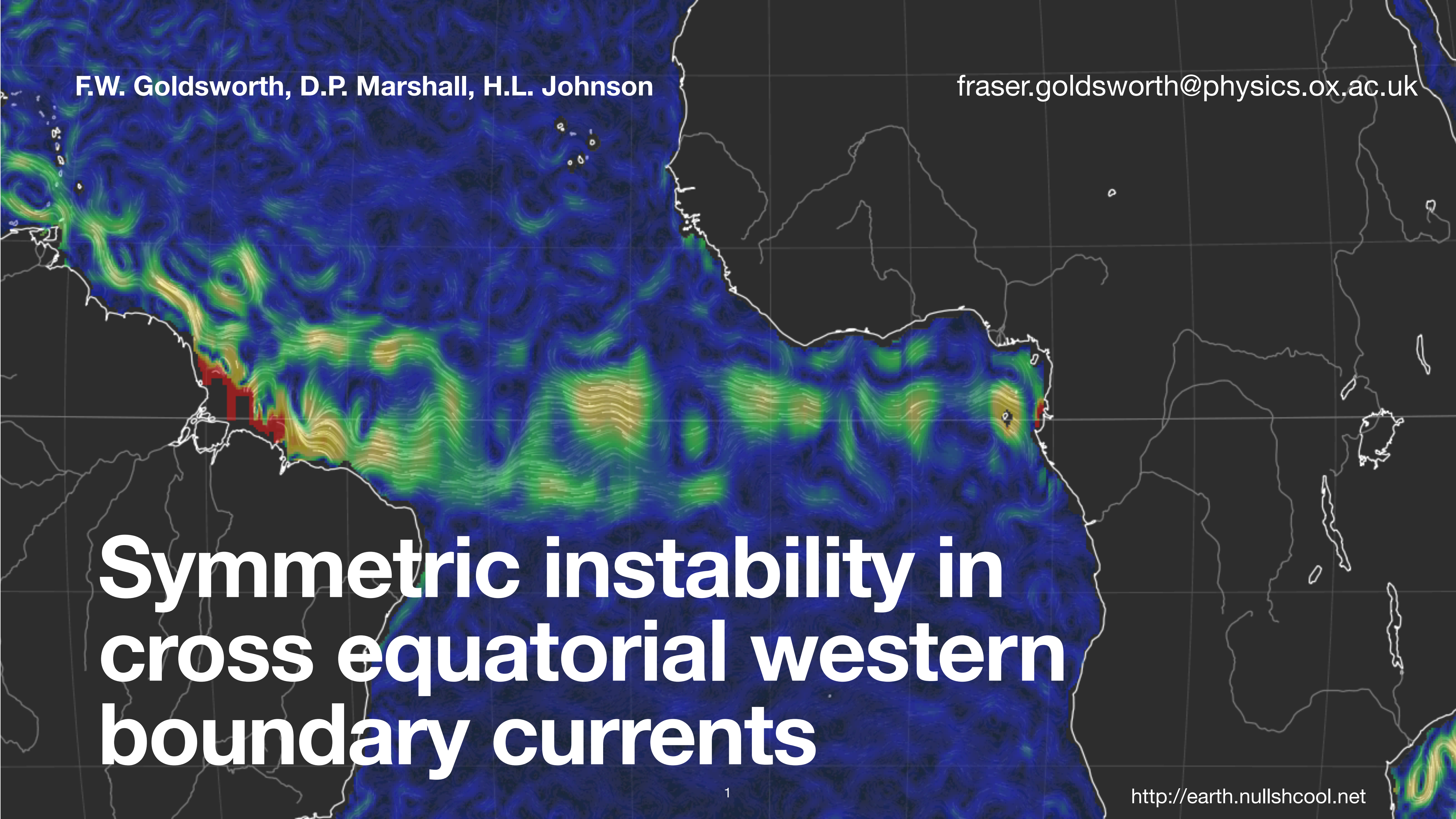


F.W. Goldsworth, D.P. Marshall, H.L. Johnson

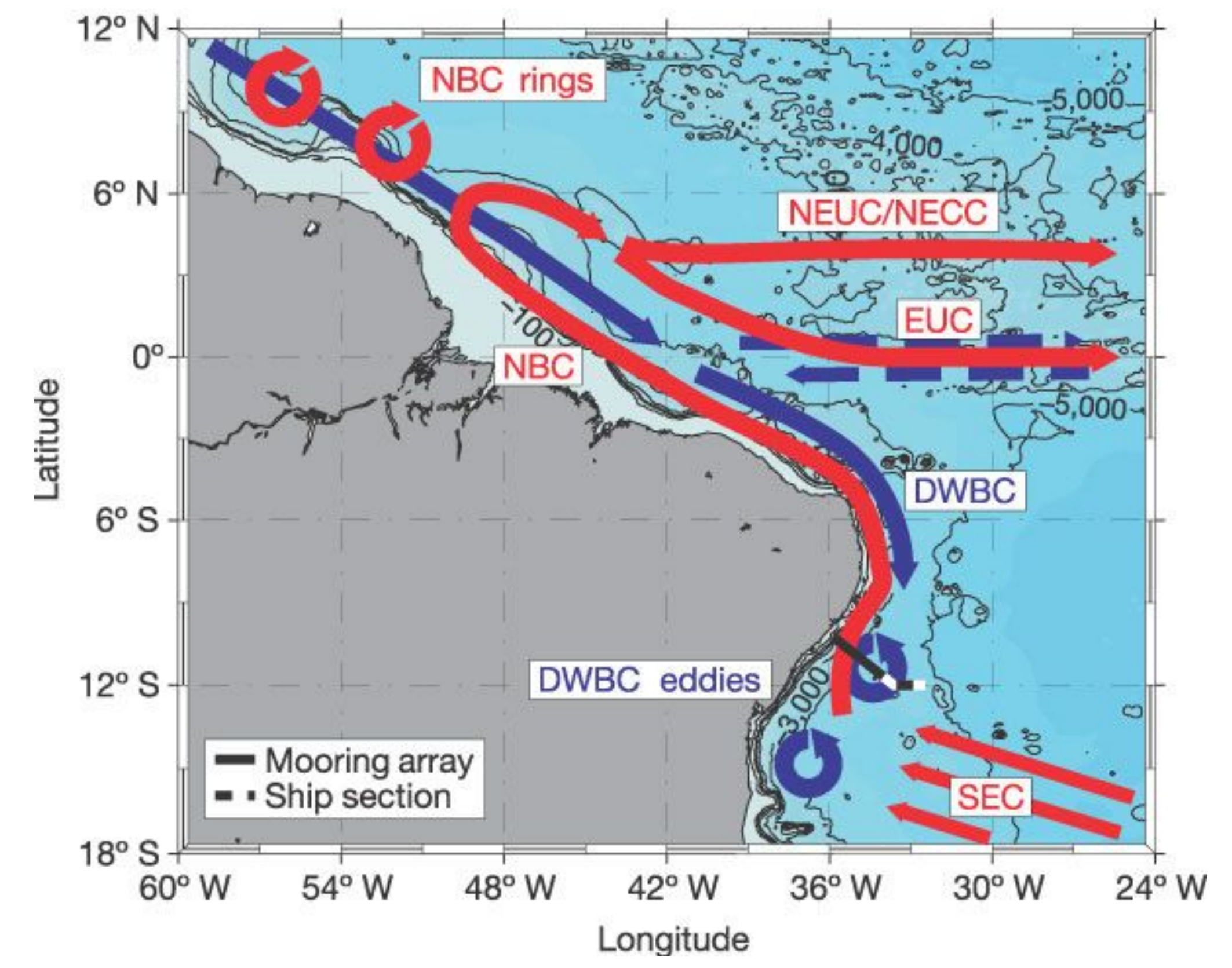
fraser.goldsworth@physics.ox.ac.uk

Symmetric instability in cross equatorial western boundary currents



Tropical circulations

- Northward flowing **surface currents**
- North Brazil Current & rings
- Equatorial Counter Current
- Southward flowing **deep western boundary current**



Dengler et al. (2004)

Potential vorticity

And a necessary condition for symmetric instability

- PV is materially conserved:

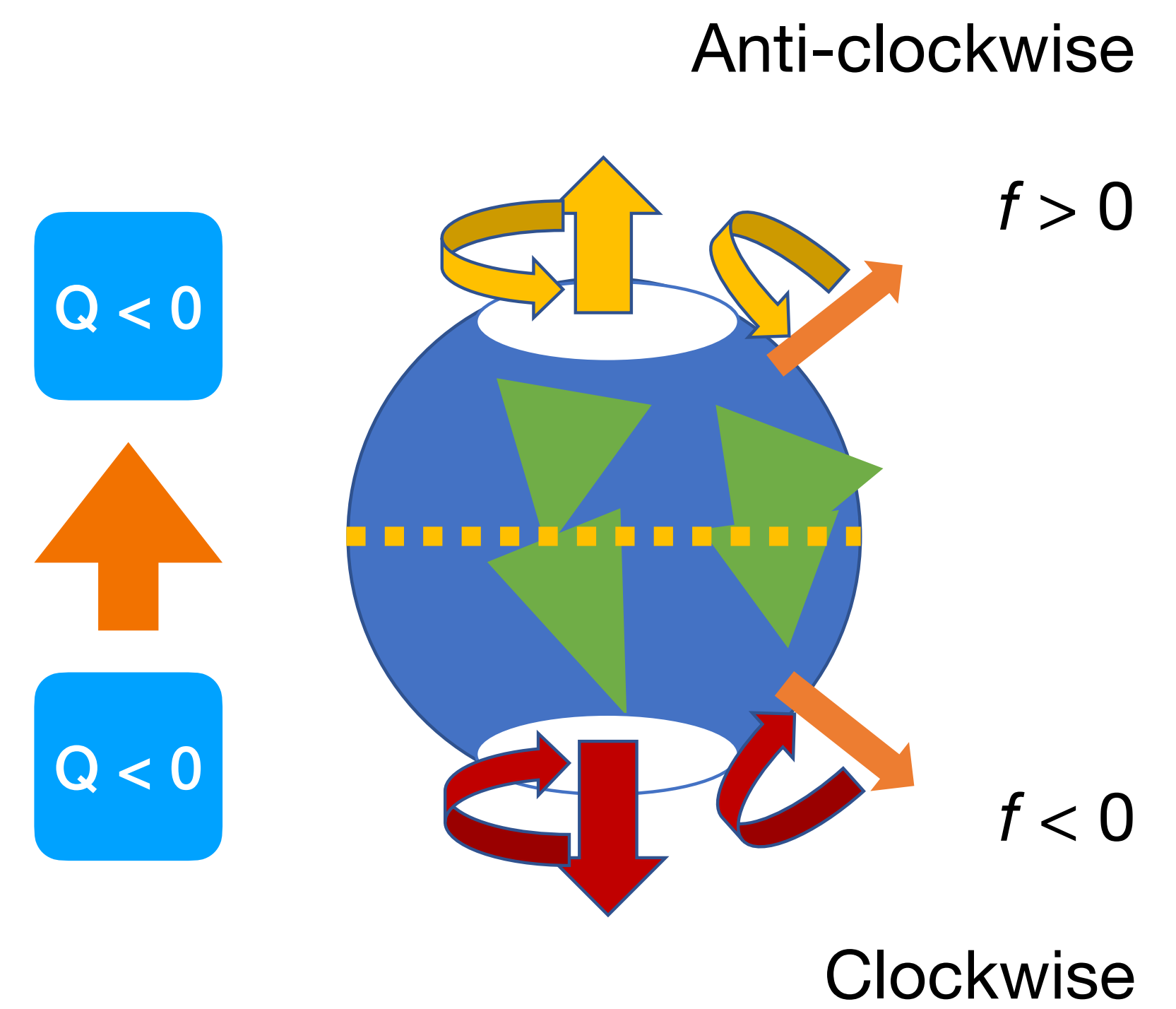
$$Q = (\mathbf{f} + \nabla \times \mathbf{u}) \cdot \nabla b$$

Planetary (red arrow pointing to \mathbf{f})

Relative (orange arrow pointing to $\nabla \times \mathbf{u}$)

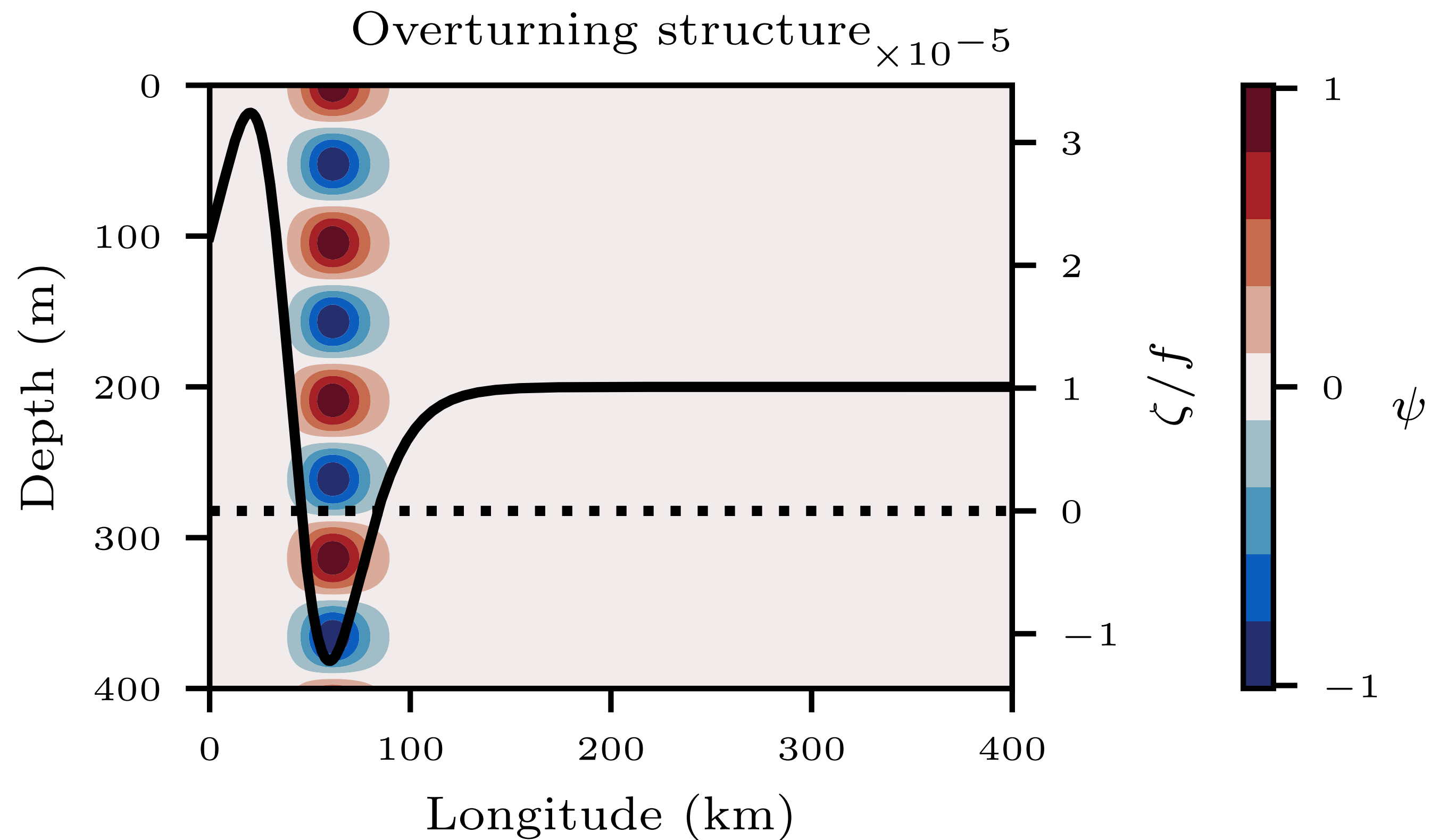
Stretching/baroclinicity (green arrow pointing to ∇b)

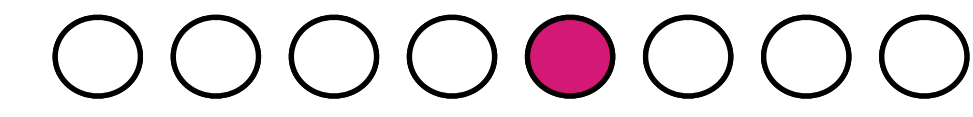
- SI if $fQ < 0$
 - If initially stable in SH, not stable in NH
- Growth rate $\sigma^2 \approx -f(f + dV/dx)$
 - Need large horizontal shear



What does symmetric instability look like?

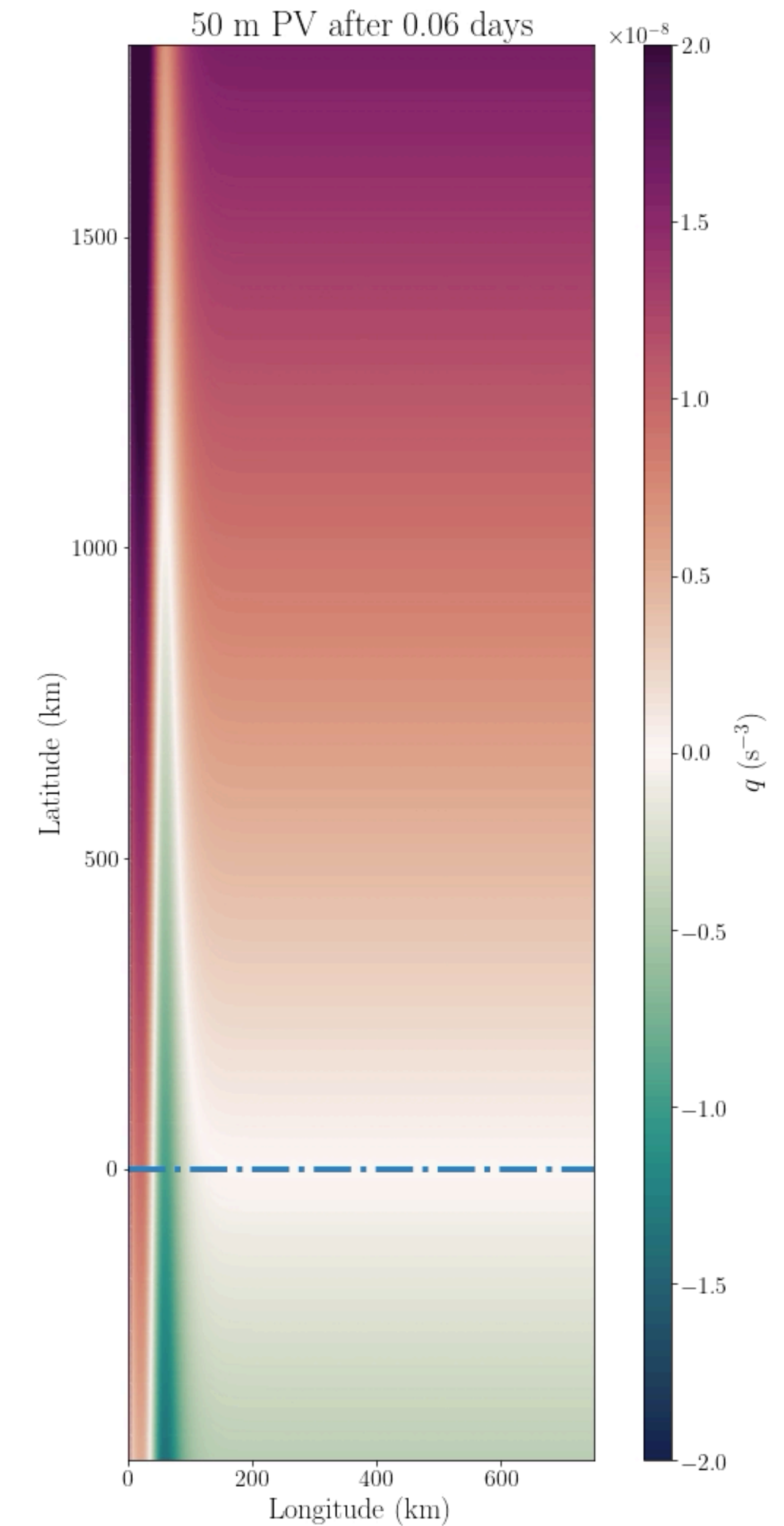
- Predictions from a linear stability analysis
- Stacked overturning cells
- Localised in regions of negative PV
- Could be an important mixing mechanism





An idealised model

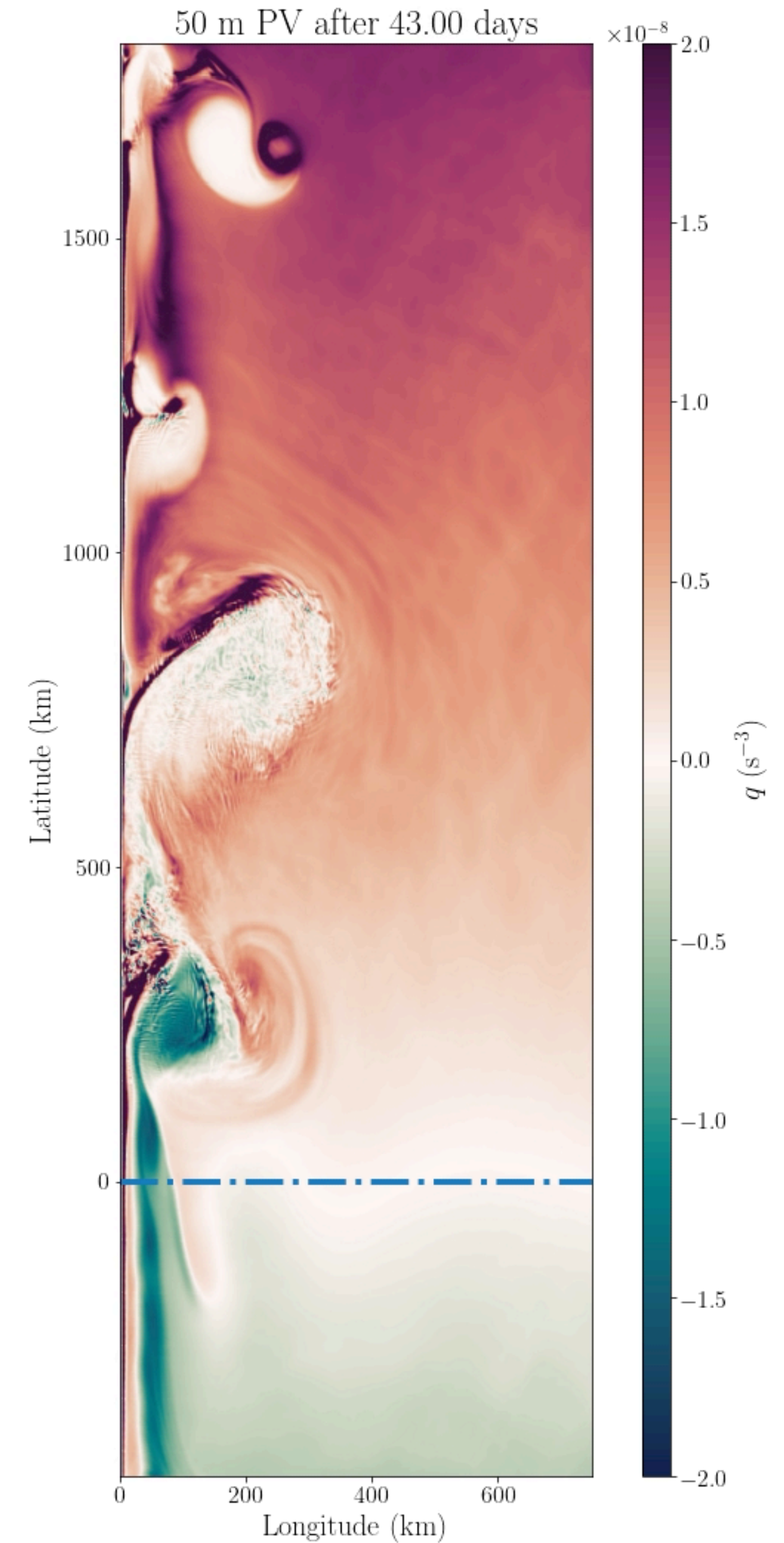
- Simplifications made:
 - Brazil is a straight line
 - No topography
 - Open boundaries (sponged)
- Unstable regions have **negative PV** in the northern hemisphere



An idealised model

Two types of instability

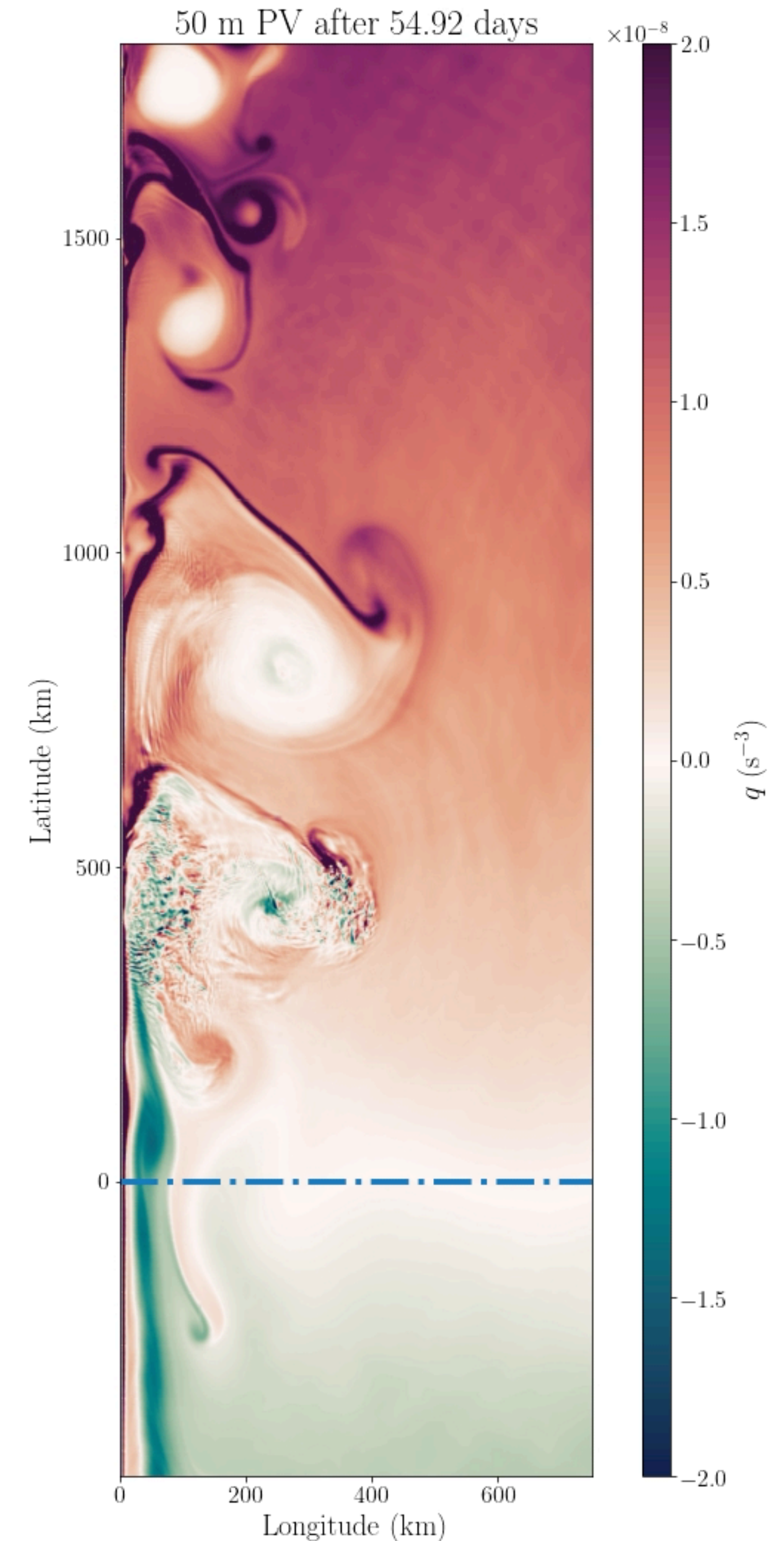
- PV at 50 m
- What's going on?
 1. **Eddy field** develops as fluid crosses the equator — e.g. Edwards & Pedlosky, 1998; Goes et al. 2009.



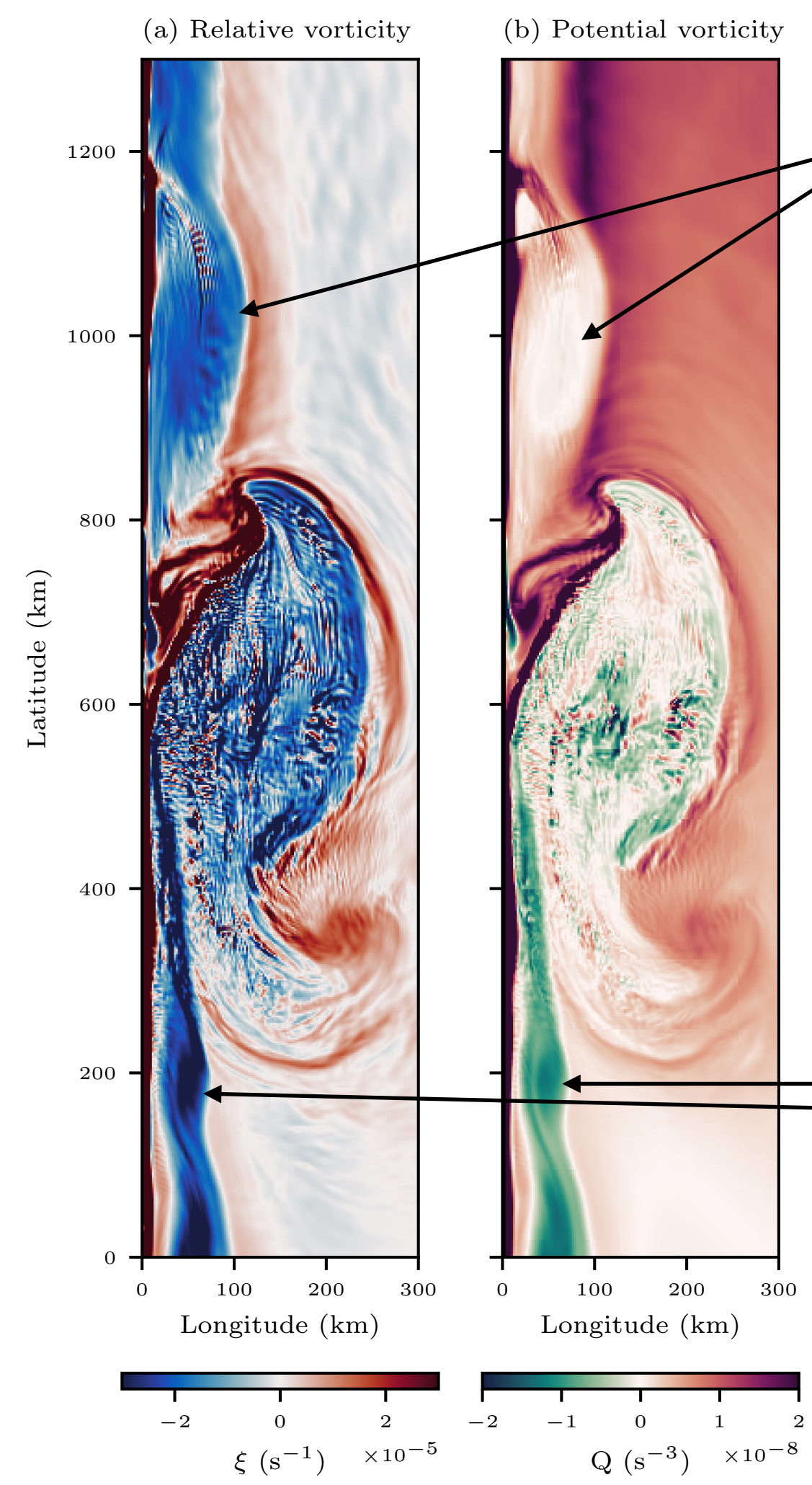
An idealised model

Two types of instability

- PV at 50 m
- What's going on?
 1. **Eddy field** develops as fluid crosses the equator — e.g. Edwards & Pedlosky, 1998; Goes et al. 2009.
 2. **SI is excited** from 300 km North of the equator.



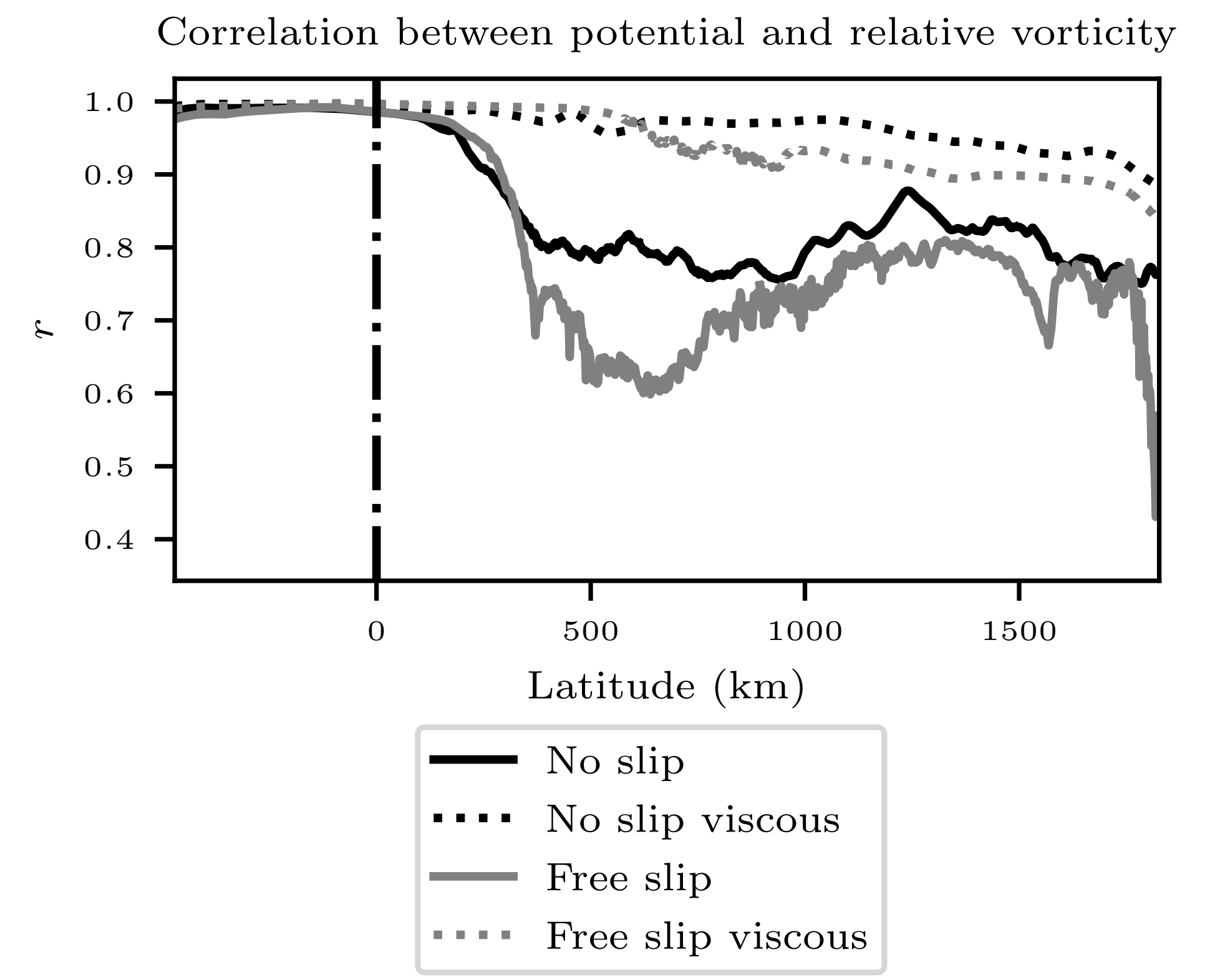
PV and relative vorticity

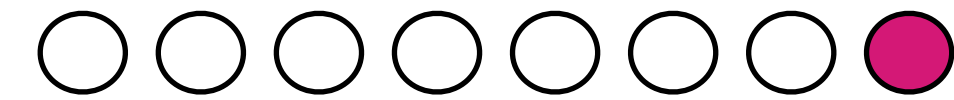


Sign of **RV** and **PV** uncorrelated \implies planetary vorticity dominates

Symmetric instability **alters** the PV balance

Sign of **RV** and **PV** match \implies relative vorticity dominates

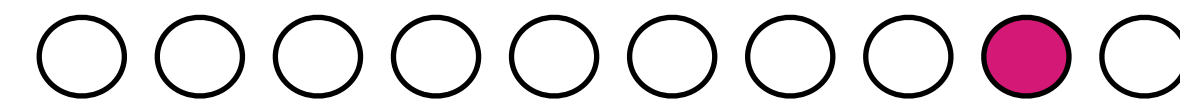




Conclusions

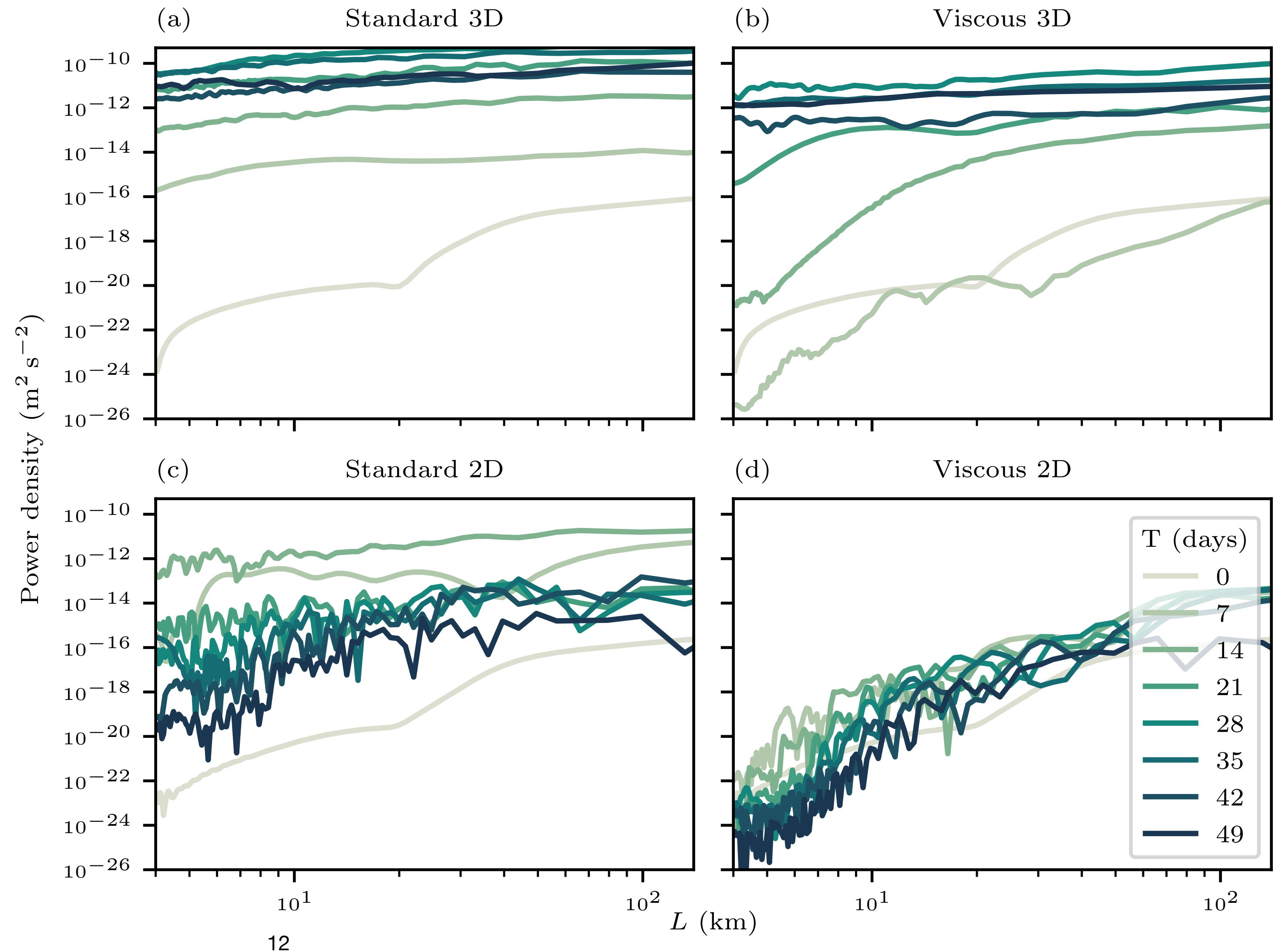
- From **theoretical considerations** we might expect to observe **symmetric instability** in **cross-equatorial** flows.
- **Symmetric instability** has been **observed** in an **idealised model** of cross-equatorial western boundary currents.
- We can see the effects of the instability on:
 - potential vorticity of **eddy cores**
 - **correlations** between potential and relative vorticity
- Next steps:
 - Deep western boundary currents
 - LLC4320 model
 - Existing glider datasets?
- For more details see Goldsworth et al. (2021), *Journal of Physical Oceanography* (early online release)

Supplementary slides



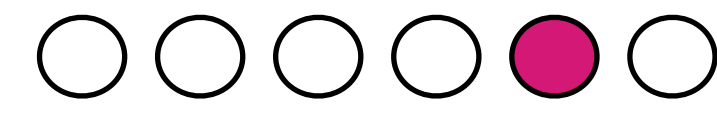
Instability & power spectra

- Spatially Fourier transform $w^2/2$ at 50 m
- Plot at week long intervals
- SI causes flattening and lifting of the spectra
- See also Yankovsky & Legg (2019)



Evidential summary

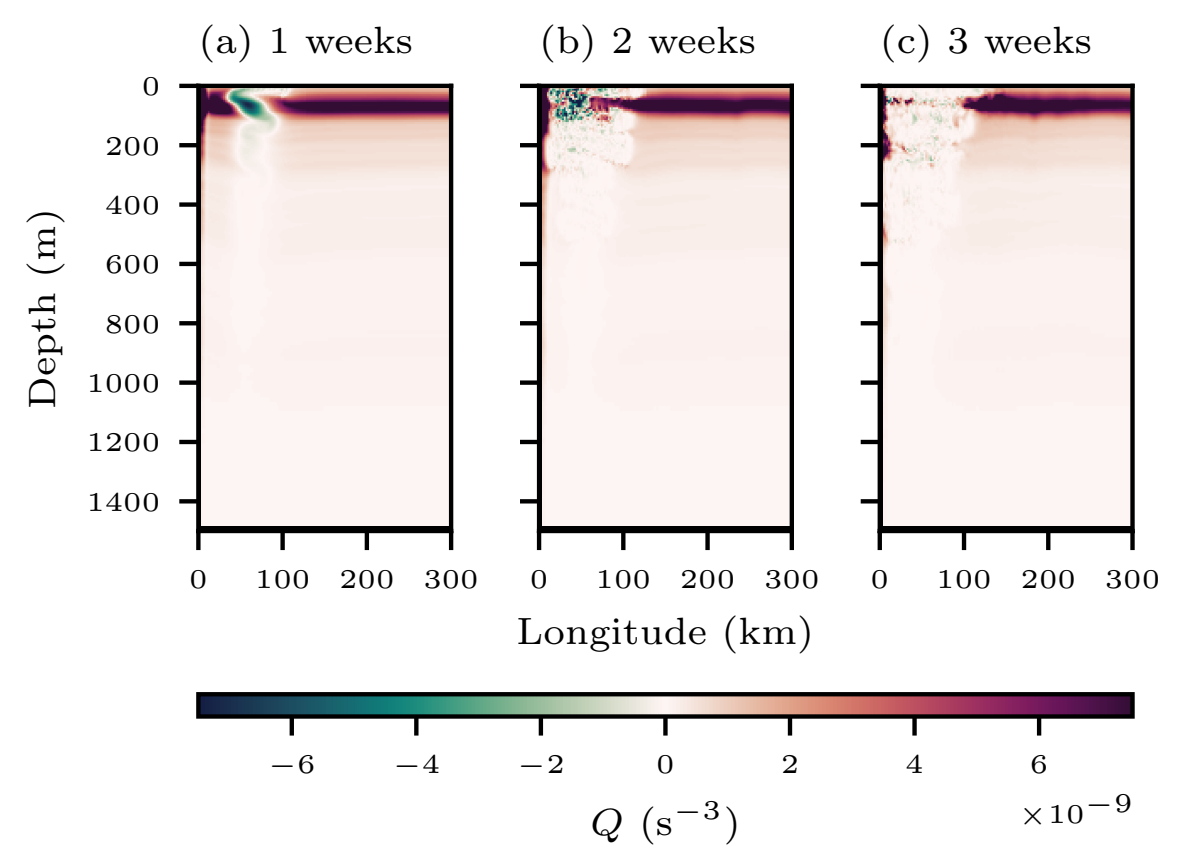
1. Regions of negative PV are unstable
2. Vertical scale
3. Horizontal scale
4. Time scale
5. Viscosity dependence
6. Structure of overturning
7. General agreement between linear stability analysis, simplified two dimensional models and 3 dimensional models
8. Power spectra of the vertical KE
9. Correlations between PV and relative vorticity



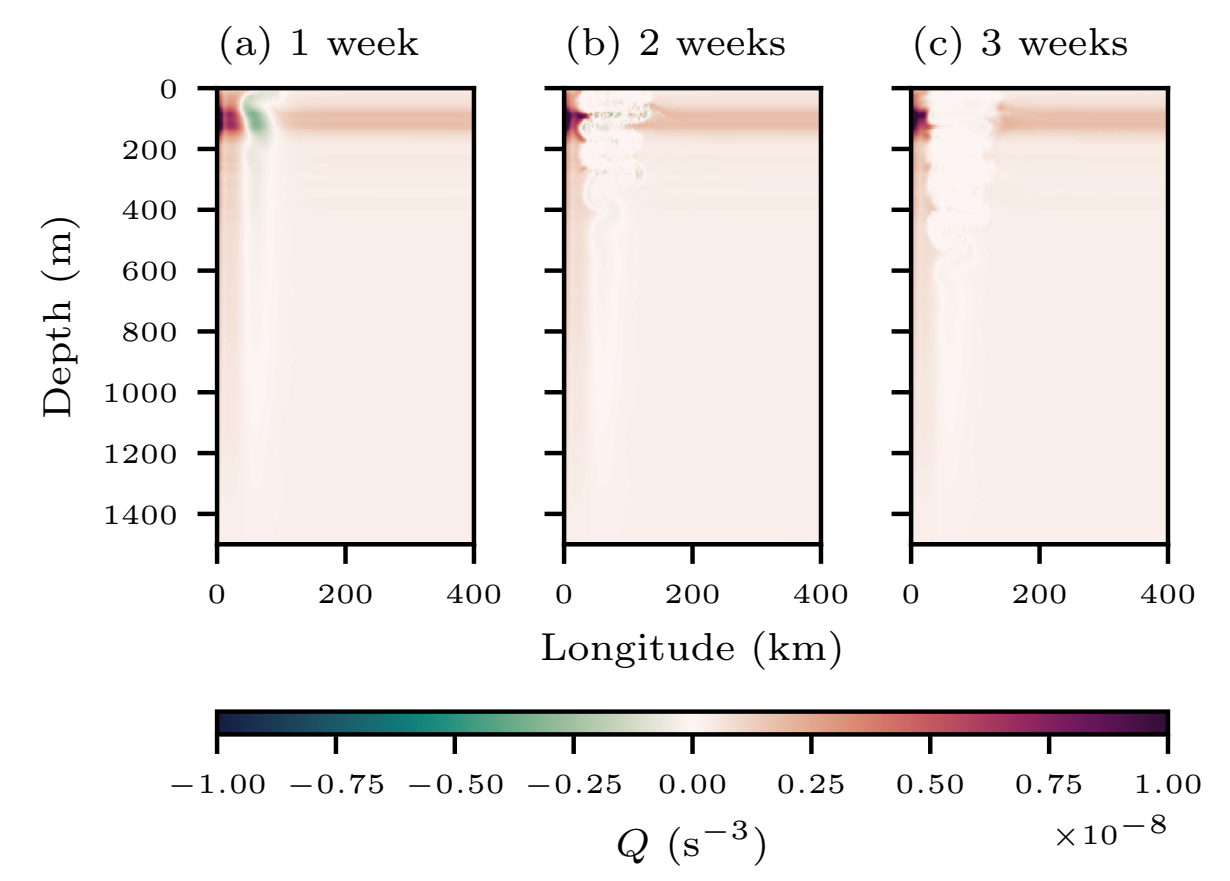
Why symmetric instability?

Predictions from a hierarchy of models

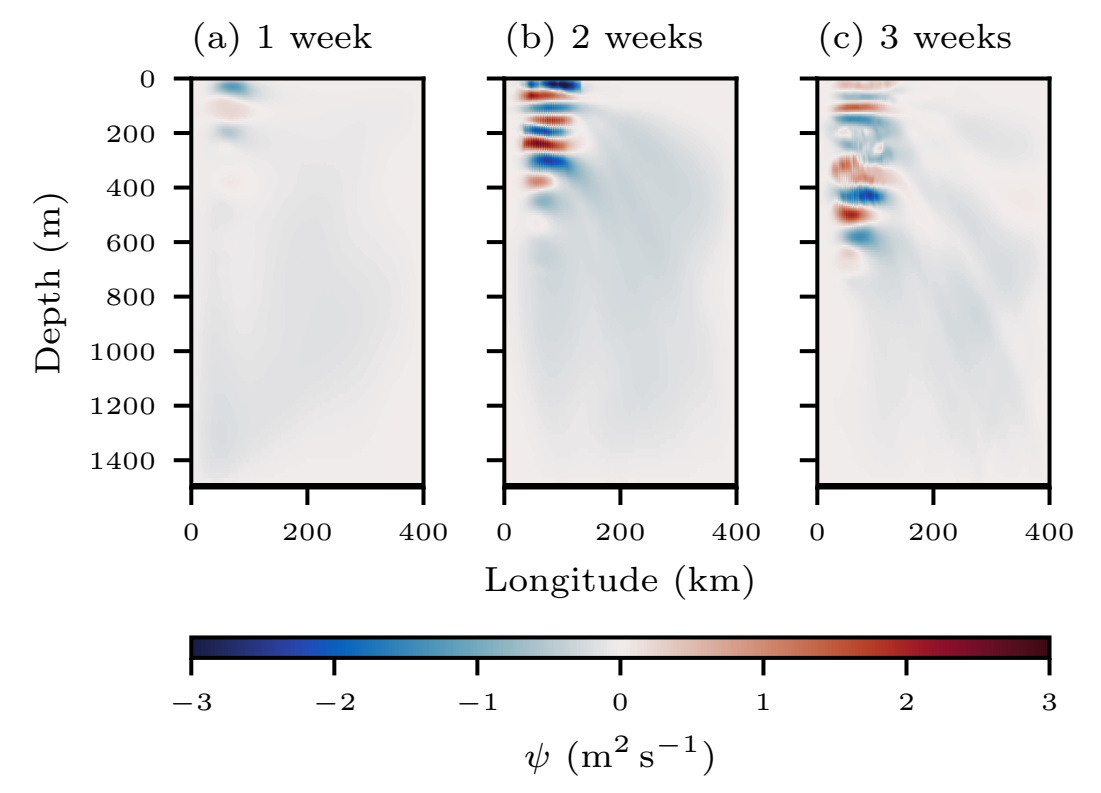
3D numeric PV



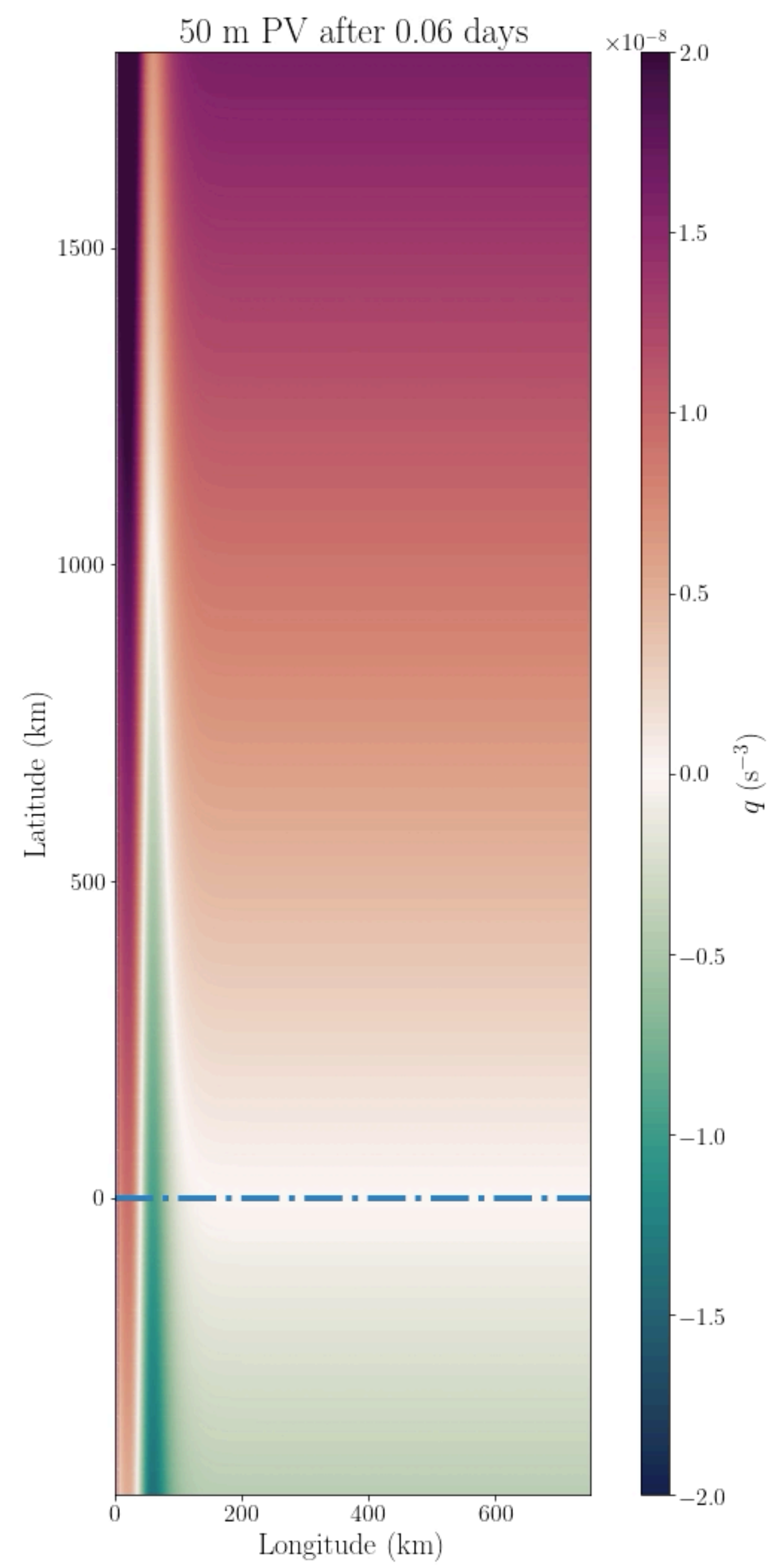
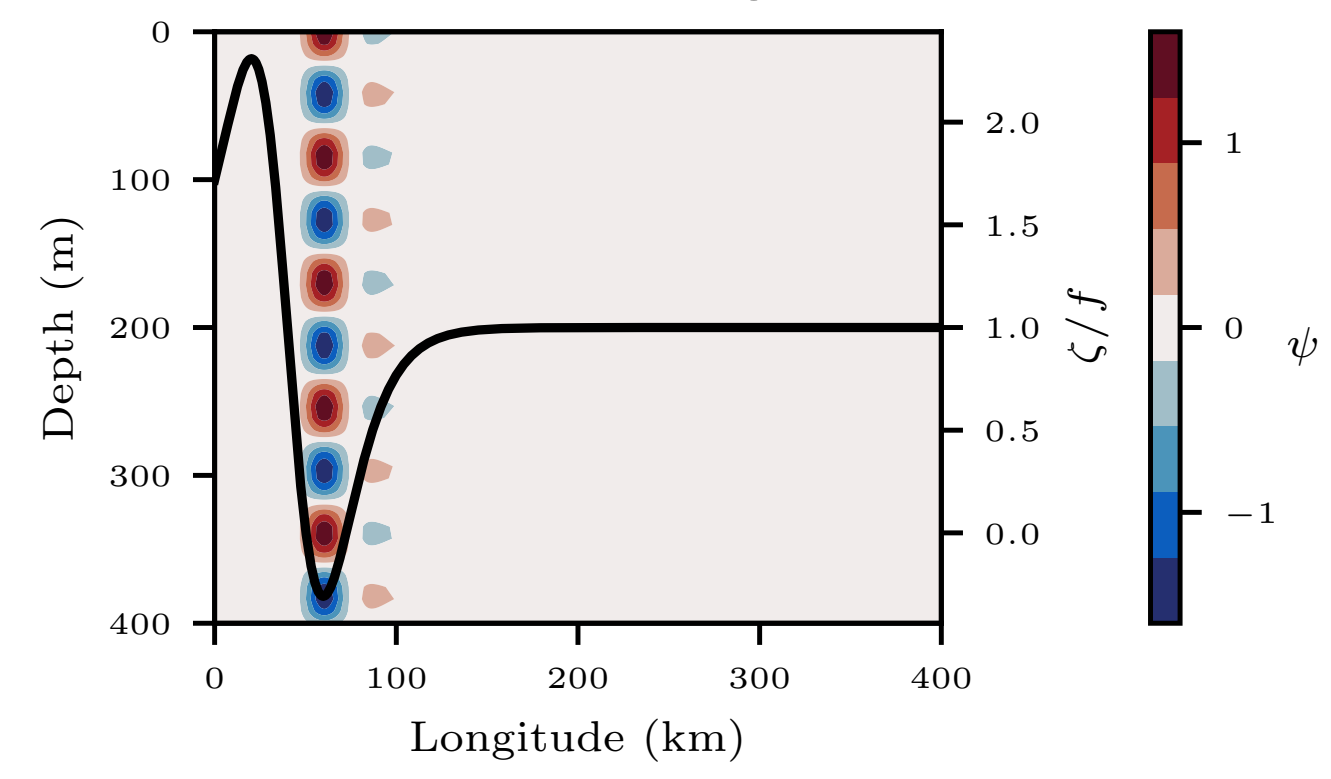
2D numeric PV



2D numeric overturning streamfunction



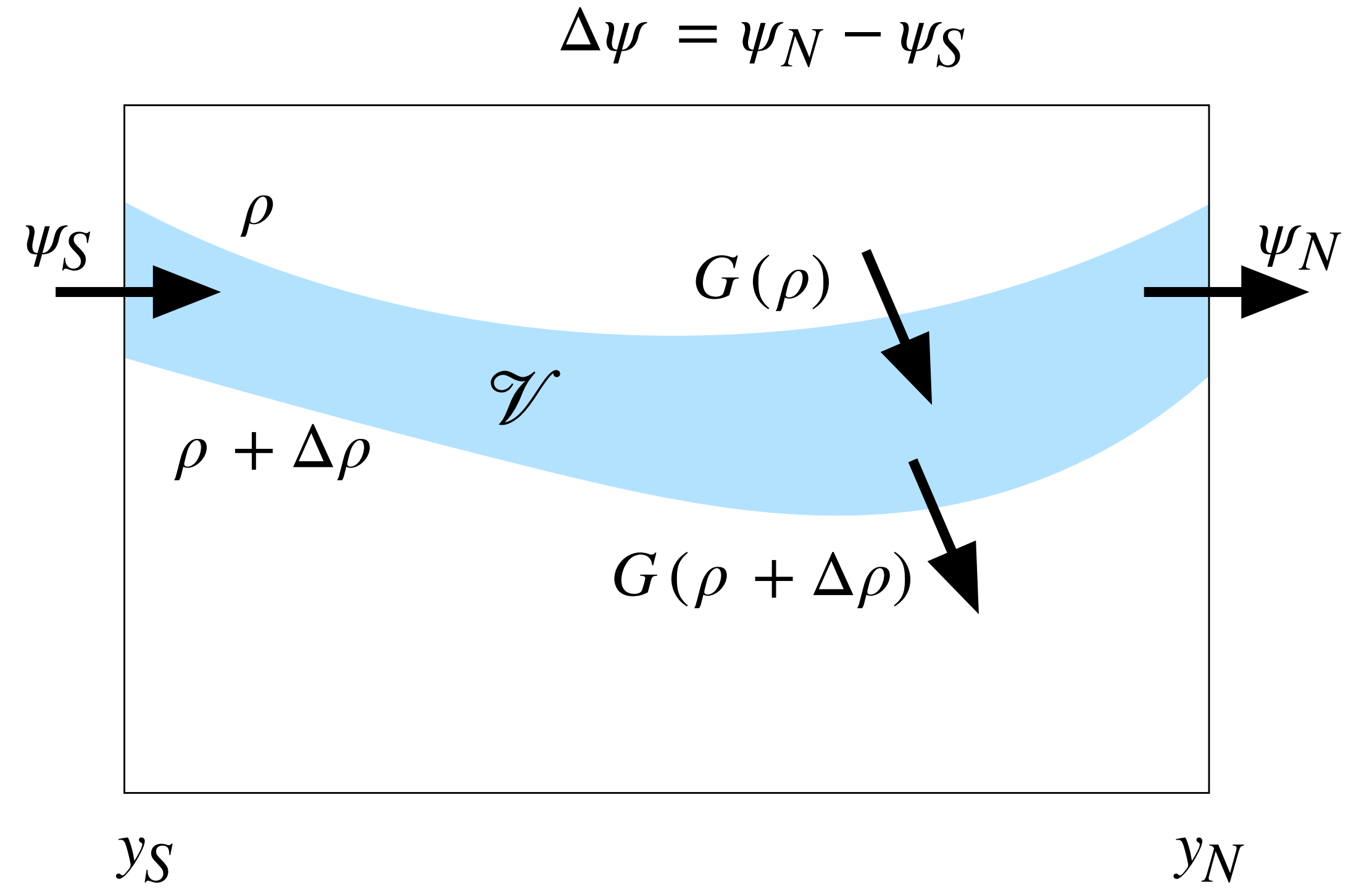
Theoretical overturning streamfunction



Water mass formation

The Wallin framework

$$F = \frac{\partial \mathcal{V}}{\partial t} + \Delta\psi = G(\rho) - G(\rho + \Delta\rho)$$



Water mass formation

The Wallin framework

