

Validation of COSMIC-2 Space Weather Science Products



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COSMIC-2 Space Weather (SpWx) Cal/Val Overview



- On the US side of the COSMIC-2 program, Cal/Val is a joint NOAA-USSF effort
 - NOAA was responsible for TGRS neutral atmospheric Cal/Val (now completed)
 - USSF is responsible for space weather Cal/Val (all sensors)
- SpWx Cal/Val is organized into 4 work areas
 - TGRS TEC
 - TGRS Scintillation
 - IVM in-situ ion density, drifts, composition and temperature
 - RF Beacon Scintillation & TEC



TGRS TEC Requirements

Per COSMIC-2 JL1PRD

<u>Parameter</u>	<u>Requirement</u>	<u>Status</u>
• Measurement uncertainty (relative)	0.3 TECu	Verified
• Measurement uncertainty (absolute)	3 TECu	Verified*
• Tangent altitude range (occultations)	60 km to s/c alt.	>90 km
• # of occultations & “overhead arcs”		
– <i>Limb TEC profiles</i>	>6,000/day	
– <i>Overhead TEC arcs</i>	>3,675/day	
– <i>Total profiles + arcs</i>	>12,000/day	~10,000/day
• Median data latency	30 minutes	~35 min

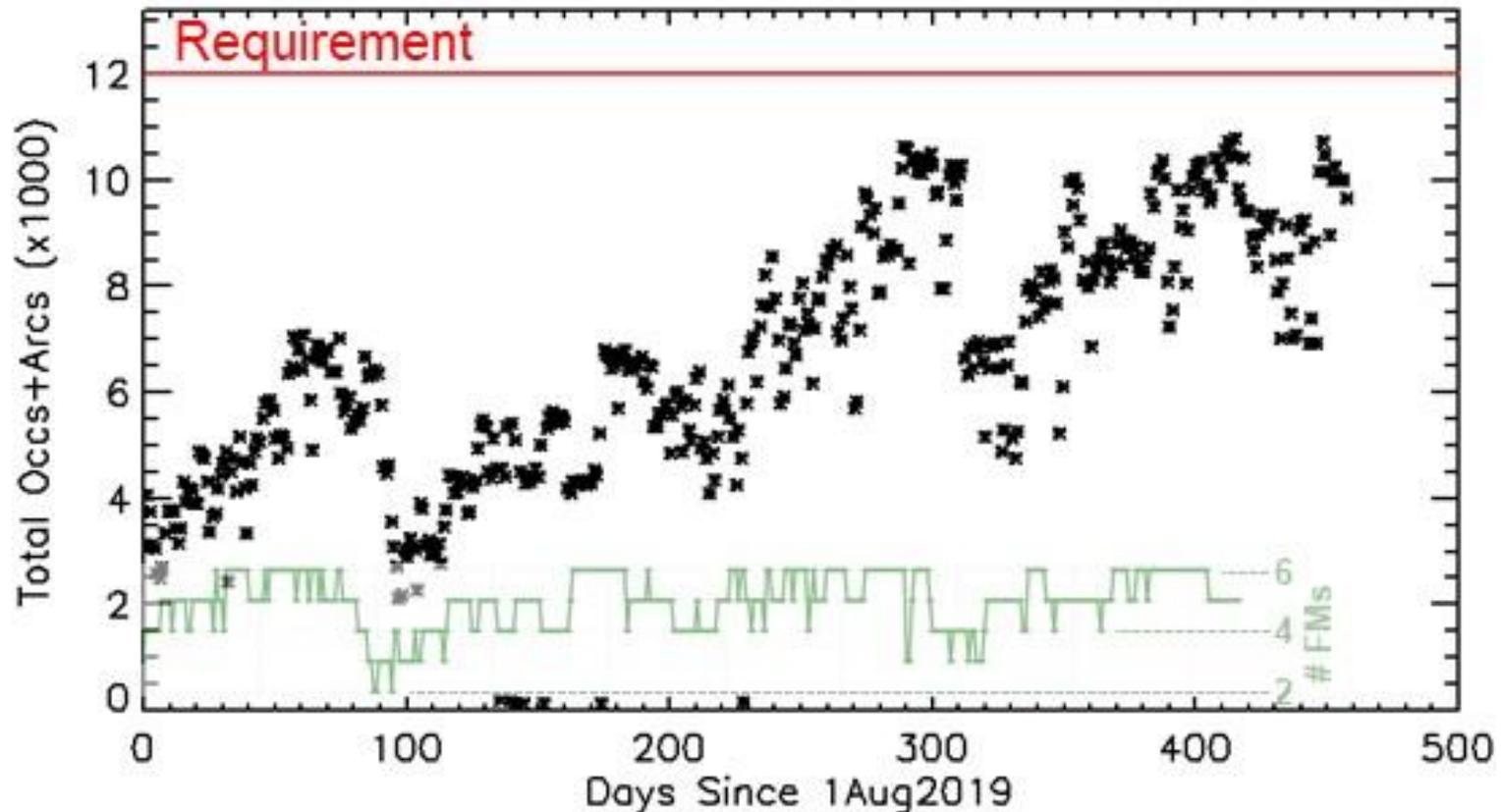
***See Pedatella et. al. for details**



Quantity of TGRS Observations is Increasing

Getting close to meeting Arc & Occultation count requirements

- Continued TGRS flight software refinements have improved performance
 - Better quality GLONASS data, more efficient use of sensor resources
 - Most occultations now span required altitude range



Further flight s/w refinements are planned to fully satisfy requirements

TGRS Ionospheric Scintillation Requirements



COSMIC-2 JL1PRD Requirements

<u>Parameter*</u>	<u>Requirement</u>
• S4 Measurement Range	0 to 1.5
• S4 Measurement Uncertainty	0.1
• σ_ϕ Measurement Range	0-20 radians
• σ_ϕ Measurement Uncertainty	0.1 radians
• On-board S4 calculations	All tracks (every 10s)
• High rate (HR) SNR/phase sent to ground up to 10% of TGRS telemetry volume	
– <i>TGRS sends down high rate (50 Hz) amplitude & phase profiles when the calculated on-board S4 exceeds a programmable threshold</i>	

*All scintillation parameters measured at GNSS L1 & L2

Cal/Val Objectives to Optimize Operational Utility

- Geolocate irregularity regions to 100-500 km accuracy
- Develop/assess an RO-based “All Clear” product
- “Limb-to-Disk”: Conversion of RO limb S4 to ground-to-space geometry
- Develop & evaluate “Bubble Maps” (joint TGRS/IVM product)

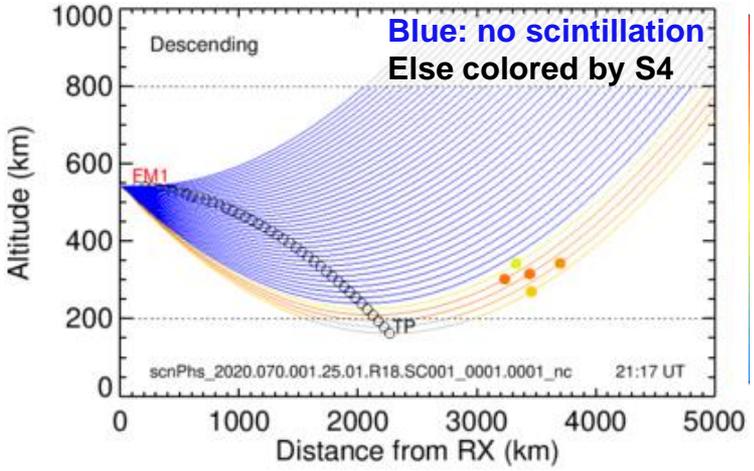
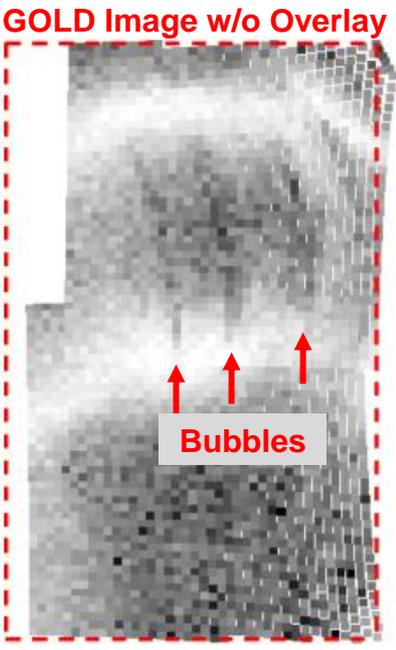
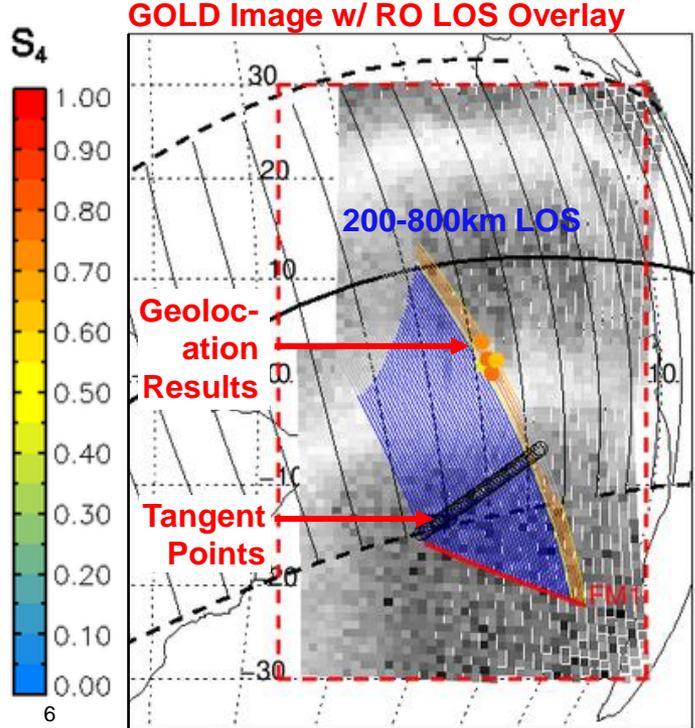
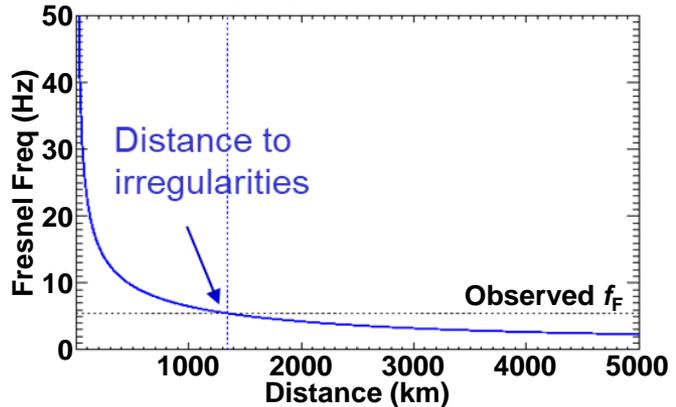
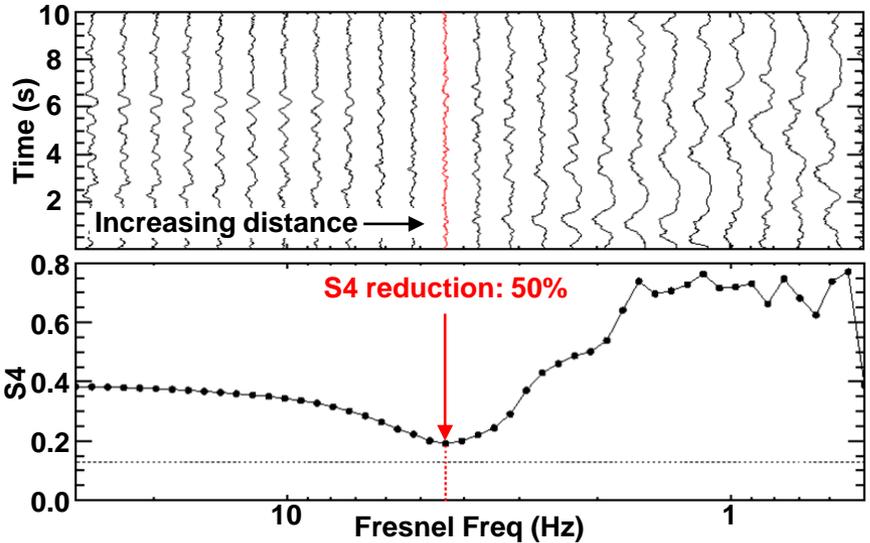
Public release of quality controlled on-board S4, σ_ϕ and HR data expected April 2021

Note: Geolocation, limb-to-disk and σ_ϕ are derived from high rate SNR and/or phase data



Geolocation via Back-Propagation/GOLD Comparisons

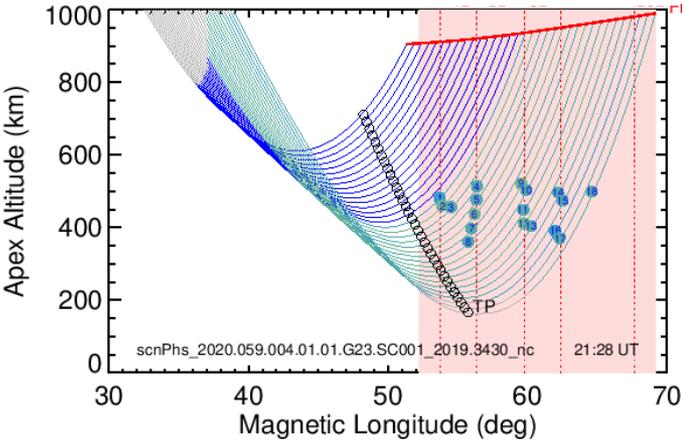
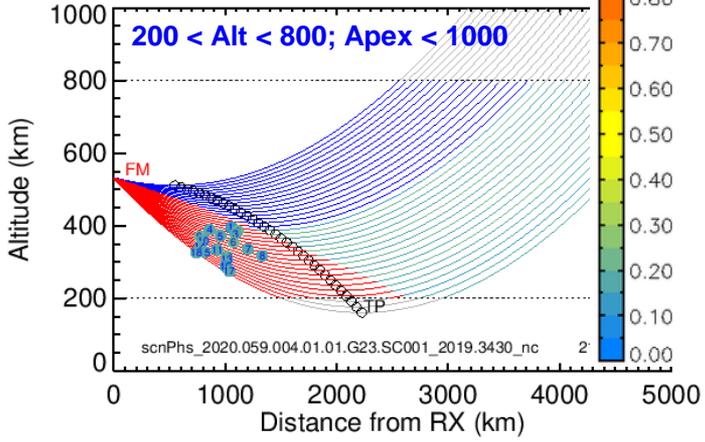
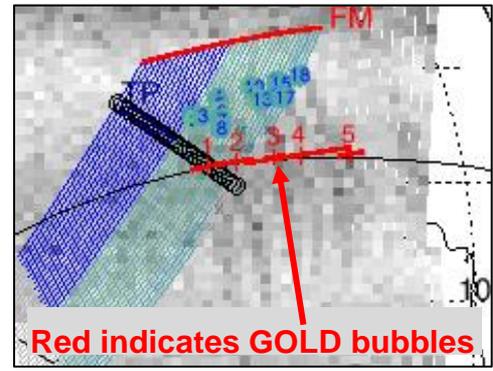
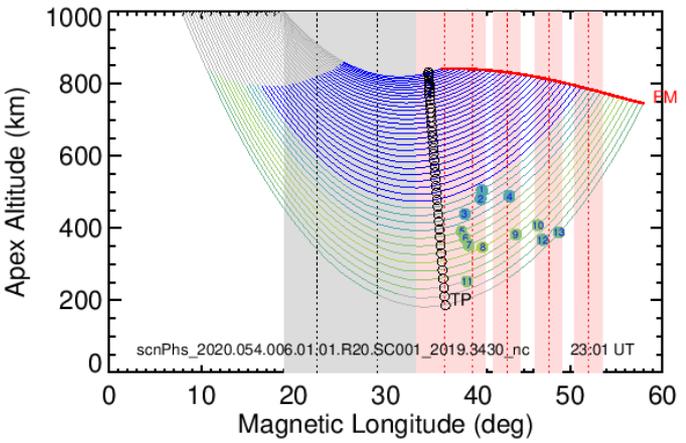
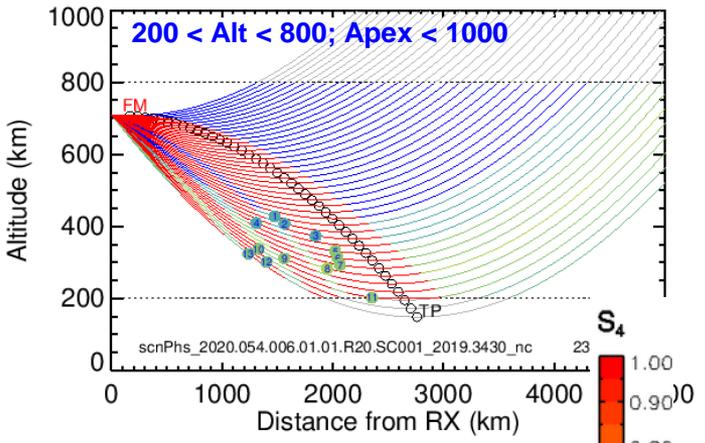
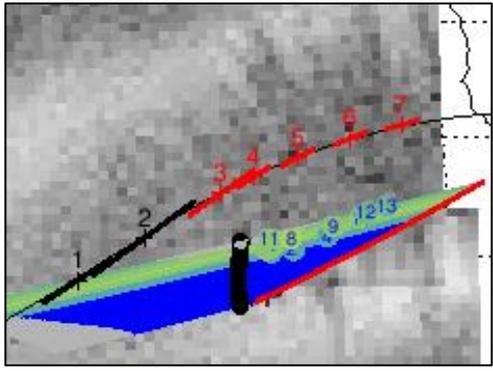
- Boston College (Carrano) back-propagation code operates in time domain, with Fresnel frequency as independent variable
- Rino (1977) scintillation model generalized to RO geometry relates Fresnel frequency to Fresnel scale & then to distance to irregularities
- GOLD 135.6 nm bubble imagery used for validation





Multiple Bubbles Detected in Single Occultations

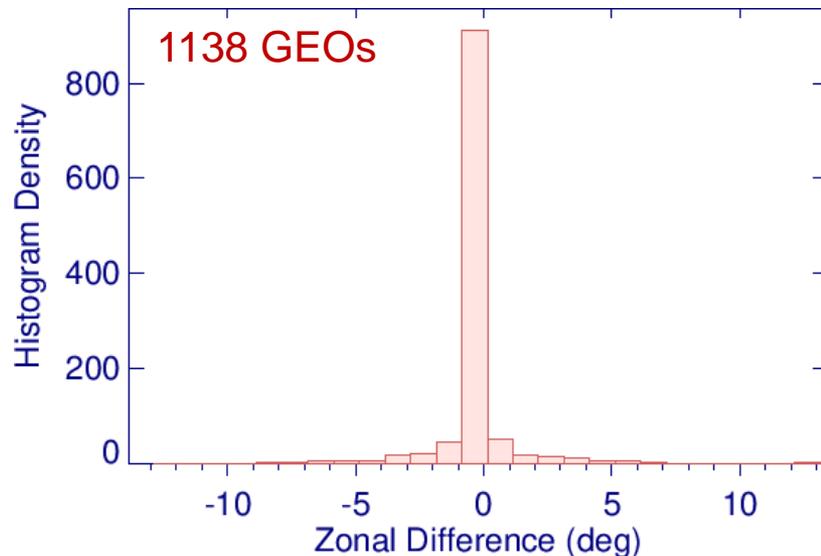
- Bubble locations/widths in GOLD data initially identified by UCAR algorithm, then manually sanity checked by BC/UCAR/Aerospace Cal/Val team
 - *GOLD bubbles validated vs. ground-based VHF scintillation data (Sao Luis, Brazil): 100% agreement wrt pre-midnight scintillation*
- Bubble geolocation determined based on 10s segments of high rate SNR/phase data
- Various occultation geometries result in potential detection of multiple bubble regions





Results of Validation with GOLD

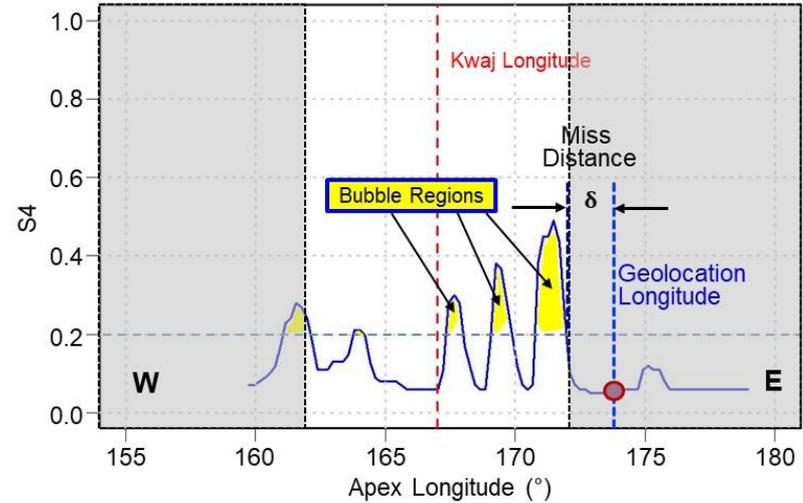
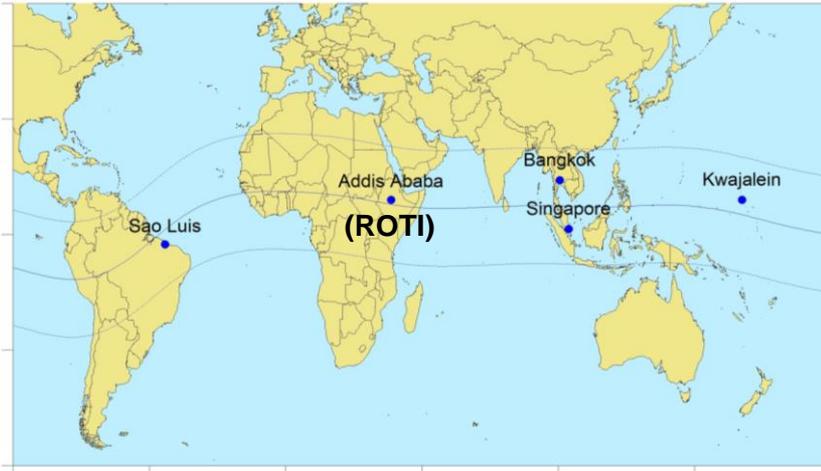
- February-April 2020 selected for performance evaluation
- Comparison of RO geolocations with bubbles detected by GOLD indicate agreement to within 2° geomagnetic longitude for better than 90% of cases
- Outliers are mostly associated with cases with poorly defined GOLD imagery structures, so truth was not well established (bubbles may be present at geolocation points)
- Caveat: in S. American/Atlantic sector scintillating nights often have long trains of bubbles – it may be easier to “hit” a bubble than to miss it!
- However, restricting the observational dataset to only cases containing isolated GOLD bubbles does not change the validation statistics
 - *This bolsters our confidence in the geolocation algorithm*



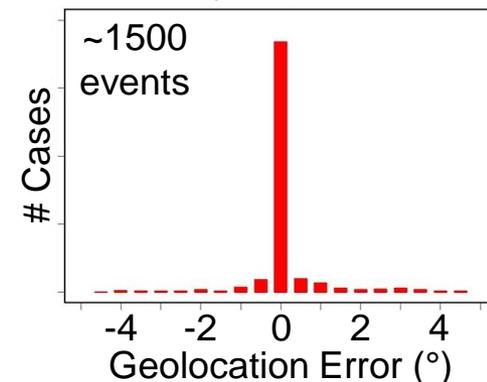
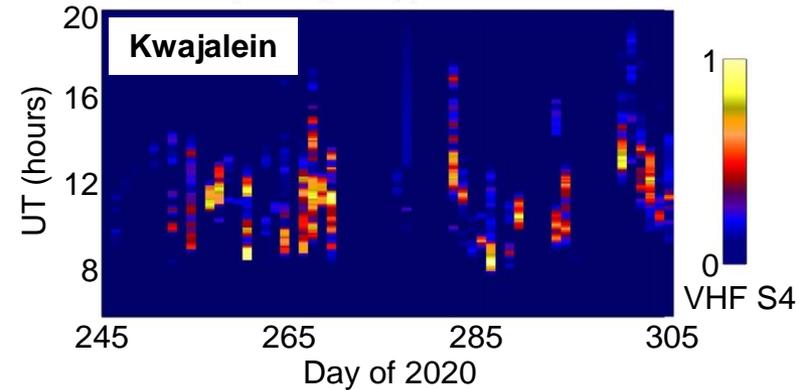
Error	All Bubbles % Cases	Isolated Bubbles % Cases
0°	71	73
$<1^\circ$	85	84
$<2^\circ$	91	92
$<5^\circ$	97	98



Geolocation Validation Outside of S. America



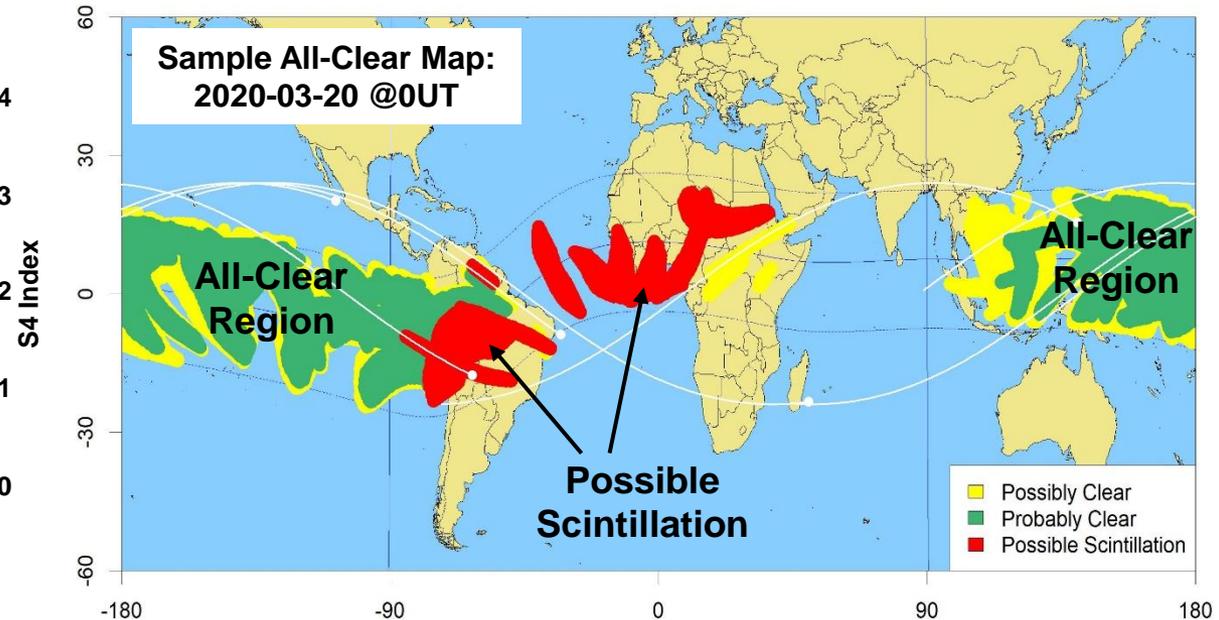
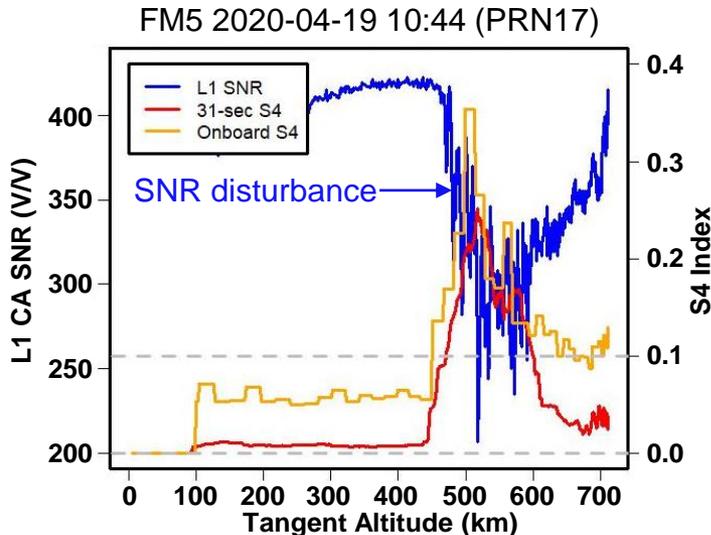
- Ground sensor for geolocation accuracy evaluation
 - Receivers tracking VHF geostationary satellite signals
 - GPS receiver ROTI data (Addis Ababa station only)
- Analysis assumes bubble structures observed by ground are “frozen in” and drift at 75 m/s
 - Bubbles identified relative to 0.2 S4 sensor noise floor & projected out for 2 hours (5° longitude)
 - Pre-midnight events only
- Geolocation accuracy for 4 non-S. American stations exceeds GOLD analysis (90/96% are within 1°/2° longitude)
- Restricting comparisons to isolated bubbles doesn't significantly degrade accuracy





TGRS Scintillation “All-Clear” Product

- All-Clear algorithm identifies irregularity-free regions using bubble physics to determine where along RO ray paths scintillation would be detected if present
 - Irregularity locations limited by apex altitude & plasma density (based on climatology)
 - Algorithm specifies absence of UHF scintillation for a ground-based observer
 - Employs “pseudo-S4” derived from 31s of TGRS L1 1Hz SNRs
- TGRS SNR disturbances caused by ground-based RFI create spurious pseudo-S4 values
 - RFI filtering algorithms developed to exclude bad data
- Validation w/ Kwajalein, Singapore, Sao Luis VHF stations
 - < 3% error in All Clear at Kwaj/Singapore, but almost 20% error at Sao Luis
 - Problem has been traced to incorrect exclusion of scintillated RO data (RFI filtering error)
 - Improved RFI filtering algorithm developed by JPL and will be tested





IVM Performance Requirements

COSMIC-2 JL1PRD Requirements

Parameter

- In-situ density accuracy/precision
- In-situ density measurement range
- In-track drift accuracy/precision
- Cross-track drift accuracy/precision
- Plasma drift measurement range
- Ion composition accuracy/precision
- Ion temperature accuracy/precision

Requirement

- 5% & 1%
- 10^3 to 5×10^6 cm⁻³
- ± 10 m/s & ± 5 m/s
- ± 5 m/s & ± 1 m/s
- ± 1000 m/s
- 5% & 5%
- 10% & 5%

Status

Verified*
 8×10^3 to 5×10^6 cm⁻³

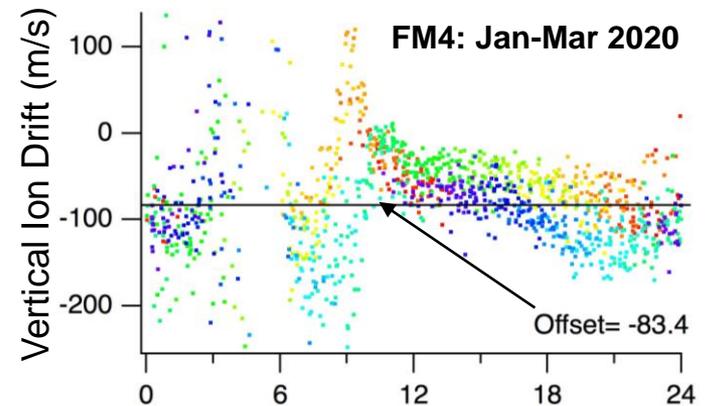
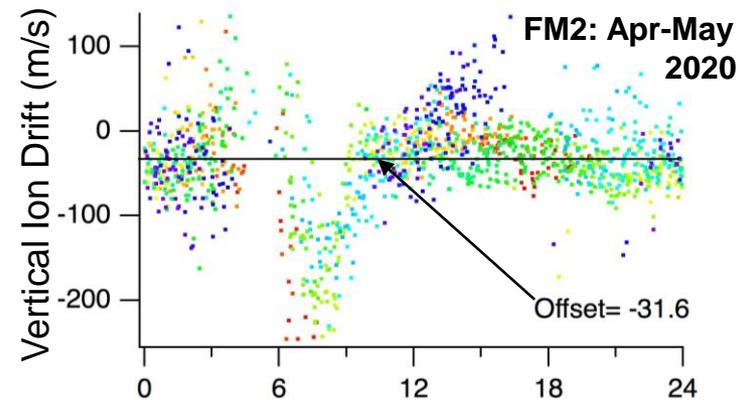
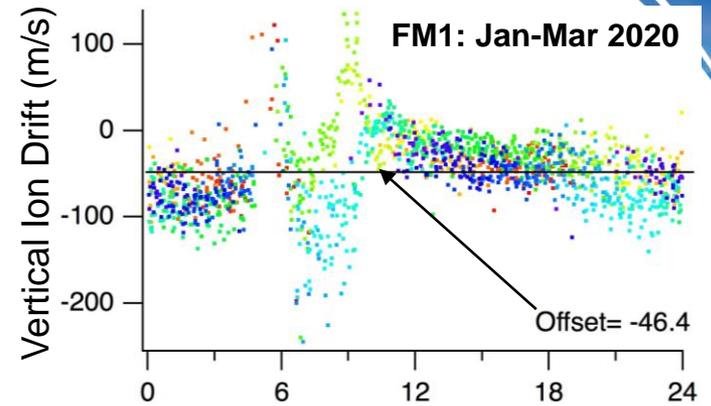
Cal/Val Objectives to Optimize Operational Utility

- Develop & evaluate “Bubble Maps” (joint TGRS/IVM product)

***See Wu et. al. poster for details**

IVM Ion Drifts

- Ion drift products can only be produced for COSMIC-2 satellites after lowering to 550 km
 - H^+ ion dominance at 720 km corrupts drift data
 - Even at 550km, low solar minimum nighttime densities & high H^+ fractions result in noisier drifts
 - Daytime data are highest quality, but pre-noon data required correction for photoelectron (PE) effects
 - Initial PE corrections have been developed, but further improvements are needed
- Local time variation of vertical drift at magnetic equator determined over ~60 day periods
 - Drift offsets determined for FM1, FM2 & FM4 for median drifts at 1800 LT
 - Currently investigating bias changes that occurred near the end of 2020
 - Offsets will be validated through comparisons to Jicamarca ISR



UTD analysis

RF Beacon Scintillation Requirements & Plan



COSMIC-2 JL1PRD Requirements

Parameter

- S4 Measurement Range
- S4 Measurement Uncertainty
- σ_ϕ Measurement Range
- σ_ϕ Measurement Uncertainty

Requirement

- 0 to 1.5
- 0.1
- 0-20 radians
- 0.1 radians

Cal/Val Plan

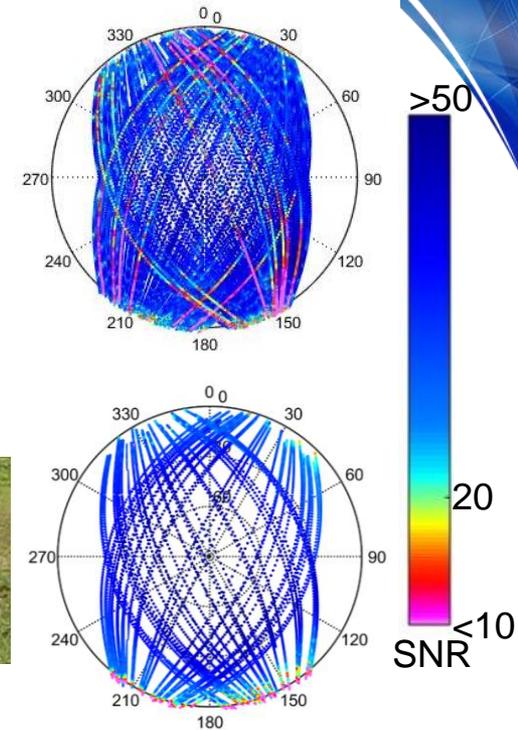
- Two RF Beacon Receiver (RFBR) ground sites
 - *Kwajalein (AFRL Roi-Namur site)*
 - *Sao Luis, Brazil*
- Areas of work
 - *Quality control assessment (leading to RFBR s/w improvements)*
 - Algorithms to distinguish between real scintillation & multipath effects
 - Mapping the RF Beacon antenna pattern on the COSMIC-2 spacecraft (particularly L-band)
 - *ALTAIR radar validation campaigns*
 - Dry run in October 2020/final in April or September 2021
 - ALTAIR measures 2-way scintillation to compare with RFBR 1-way measurements
 - Signal propagation models will be used to extrapolate between radar & RFBR frequencies

RF Beacon Cal/Val Status

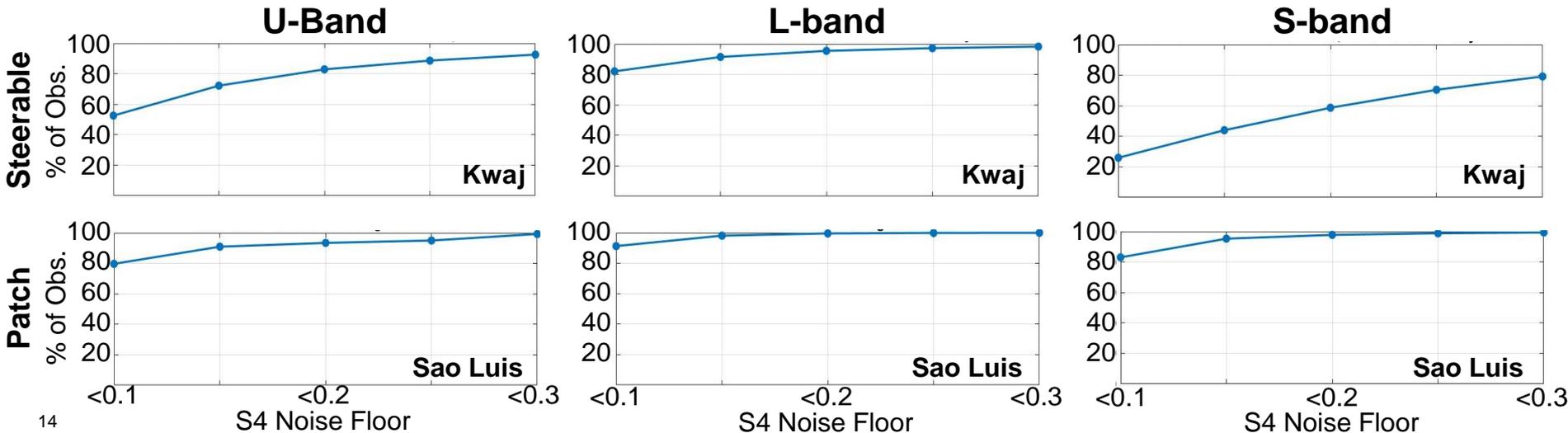
- Cal/Val is comparing two different RFBR antennas
 - Higher gain steerable Yagis (@Kwaj May 2019)
 - Fixed tri-band patch (@Sao Luis Feb 2020)
- COVID-19 site restrictions delayed 2nd RFBR installs & has slowed system troubleshooting
 - Patch system added to Kwaj (Aug 2020)
 - Steerable system added to Sao Luis (Oct 2020)
- Steerable system performance problems caused by antenna polarization errors → corrected!
- Working to understand predicted vs actual performance for the different RFBRs & extract accurate phase observations



UHF SNR Maps



Figures courtesy AFRL/Boston College





Summary

- The COSMIC-2 space weather Cal/Val is ongoing
- COSMIC-2 sensor products in the areas of ionospheric density, scintillation and plasma drift specification are being validated for operational and scientific use
 - *TGRS GPS & GLONASS TEC product validation is complete*
 - *IVM in-situ density validation is complete*
 - *TGRS scintillation & additional IVM products will be made available for operational and scientific use as validation work completes (spring – summer 2021)*
- RF Beacon signals are available to any one that fields a receiver capable of measuring them

The COSMIC-2 sensors provide unprecedented coverage and refresh in support of equatorial space weather applications