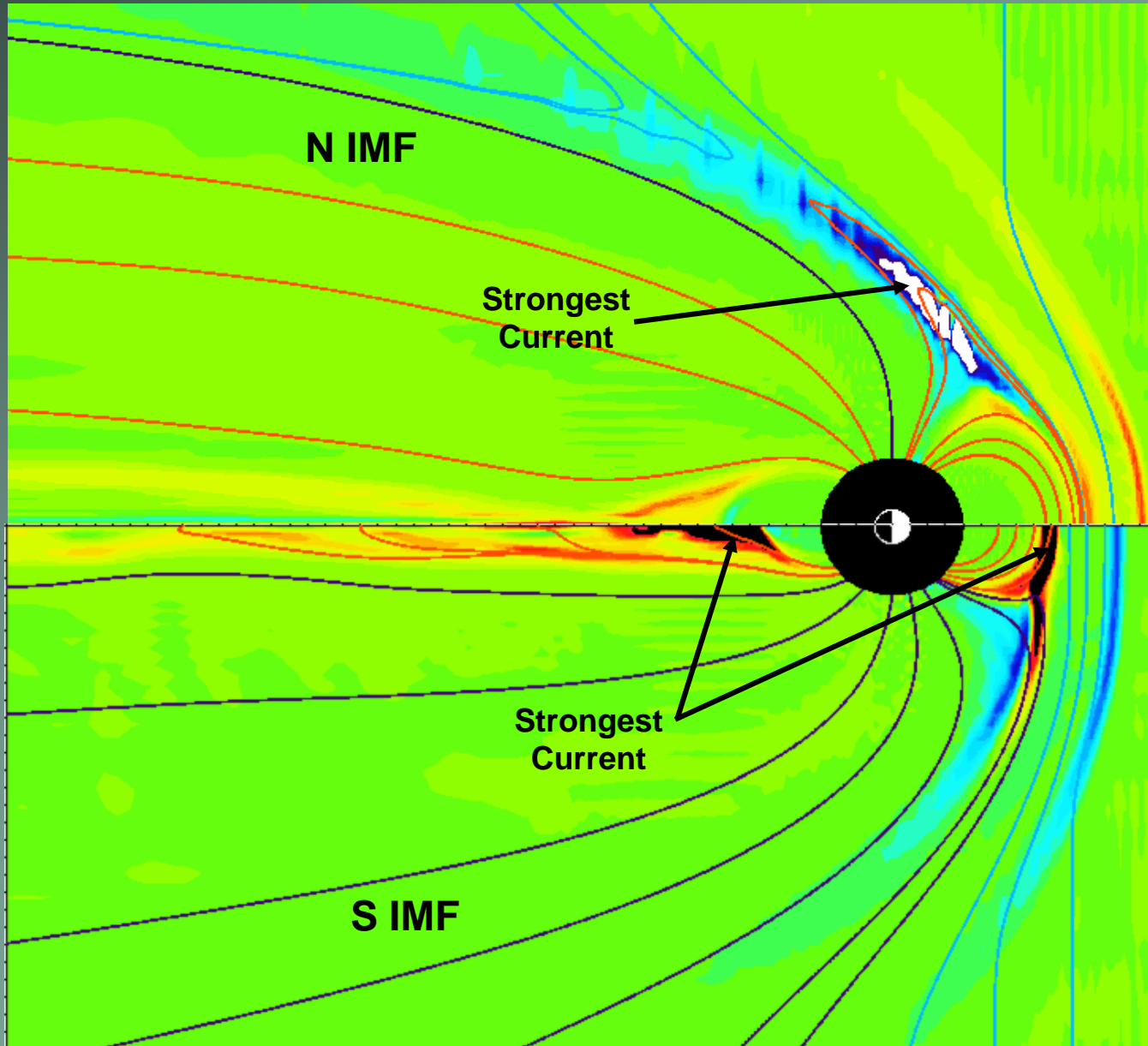


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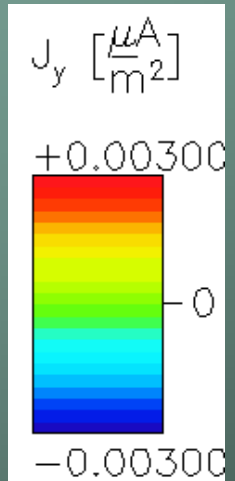


Magnetopause Current

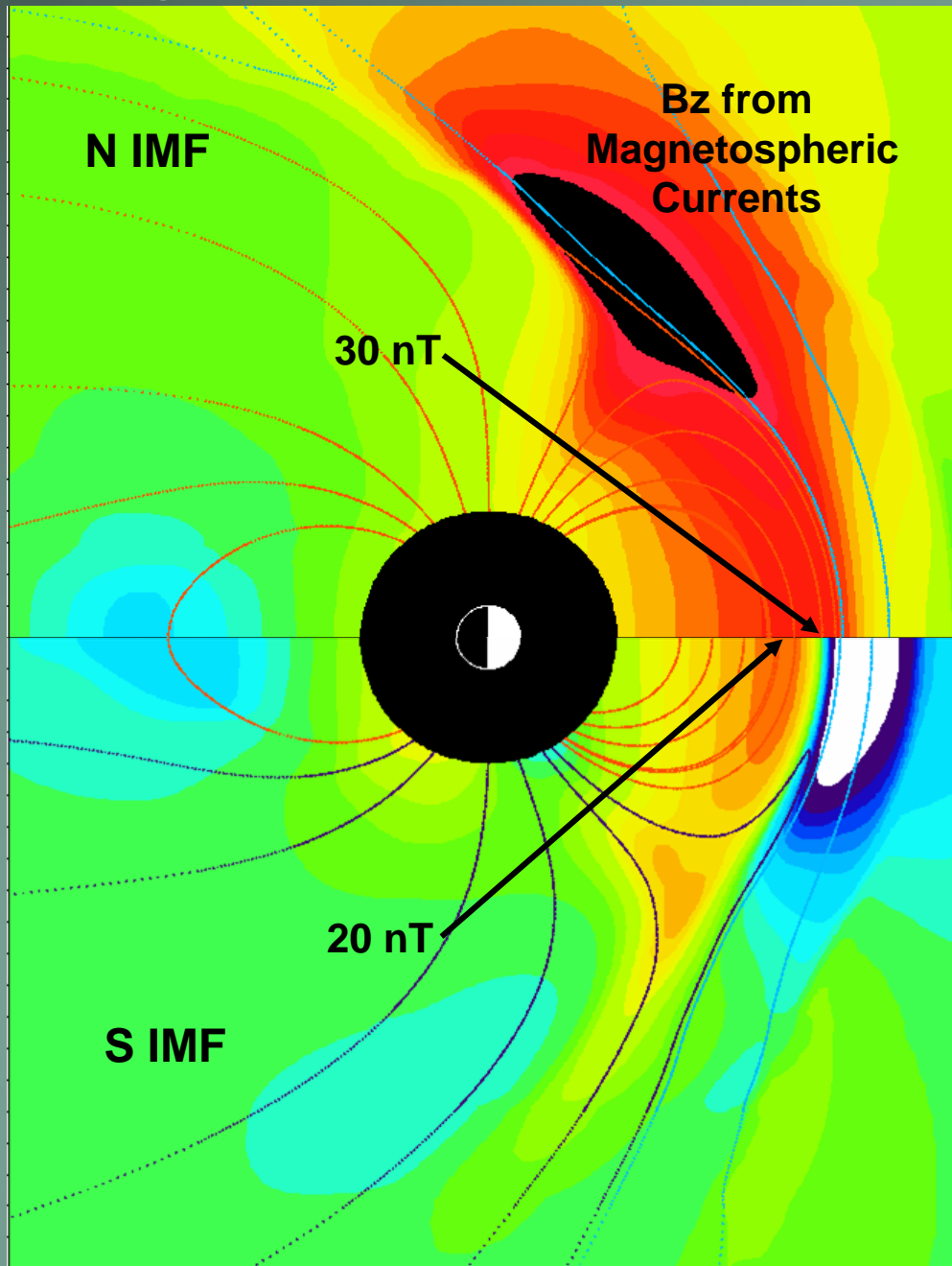


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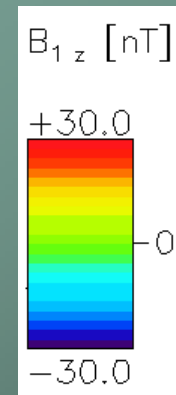
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Dayside Compression

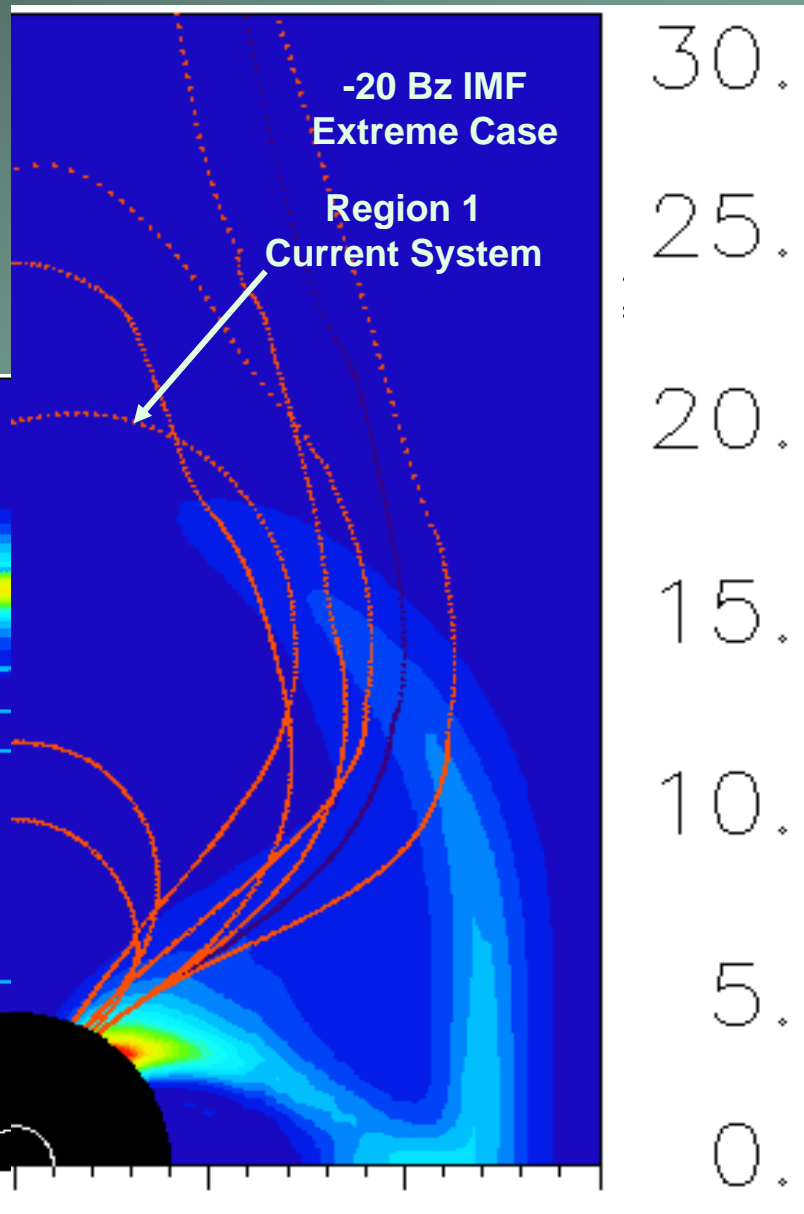
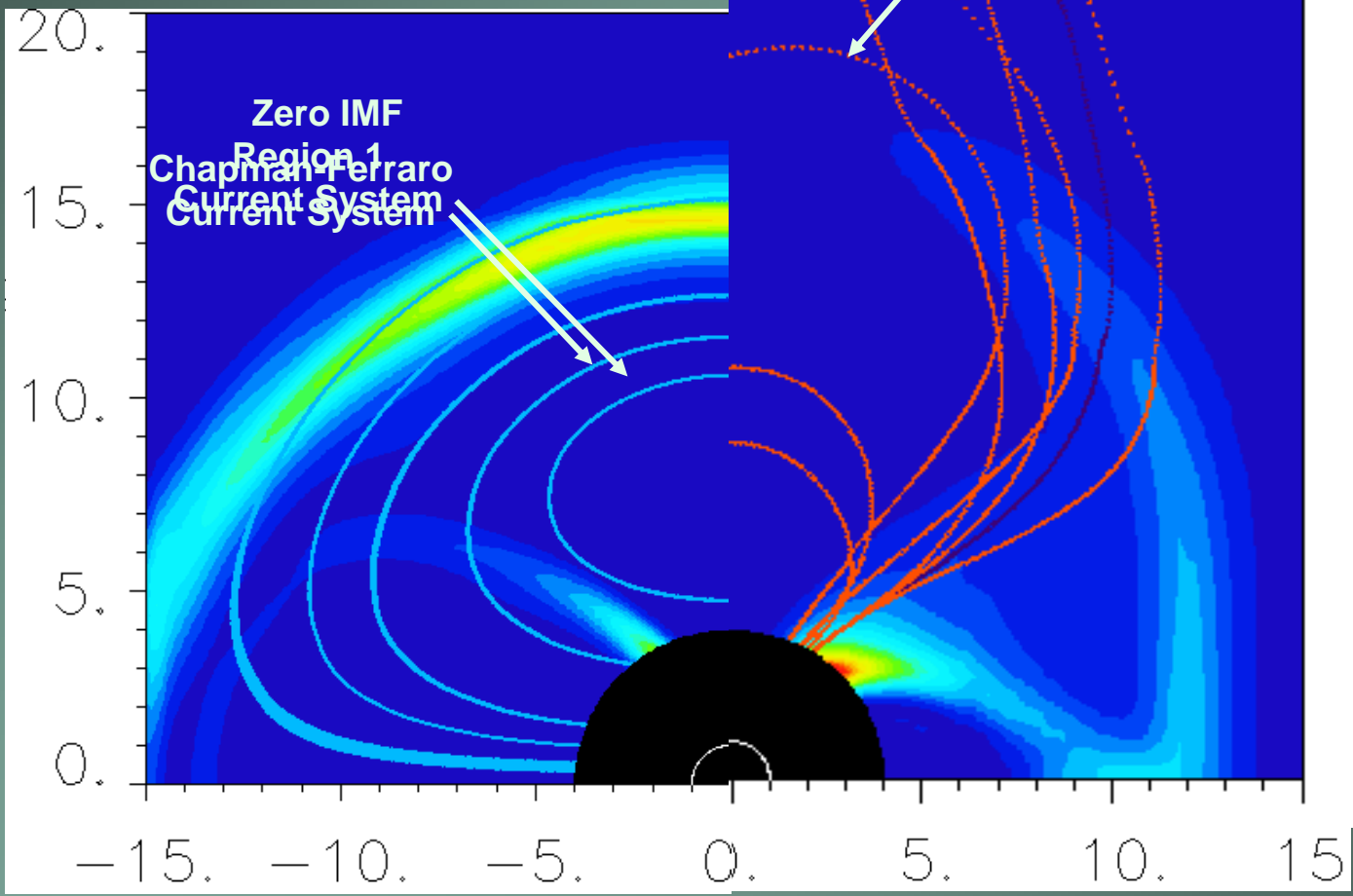


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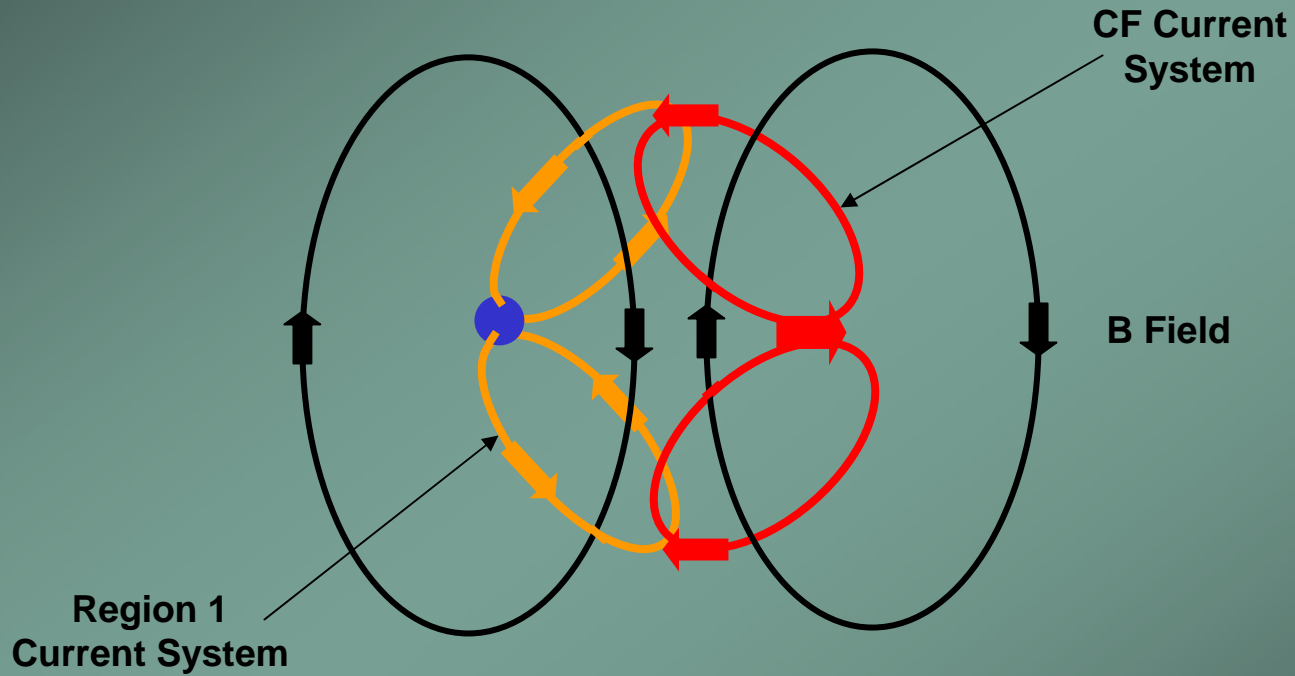
South IMF reduces compression

Strong IMF conditions Compress the current sheet?



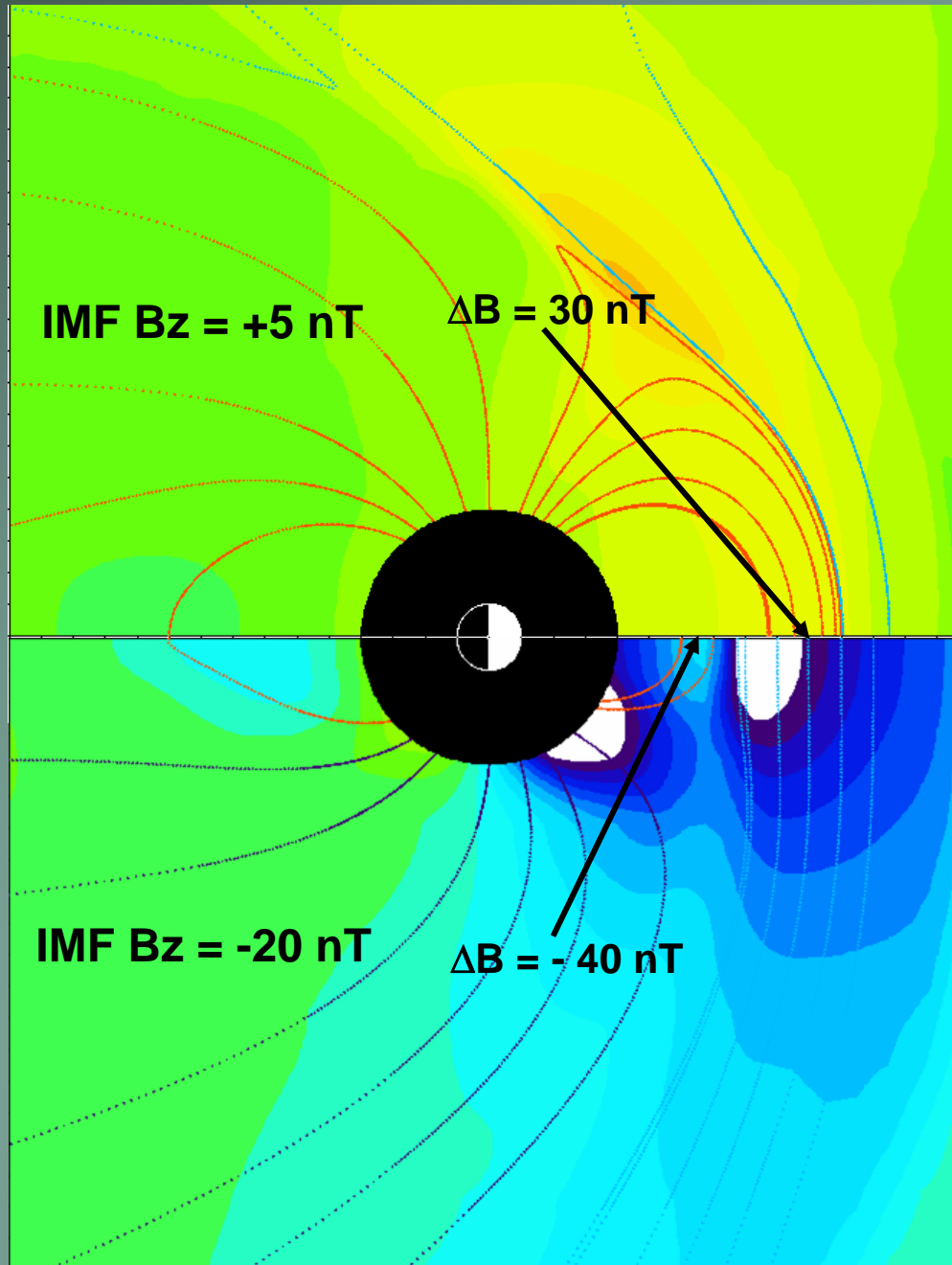
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Explanation



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Extreme Case



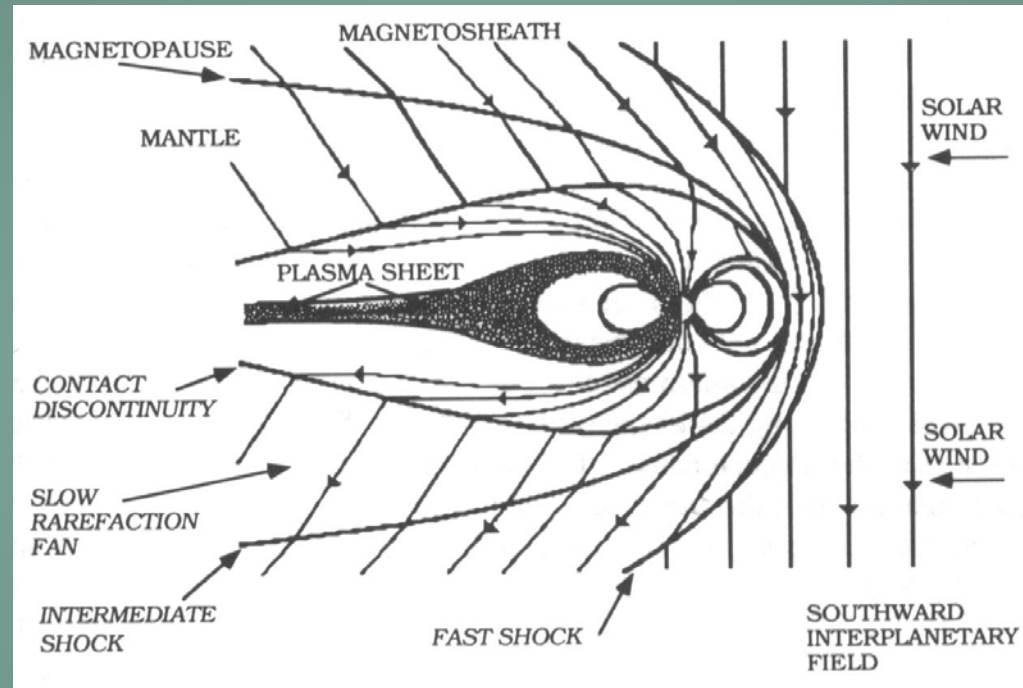
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Negative Dayside
Compression

The Plasma-Entry Problem

How does solar wind plasma get into the plasma sheet?



(Coroniti and Kennel, 1979)