

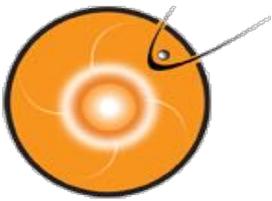
# International Forum for Space Weather Capabilities Assessment

*Focus: Radiation and Plasma Effects Working Team's progress*

Yihua Zheng, Natalia Ganjushkina, Timothy B Guild,  
Pier Jiggins, Insoo Jun, Joseph E. Mazur, Matthias  
Meier, Joseph I Minow, Dave Pitchford, Paul O'Brien,  
Yuri Shprits, W. Kent Tobiska, Michael A. Xapsos, M.  
M. Kuznetsova

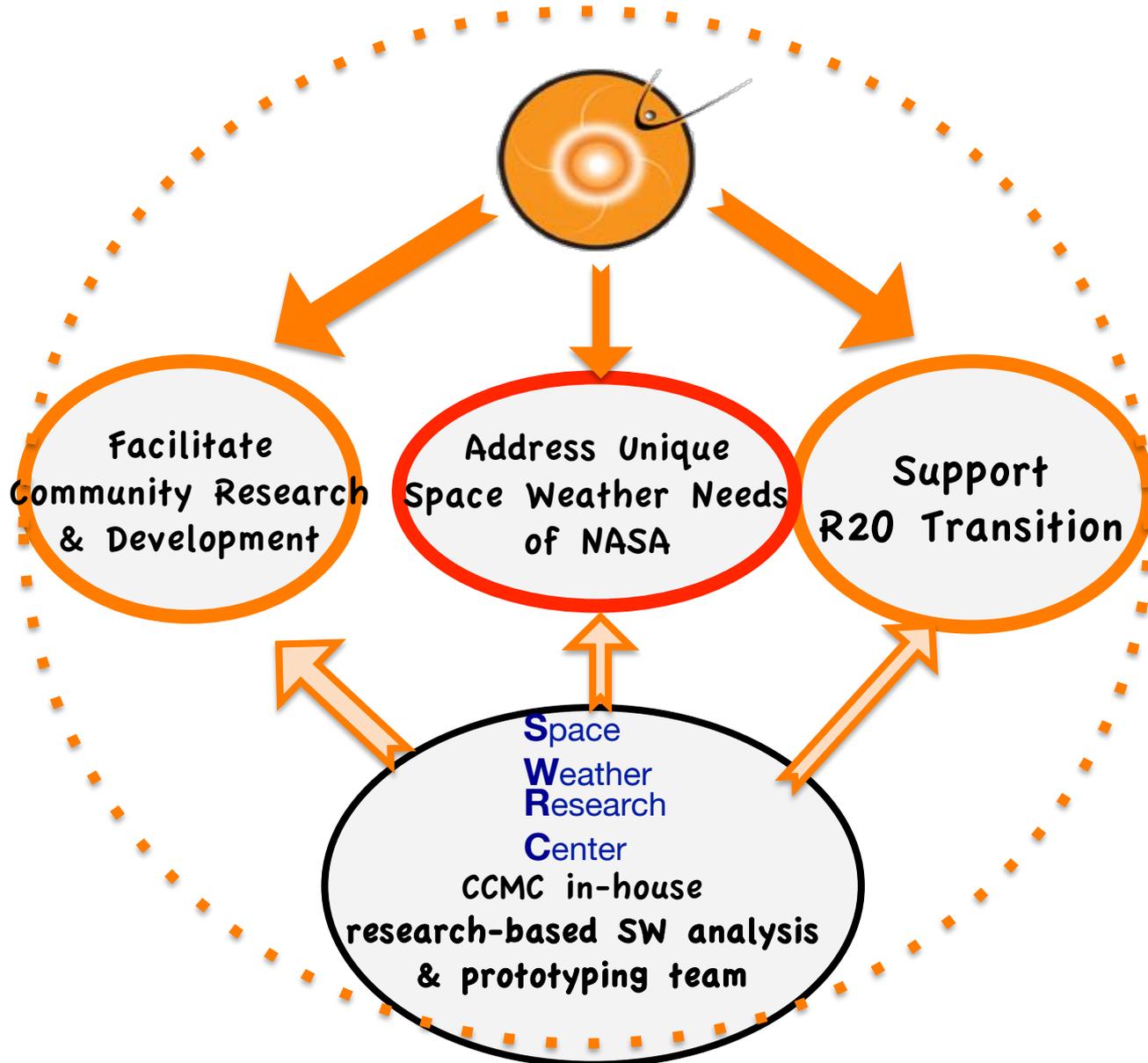
**Space Environment Engineering and  
Science Applications Workshop**

8 September 2017



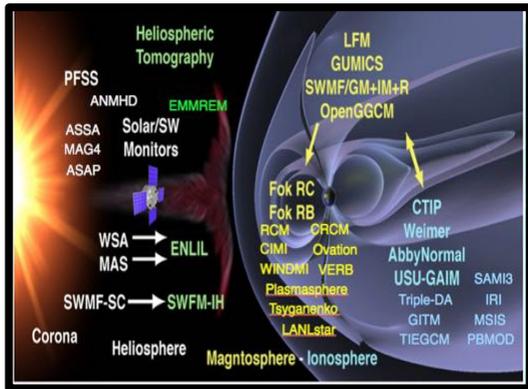
# CCMC

The Community Coordinated Center



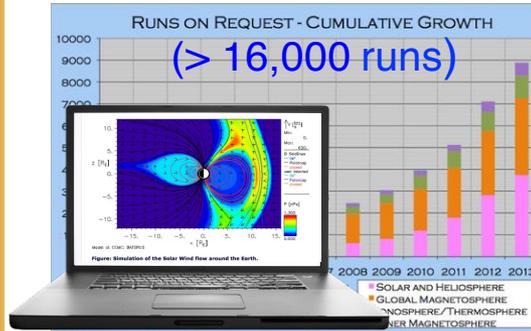
# CCMC Assets & Services

## Models

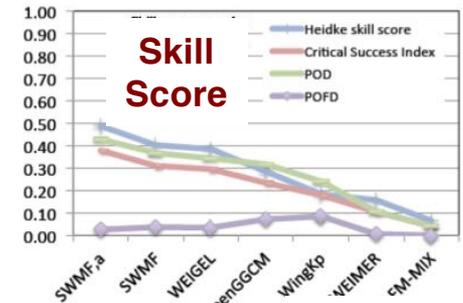


(expanding collection: > 80)

## Simulation Services



## Assessment



Scoreboard



**Dissemination**  
Flexible Infrastructure  
Tools, Databases  
Actionable Displays



**KAMELEON**

CONVERSION • ACCESS • INTERPOLATION

## Space Weather Services

for NASA's missions



**Space Weather  
Research Center**

## Hands-on Education



# International Forum on Space Weather Capabilities Assessment

The Forum brings together space environment experts, models developers, data providers, forecasters and end-users to

- define **internationally recognized metrics** that are meaningful and informative to end-users, developers, and decision makers;
- evaluate the **current state** of space environment models, applications and forecasting techniques;
- establish a procedure to **quantify and track progress** over time;
- facilitate **communications** between forecasters and researchers;
- **address challenges** in data-model comparison



<https://ccmc.gsfc.nasa.gov/assessment/>

## **SUPERTOPIC: QUANTIFYING SCIENTIFIC PROGRESS**

- Assessment of Understanding and Quantifying Progress Toward Science Understanding and Operational Readiness

### **SOLAR**

- Solar Flare Prediction Coronal & Solar Wind Structure
- 3D CME kinematics and topology
- Solar Indices and Irradiance

### **GEOSPACE: Geomagnetic Environment**

- Ground Magnetic Perturbations: dBdt, delta-B, GICs, FACs
- Geomagnetic Indices
- Magnetopause location and geosync. orbit crossing

### **HELIOSPHERE**

- CME Arrival Time
- IMF Bz at L1
- SEPs

### **GEOSPACE: Auroral Region**

- Auroral precipitation and high latitude ionosphere electrodynamics

### **RADIATION and PLASMA EFFECTS**

- Surface Charging
- Internal Charging
- Single Event Effects
- Total Ionizing Dose
- Radiation effects for aviation

### **IONOSPHERE**

- Neutral Density and Orbit Determination at LEO
- Global & Regional TEC
- Ionosphere Plasma Density: NmF2/foF2, hmF2, TEC
- Ionosphere Scintillation

### **INFORMATION ARCHITECTURE**

- Information Architecture for Interactive Archives (IAIA)

# The Radiation & Plasma Effects Working Team

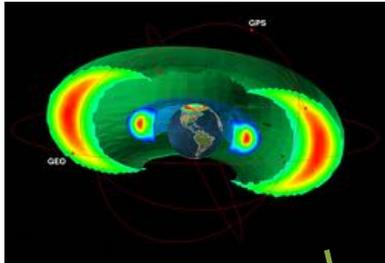
- **Radiation and Plasma Effects Working Team deals with five different subtopics with a variety of plasma & particle populations**
- **It bridges the space environments, engineering and user community.** Close collaboration among them becomes even more paramount. **Choosing proper metrics** that are simple yet meaningful to different groups of people, and measurable over a long period of time, is a challenge.
- **International Forum for Space Weather Capabilities Assessment** intends to be a long-term, community wide effort. Focus of this presentation: what has been achieved at the first workshop (April 3-7, 2017, Cape Canaveral, FL) arising from the forum.

**Electron radiation storms**

**Ion radiation storms**

**Electrons < 100s keV**

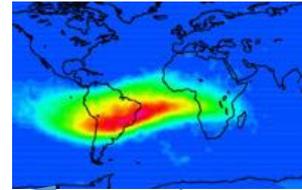
Radiation Belts



SEPs

GCRs

SAA

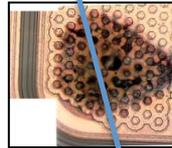


Ring current, aurora, plasma sheet

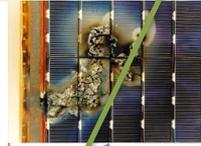
Internal electrostatic discharge



Single event effects



Surface Charging



**Radiation Impacts on Aviation**

**Spacecraft**

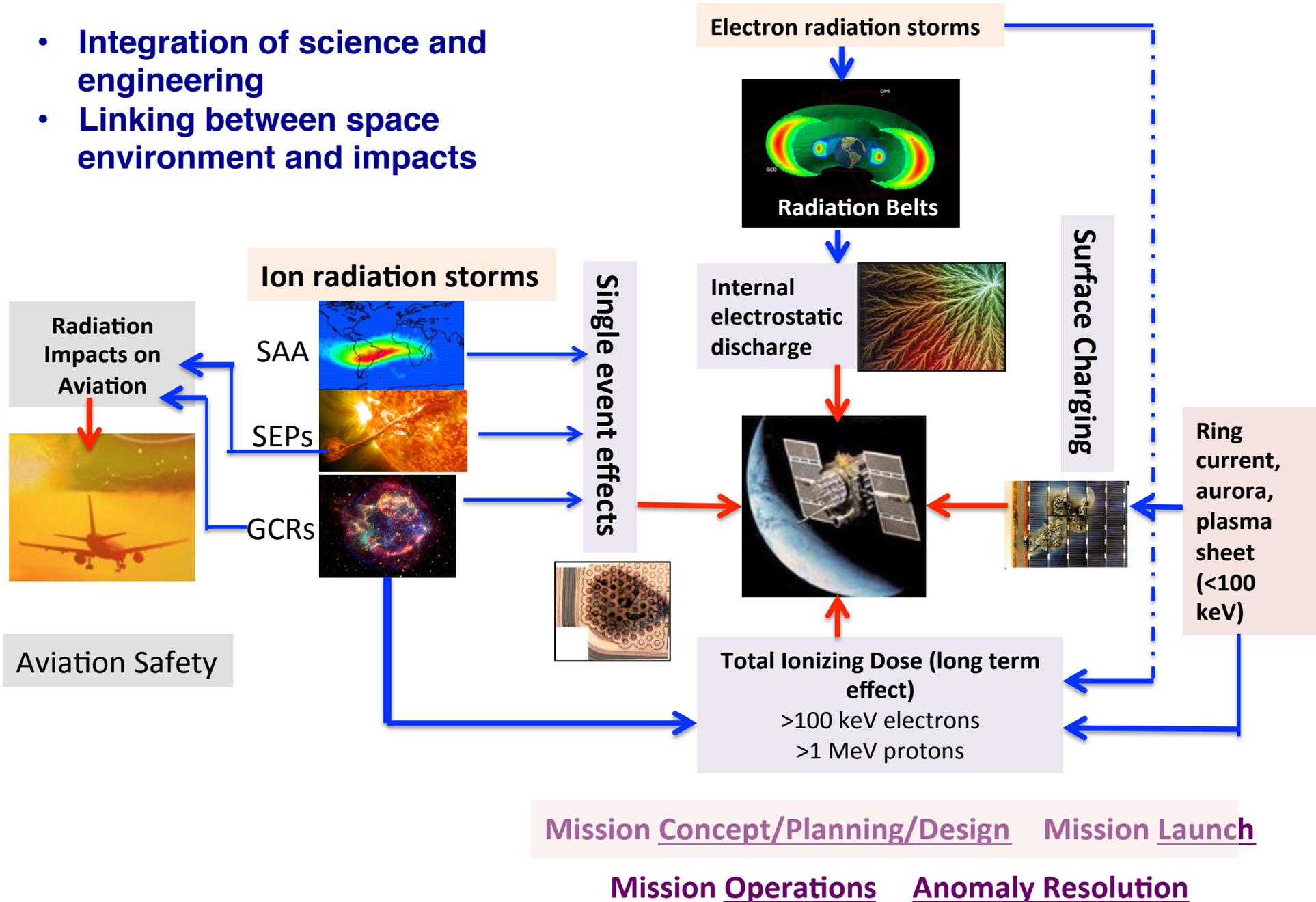
Total Ionizing Dose (long term effect)  
>100 keV electrons  
>1 MeV protons

Mission Concept/Planning/Design Mission Launch

Mission Operations Anomaly Resolution

Aviation Safety

- Integration of science and engineering
- Linking between space environment and impacts



# Radiation and Plasma Effects Working Group co-leads

Decided to focus on **space weather models**

Space weather effects include:

- Surface charging (J. Minow, N. Ganushkina, D. Pitchford)
- Internal Charging (P. O'Brien, Y. Shprits)
- SEEs (M. Xapsos, P. Jiggins, J. Mazur)
- Radiation at aviation altitudes (K. Tobiska, M. Meier)
- Total Dose in solar array and electronics due to SPEs and electron enhancements (I. Jun, M. Xapsos, T. Guild)

# Summary of Metrics

1. Fix Orbits: polar LEO, MEO, GEO, GTO, polar aircraft route
2. Derive effects metrics based on standard orbits/components.

	Effect Metric	Science Predictands	Time Period (Space Weather)
<b>Surface charging</b>	12 keV e- flux	12 keV e- flux; Te; Ne	seconds
<b>Internal charging</b>	>100 fA/cm <sup>2</sup> [100 mils]	1 MeV and >2 MeV e- flux	24-hour averaged
<b>SEEs</b>	SEE rate [100 mils]	>30 MeV p+ flux; >15 MeV.cm <sup>2</sup> .mg <sup>-1</sup> LET flux	5-min, daily, weekly
<b>TID</b>	Dose in Si [100 mils; 4 mils]	30-50 MeV p+ flux; >1.5 MeV e- flux 1-10 MeV p+	Daily, weekly
<b>Aviation</b>	Dose rate in aircraft (D-index)	2 spectral parameters (power law with rigidity)	5-min, Hourly

More can be added

# Stoplight Metrics

*Paul O'Brien*

- Satellite operators tend to prefer RED/YELLOW/GREEN stoplight indicators to real-valued quantities
- Statistical approach (over a long time interval)
- GREEN is the 75<sup>th</sup> percentile and below, RED is the 97<sup>th</sup> percentile and above.
- This fixes the proportion of GREEN/YELLOW/RED in time: 75% GREEN, 22% YELLOW, 3% RED
- If anomalies are rare, then the false YELLOW rate is 22% and the false RED rate is 3%
- A good metric, then, is what percent of *anomalies* occur when the tool is outputting GREEN?
- A unifying metrics for surface charging, internal charging, single event effects, total dose

# Surface Charging Status

- User groups include spacecraft designers, operational situational awareness, anomaly investigations, and impact on science measurements
- Metrics (team is evaluating options):
  - Statistical evaluation using O'Brien “green anomalies” technique
  - Parameters used for inputs to charging models
    - GEO, MEO, GTO: Ne, Te, Ni, Ti or other
    - LEO polar (auroral): Ne,  $E_{\text{beam}}$ ,  $\Delta E_{\text{beam}}$ , and other Fontheim parameters
    - Flux spectra at different locations
- Environment models (initial focus):
  - Ovation – CCMC implementation
  - LANL model (Vania Jordanova)
  - IMPTAM (Natalia Ganjushkina), run online in near-real time since 2013
  - CIMI (Natalia Buzulukova)
- Spacecraft charging models (secondary effort, but compare with  $\Phi_{s/c}$ )
  - NASCAP
  - SPIS
  - SPENVIS, MUSCAT, and other small group charging codes

# Internal Charging Summary

## Internal charging headline metrics:

- User Metrics: % Green anomalies for 24-hour average current beneath 100 mils Al spherical: GEO, GTO
- Science Quantity (stat TBD): Omnidirectional differential or integral flux: GEO, GTO

## Internal charging events/intervals

- 2015 has some nice big storms, RBSP data to validate
- The March, April, June, and July 2015 storms (**TBR**: need CME and CIR/HSS storms)

## Internal charging "comprehensive" metrics:

- User Metrics Add: 6-, 72-hour averages; 40, 350 mils; LEO, HEO, GNSS
- Science Quantities Add: locally mirroring flux; 0.1-1 MeV; LEO, HEO, GNSS

## Models (**Bold** indicates high probability of running benchmarks soon):

- **VERB, RBE/CRCM/CIMI, DREAM**, BAS, Rice REM, Salamambo
- CRRESELE in Ap mode
- **GREEP, SWPC REFM**, Ukhorskiy nearest neighbors, NARMAX

# Internal Charging Areas of Concern

- We are not currently addressing how the metrics account for model error: is it really a “green” anomaly if the model error bar included some yellow?
- We are not addressing mission design specs (Satellite design users, govt agency, insurers): out of scope, and hard to validate a 95% confidence value for 10-year worst case without 200 years of data.
- How do we address designer, insurer, govt agency needs? By including most severe, well-observed events in our validation set.
- We are concerned about the comparability of models with different inputs (observed initial/boundary conditions, versus initial/boundary conditions provided by a coupled model)

# Total Ionizing Dose

- Total dose is a climatological quantity, not space weather quantity
  - For mission duration
  - In order of days or years
- Total dose estimate for a mission uses a long-term average environment, not the worst case environment
- Quantities that are needed to compute total dose
  - Trapped electron and proton **fluence spectra** for a given mission duration
  - SEP mission fluence spectrum
- Empirical (climatological) models are typically used for total dose calculation for a mission
  - e.g., AP9/AE9 for trapped particles or JPL/ESP for solar protons

## (1) Identify user groups

- Satellite designer (SD) for both commercial and government
- Satellite operators and anomaly analysts (SOAA) for both commercial and government
- Scientists (SCI) for both academia and government

## (2) Identify metrics for each user group

- SD: Dose-depth for the mission
- SOAA: Dose-depth from launch to given time (there are some data available)
- SCI: proton and electron energy spectra
  - Electrons for  $> 100$  keV
  - Protons for  $> 1$  MeV

## (3) Identify empirical models for each metric

- Trapped: AE8/AP8; AE9/AP9/SPM; IGE2006/POLE (other older models are also available (e.g., CRESSELE, CRESSPRO, etc.))
- Solar: King (1972); JPL; ESP/PSYCHIC; SAPPHIRE

## (4) Identify physical model for each metric

- Trapped: SALAMMBO; DREAM
- Solar: SOLPENCO

# Future Need (TID)

- Climatological models
- Flight data to continuously improve and update the existing empirical models
  - Flux energy spectra
  - Dosimeter data

*Insoo Jun, Mike Xapsos*

# SEEs: Summary (1/2)

- Trapped protons
  - AP9 (also AP8 still used in some standards);
  - PSB97 + update (local model based on SAMPEX/PET)
- SEPs
  - ESP-PSYCHIC
  - JPL
  - MSU
  - SAPPHIRE
- GCRs
  - ISO-15390 GCR model
  - Badhwar-O'Neill (BON)
  - DLR GCR model
- Magnetospheric Modelling codes (rigidity cutoff calculation):
  - ESHIEM-MSM (magnetospheric shielding code)
  - Shea and Smart model

# SEEs: Summary (2/2)

## Relevant parameters

- a) SD+SLAO (SEE rate): proton fluxes (>30 MeV & > 50 MeV) worst-case SEP values; worst-case solar particle event (SPE) fluence
- b) SD (SEL/SEB probability): proton fluences (>30 MeV & > 50 MeV) [Orbit-averaged radiation belt flux (fluence); cumulative SEP fluence]
- c) SD+SLAO+SO: Abundance ratios and charge states of SEP heavy ions ( $Z>2$ ) [extension to event-to-event variability/distributions if possible]
- d) SD+SLAO: LET behind nominal shielding ( $1 \text{ g.cm}^{-2}$ )  
*\*\*application of particle transport codes as black box only to derive useful quantities*

- SD: Satellite designer
- SLAO: Satellite/launcher/ aircraft operators
- SO: Standards organizations (ISO/ECSS/ NASA internal)

## Validation methods

- a). Statistical evaluation using O'Brien "green anomalies" technique
- b). Event /interval based

# Summary: Radiation Effects for Aviation

1. **Data**: compare absorbed dose rate in silicon or ambient dose rate equivalent, depending on instrument characteristics
2. **Models**: compare effective dose rate
3. RMS (Root-Mean-Square) metrics for error
4. Use **version numbers** for data and models
5. **Need more data** (spectral and TID) for model comparison/validation
6. Report time (UT), lat, lon, altitude, dose rate for ease of comparison
7. Support development of D-index for aviation community and discourage use of NOAA S-scale for aviation radiation

# Summary of Metrics

1. Fix Orbits: polar LEO, MEO, GEO, GTO, polar aircraft route
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<b>Aviation</b>	Dose rate in aircraft (D-index)	2 spectral parameters (power law with rigidity)	5-min, Hourly

# Outlook

- We have made headway at the April 2017 workshop.
- Will continue to make progress via telecons or other science meetings (GEM, CEDAR, SHINE, AGU....)
  - Will have a topical discussion at the November 2017 European Space Weather Week ‘[How to assess space environment models’ capability in satellite impact analysis](#)’
- **Everyone is invited to participate**