



Assessment of COSMIC-2 Radio Occultation (RO) bending angle bias in the lower troposphere

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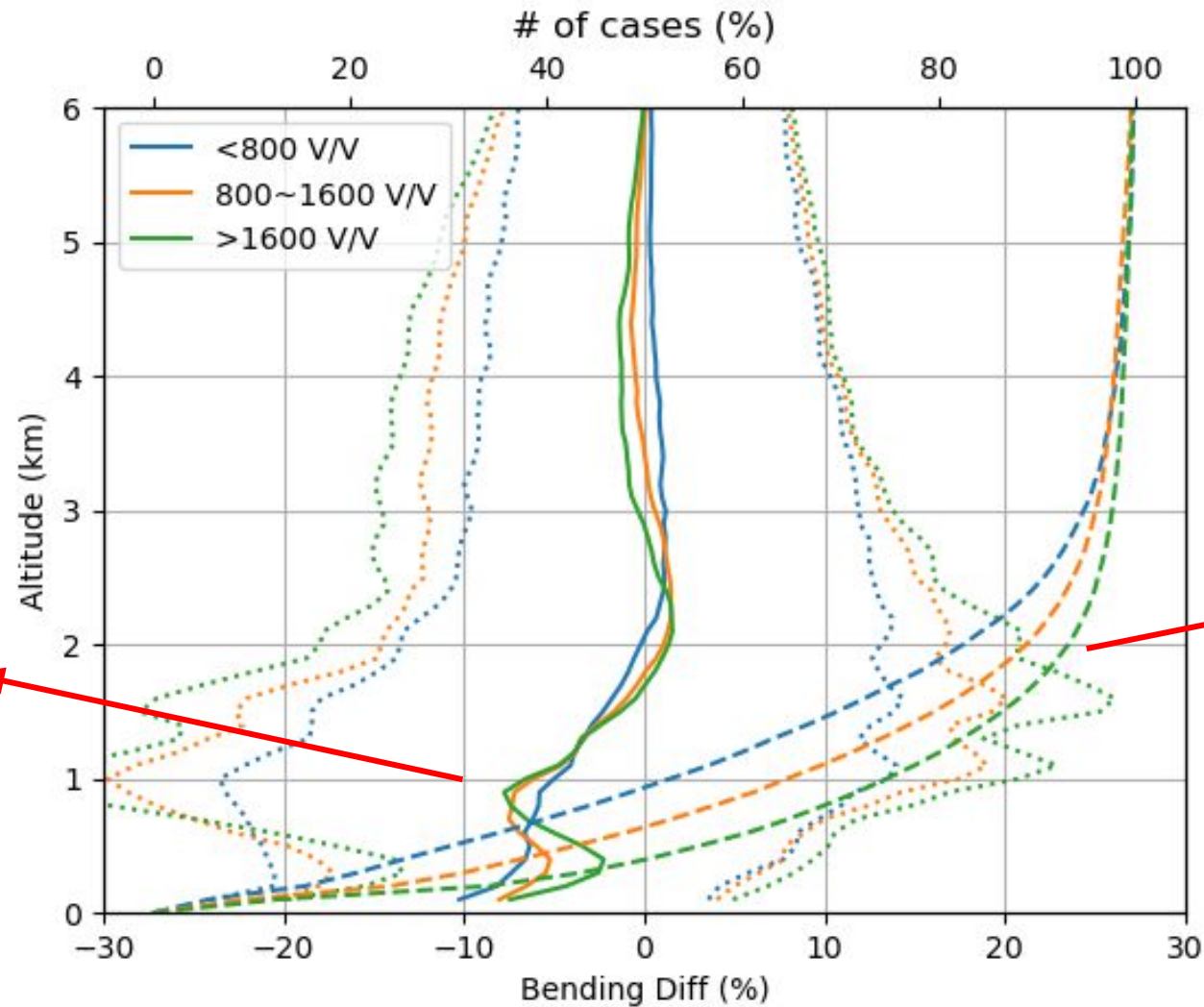
Jet Propulsion Laboratory
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This document has been reviewed and determined not to contain export controlled technical data.

Introduction

- Bending angle bias: it has been shown that >80% of COSMIC-2 observations under 2km failed the quality control (QC) of the data assimilation (DA) process due to sizeable bending angle biases
- Higher SNR (>1000 V/V) of COSMIC-2 does not improve the biases below 2 km. Several possible error sources causing the negative biases are investigated
- Error source: Measurement
 - Smoothing
 - Noise
 - Sampling bandwidth
 - Receiver model

Bending O-A (COSMIC2 2020-04 ~ 2020-05)



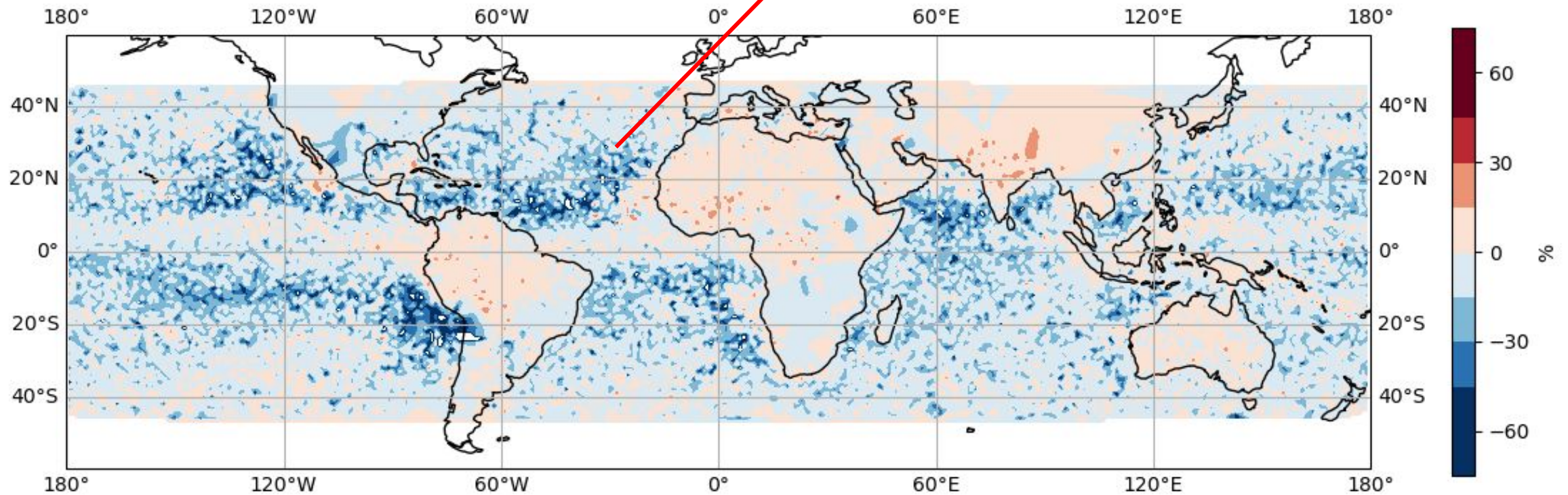
Penetration height improved when SNR is higher

Negative bias > 8% at 1 km even in strong SNR cases

Bending O-A (COSMIC2 2020-04 ~ 2020-05)

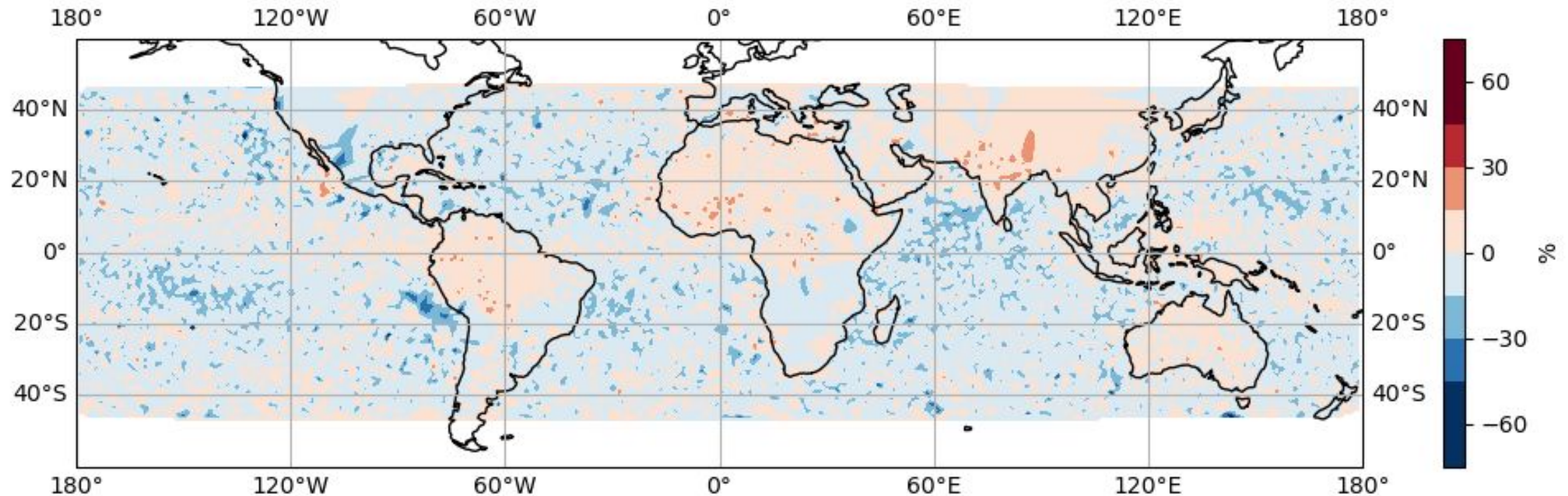
- COSMIC2 – NCEP: 1 km

Significant bending difference
in subtropical regions



Bending O-A (COSMIC2 2020-04 ~ 2020-05)

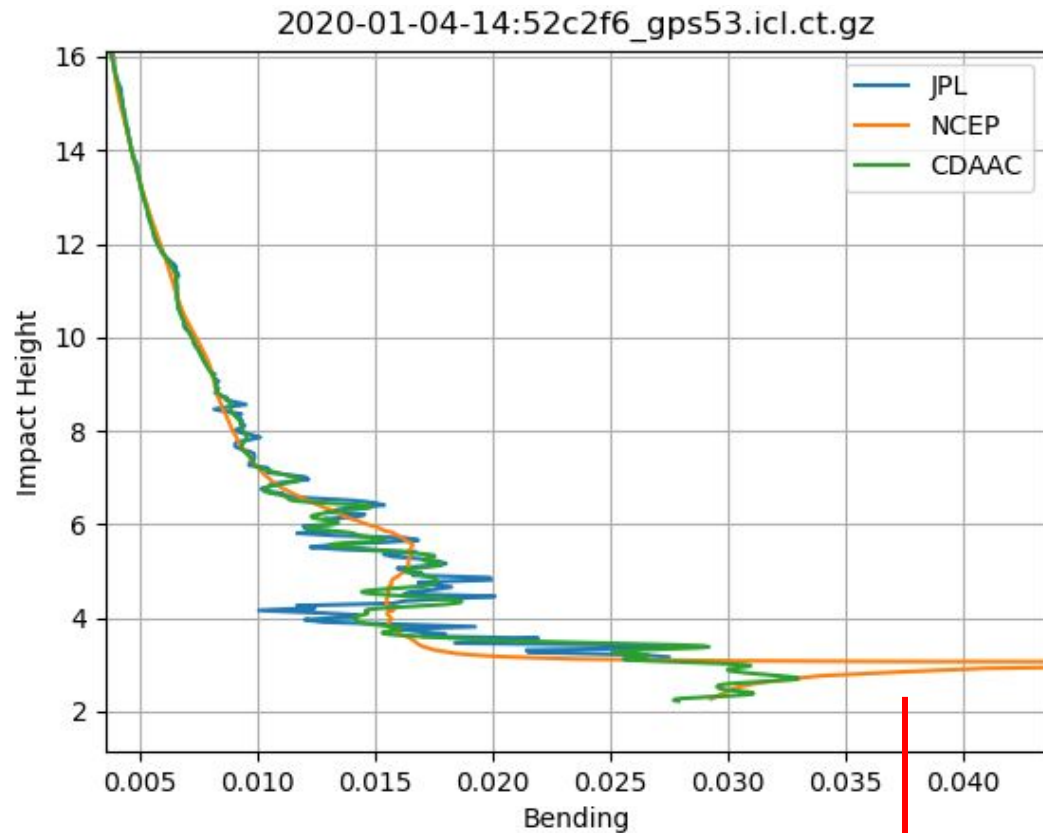
- COSMIC2 – NCEP: 1 km (max bending < 0.04 rad: 15% cases removed)



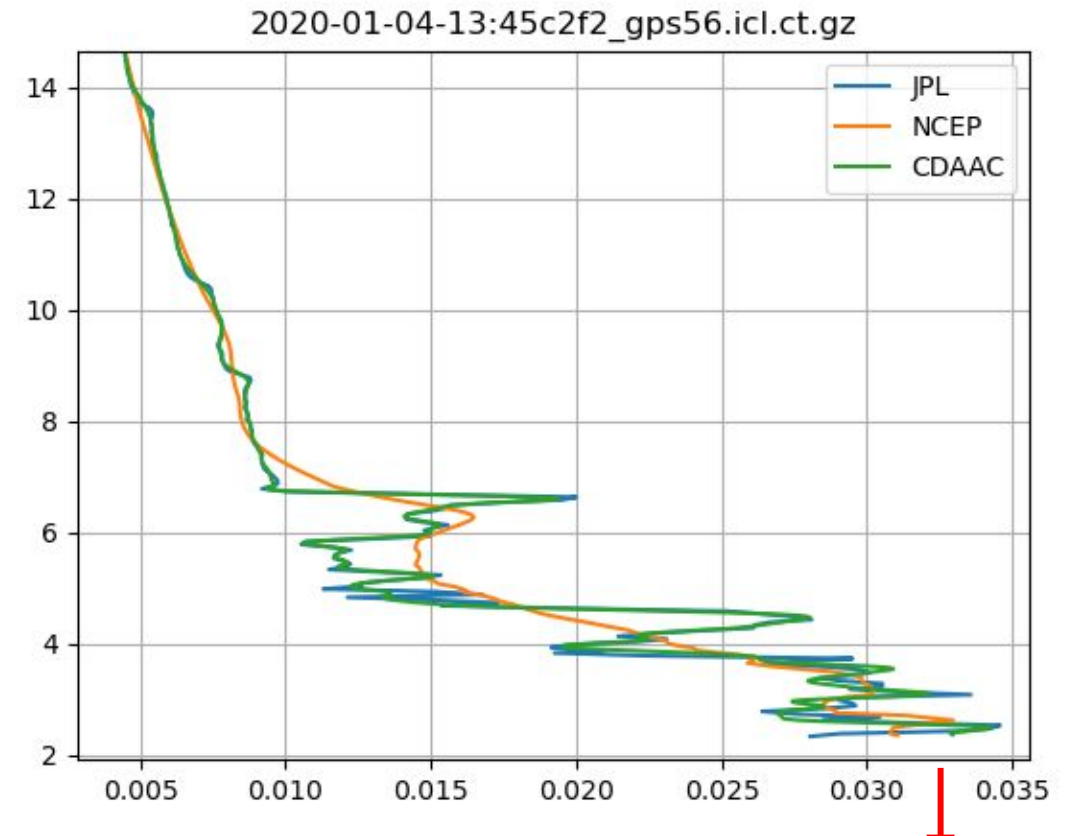
The negative bias at 1km mainly comes from the cases with strong N gradient!

Bending angle actual cases

- Individual cases (large peak BA vs small peak BA)



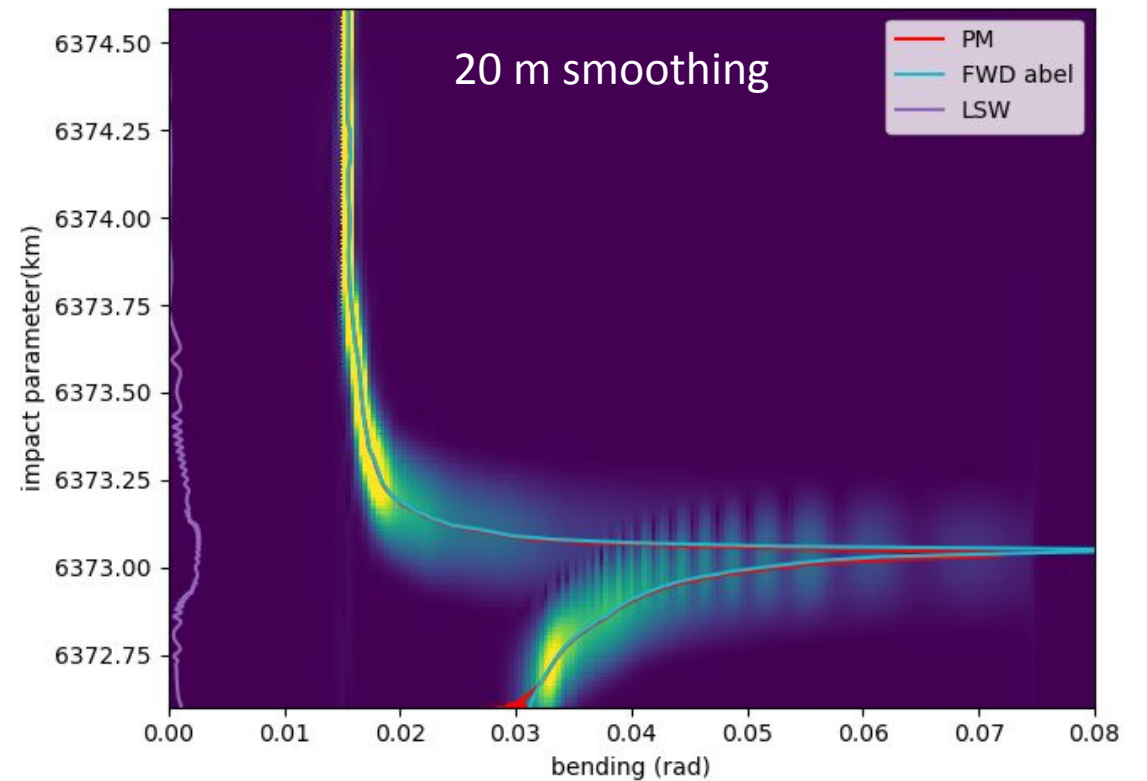
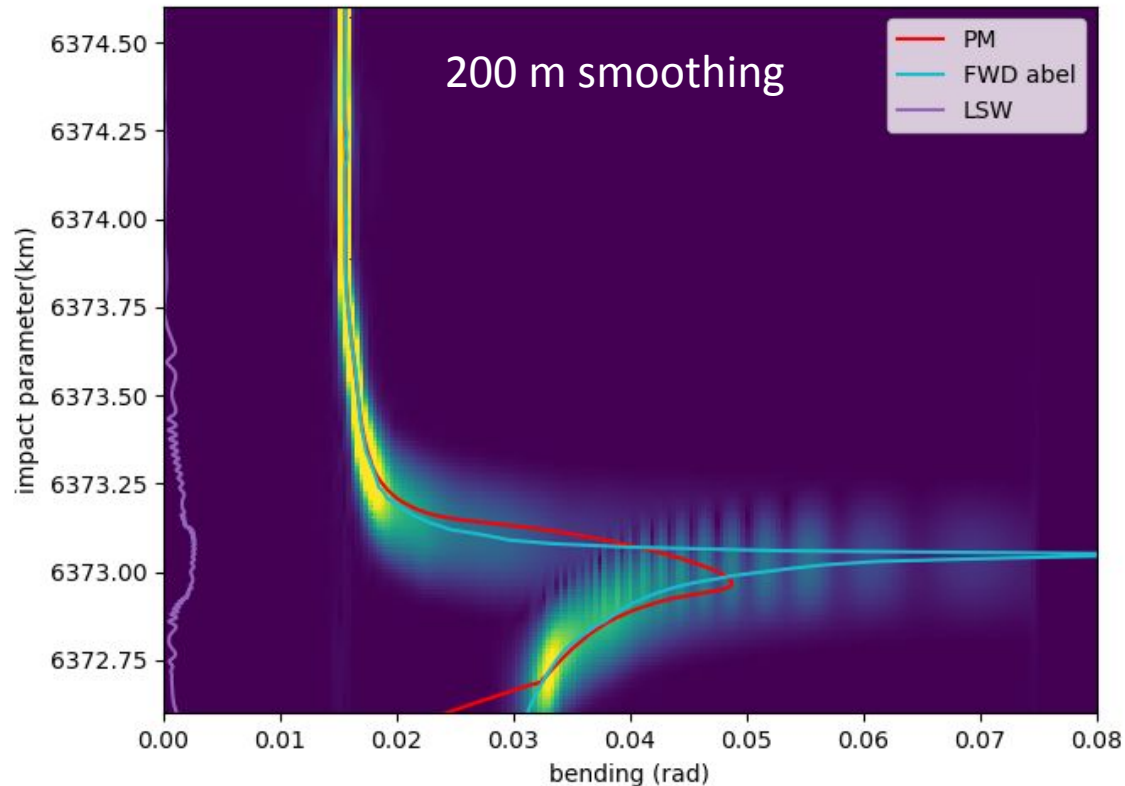
Large bias when strong peak presents



No clear bias if bending maximum < 0.04 rad

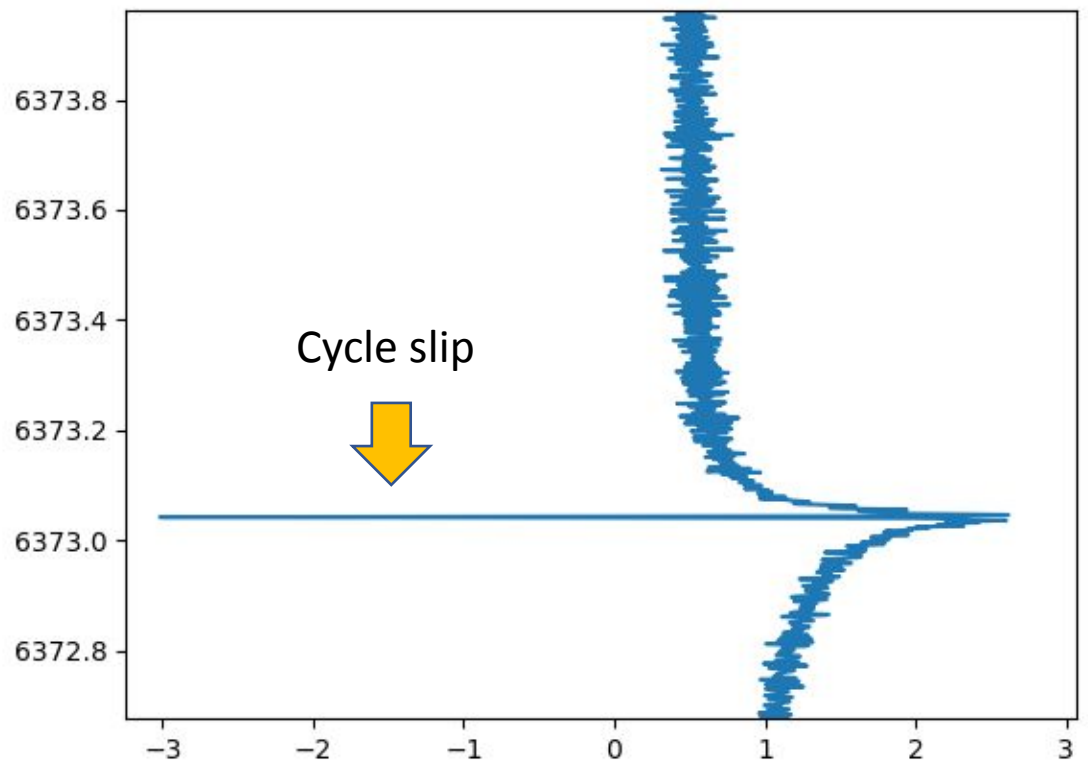
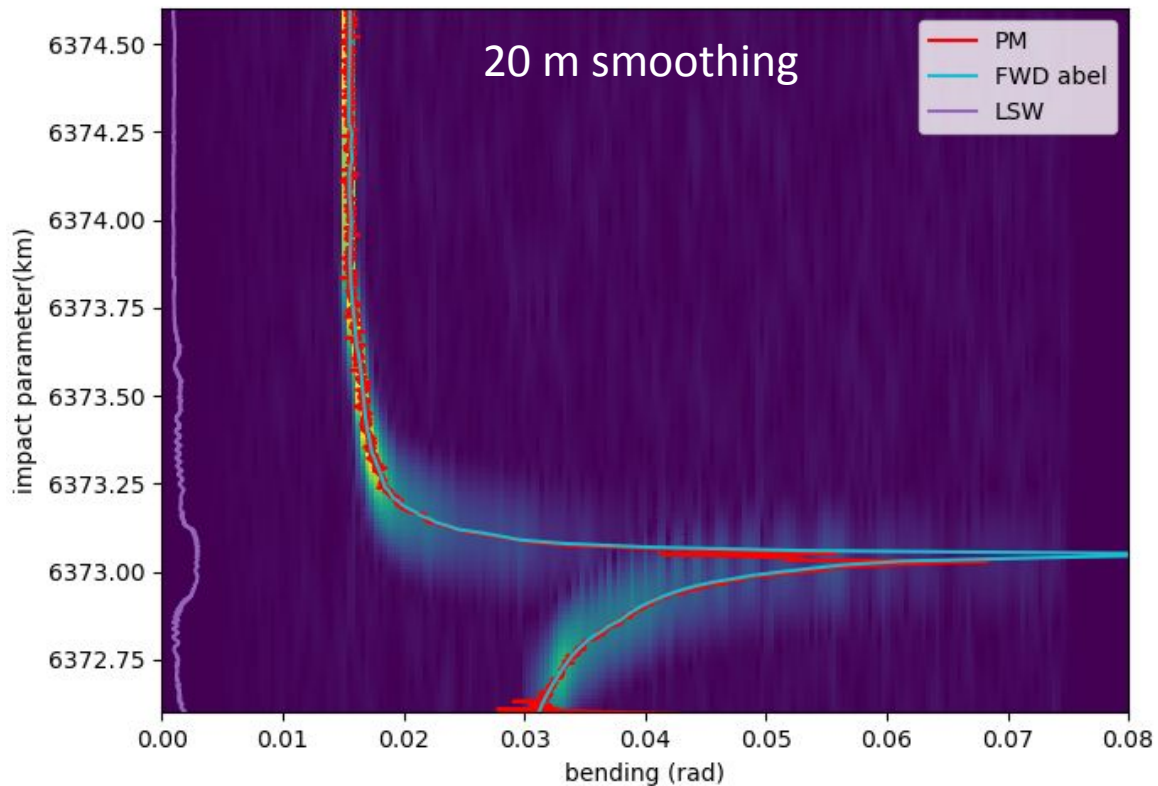
MPS simulation

- Vertical smoothing (200m vs 20m)
 - Forward Abel
 - Multiple Phase Screen (MPS) + Phase Matching (PM)



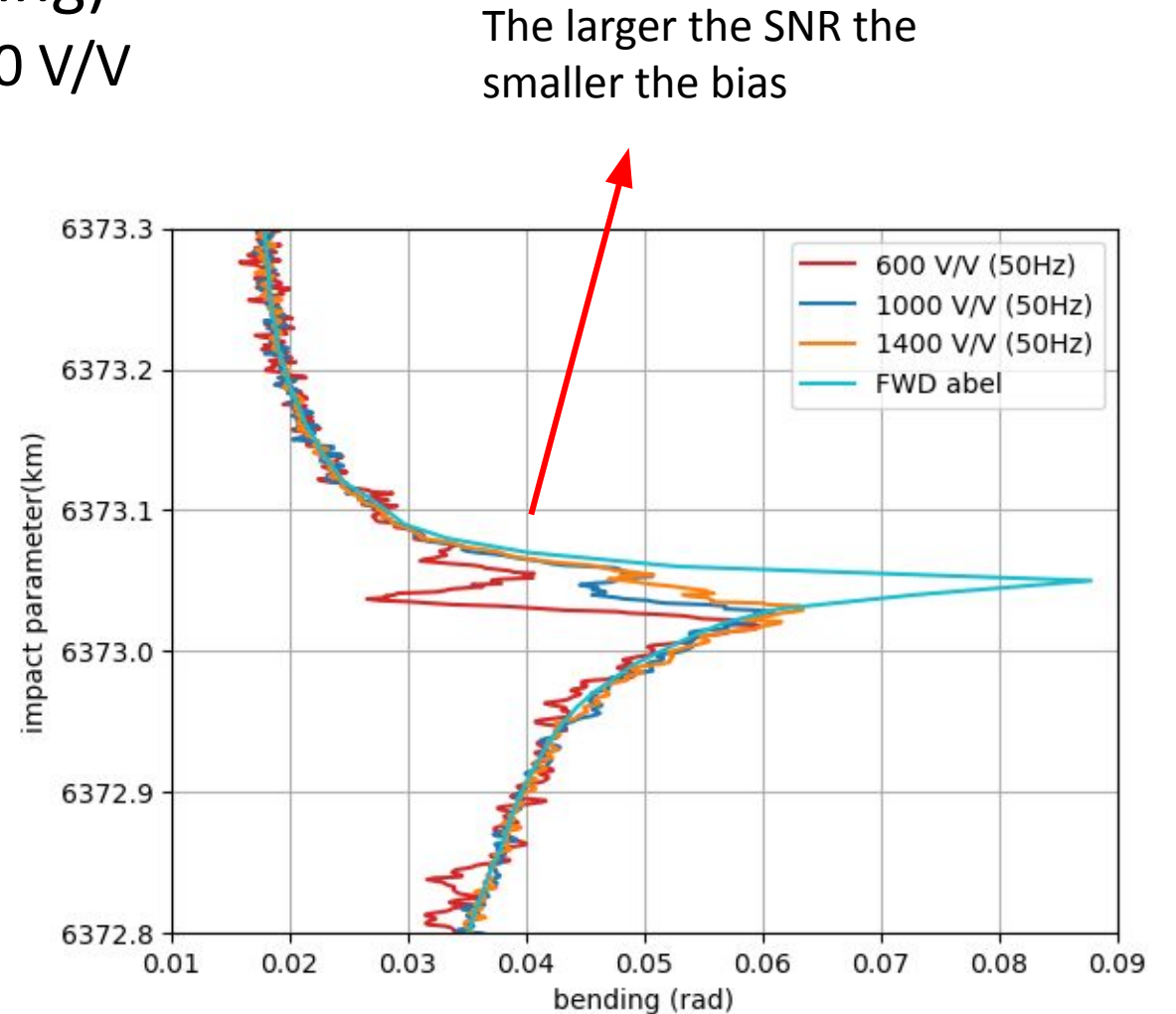
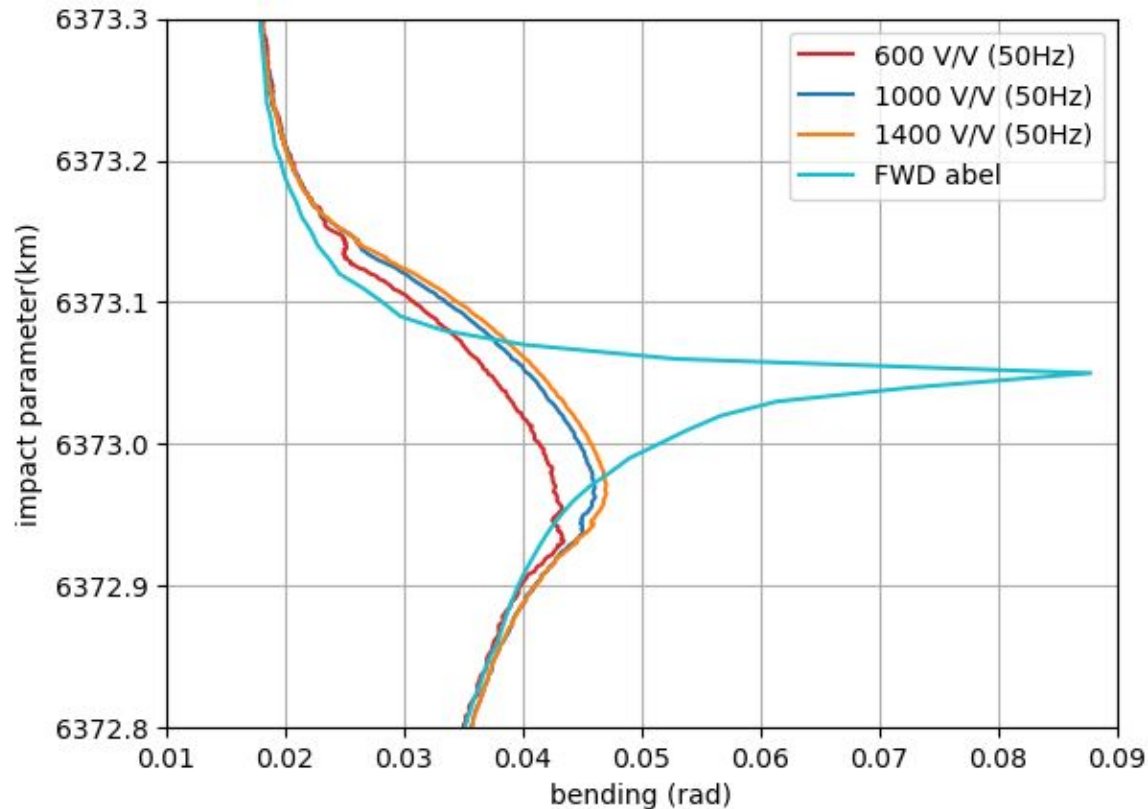
MPS simulation

- Noise (600 V/V w/ 20m smoothing)
 - Bending bias at the peak shows up when noise is applied
 - This is mainly due to noisy phase or cycle slip in the impact parameter domain



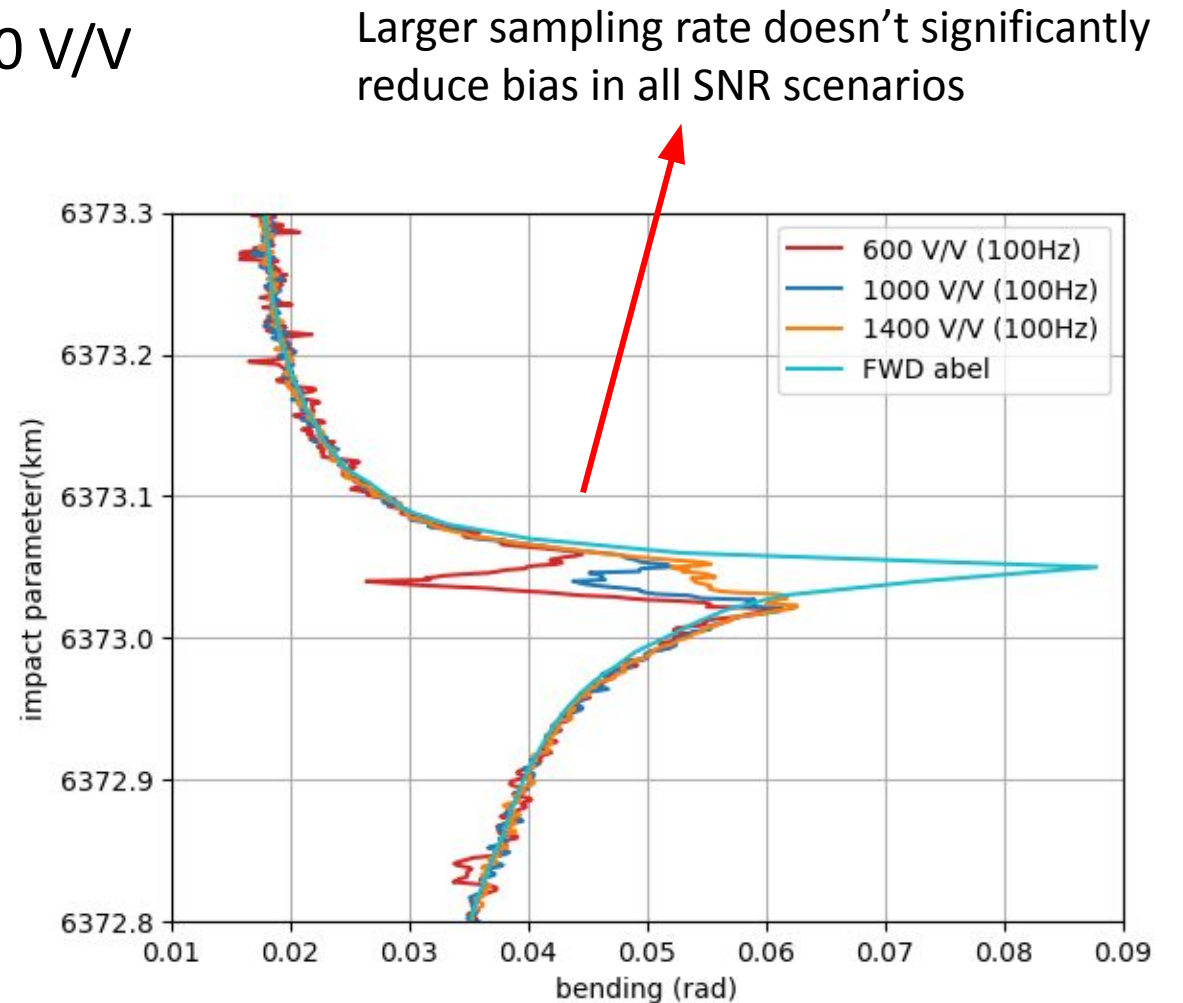
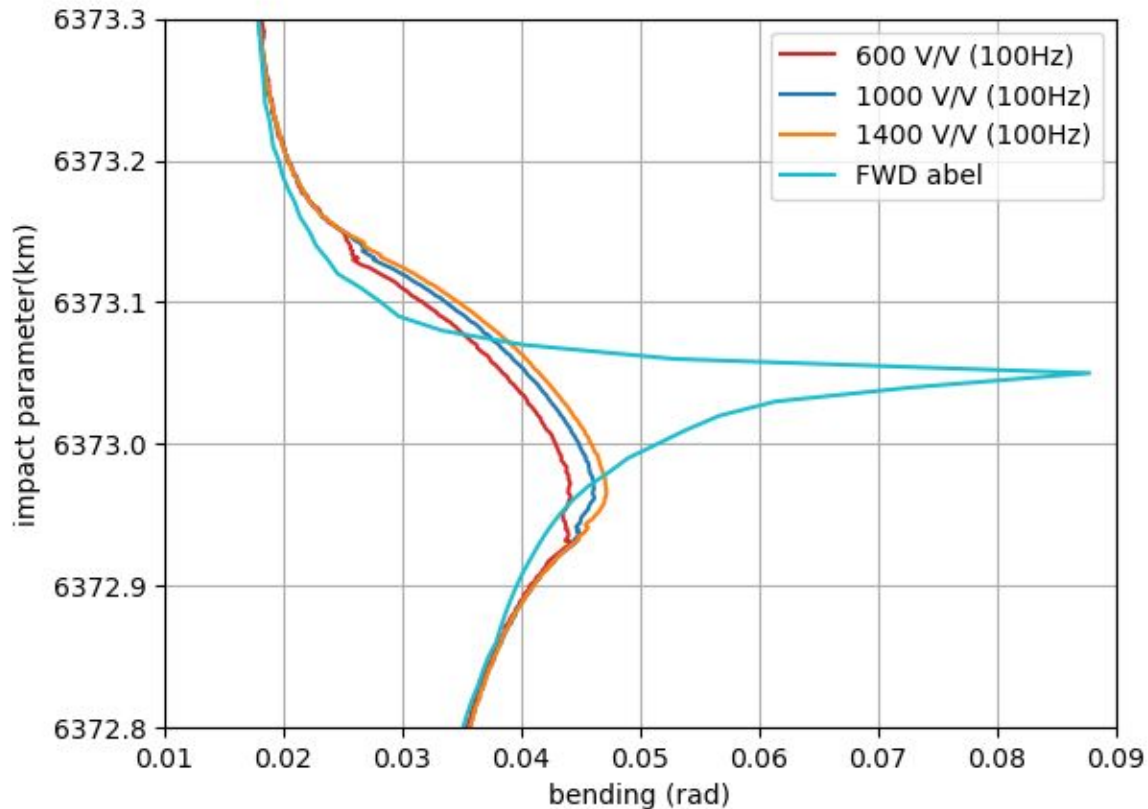
MPS simulation

- Noise - Sensitivity test (50Hz Sampling)
 - SNR scenario: 600 V/V, 1000 V/V, 1400 V/V
 - Monte Carlo analysis



MPS simulation

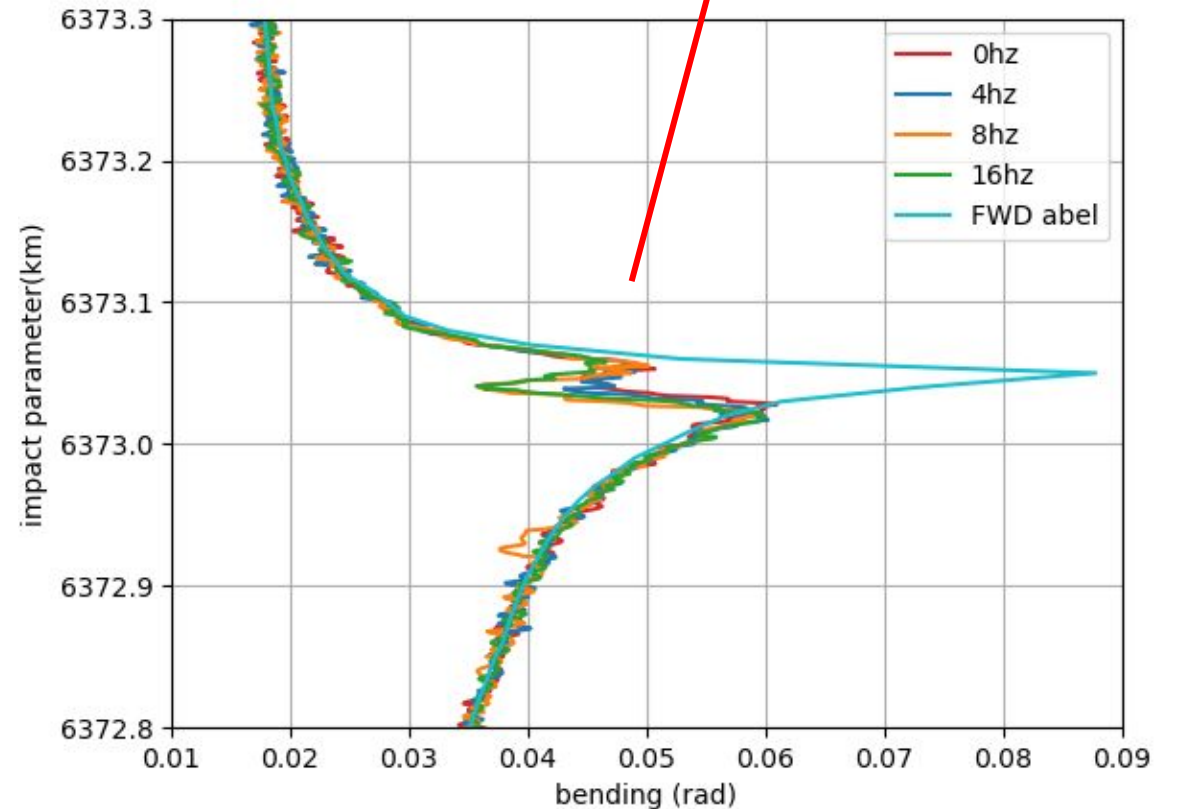
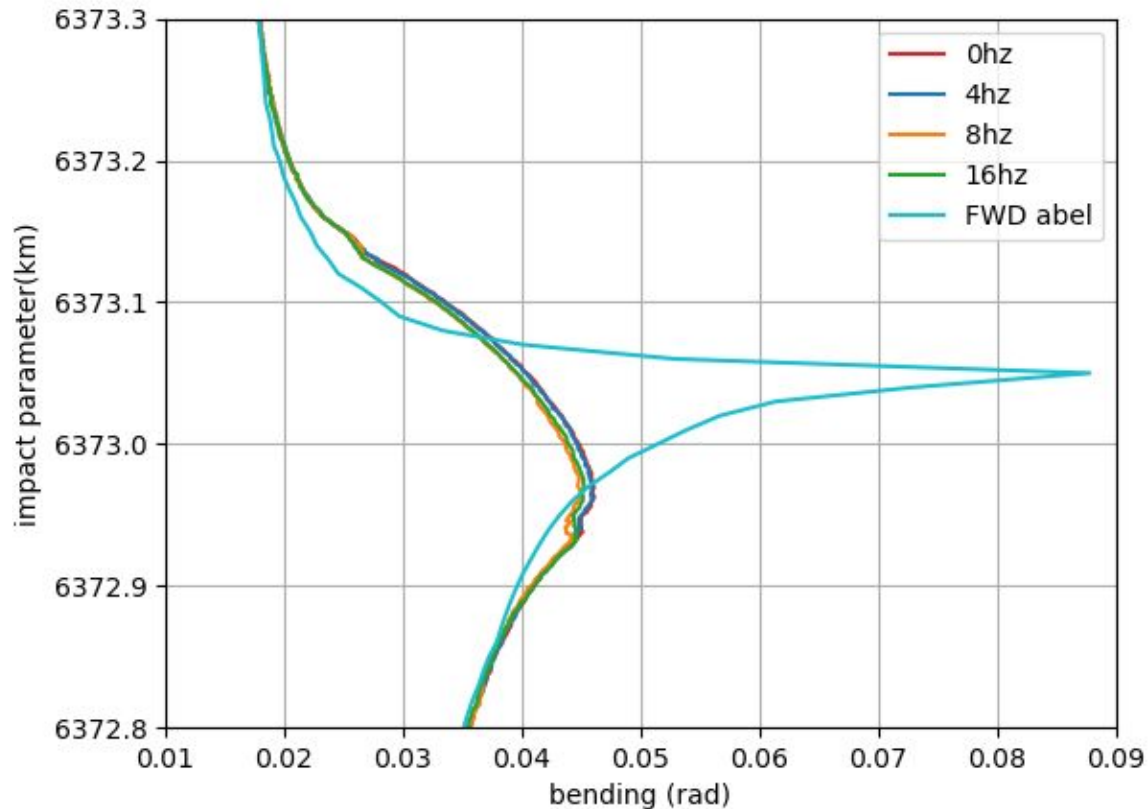
- Noise - Sensitivity test (100Hz sampling)
 - SNR scenario: 600 V/V, 1000 V/V, 1400 V/V
 - Monte Carlo analysis



MPS simulation

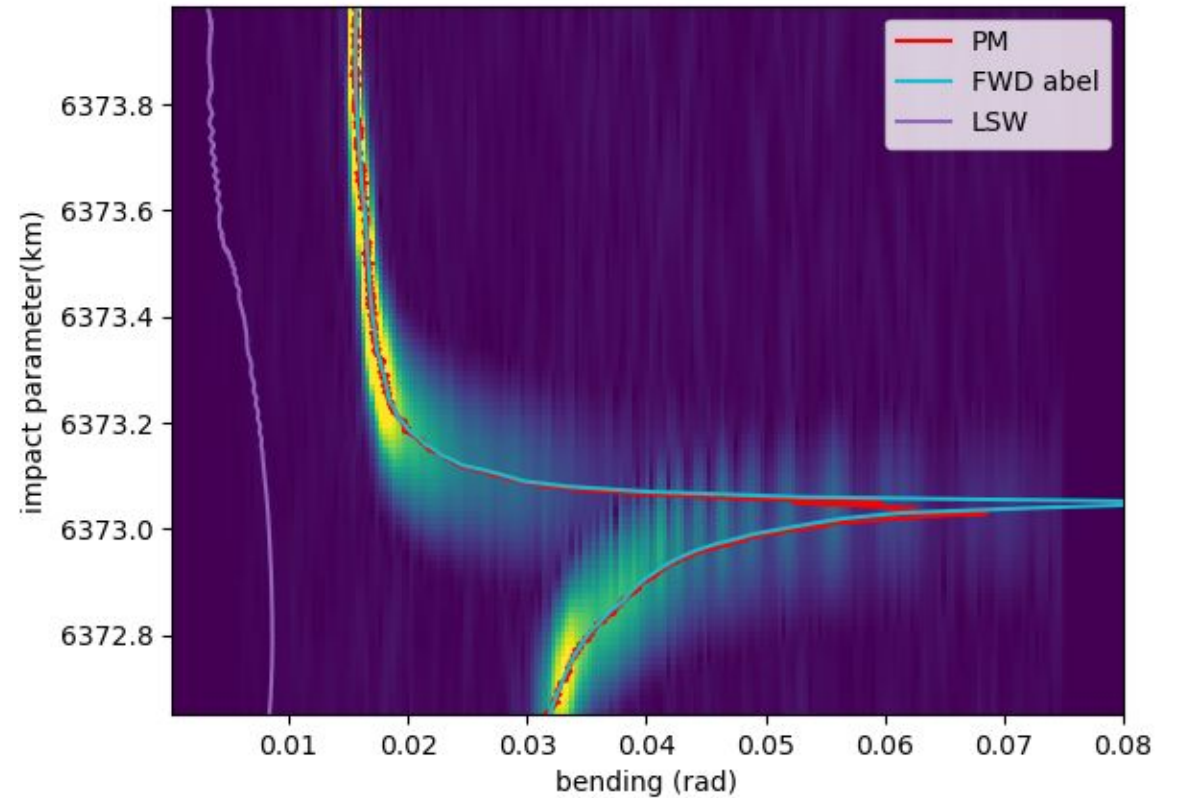
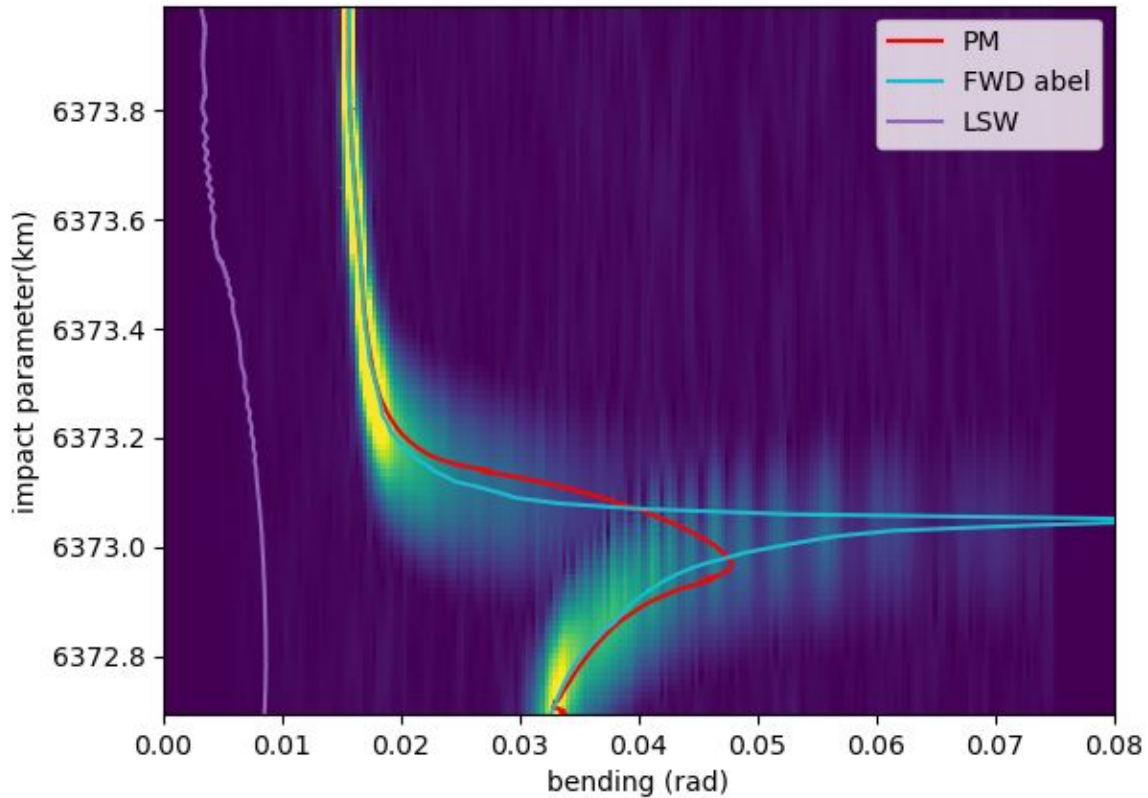
- Receiver model (1000V/V and 50Hz sampling)
 - Model deviation scenario: 0Hz, 4Hz, 8Hz, 16Hz
 - Monte Carlo analysis

The receiver model has limited influence to the bending bias



MPS simulation

- Receiver model (1000V/V and 50Hz sampling)
 - $N=280 \cdot \exp(-z/7)$



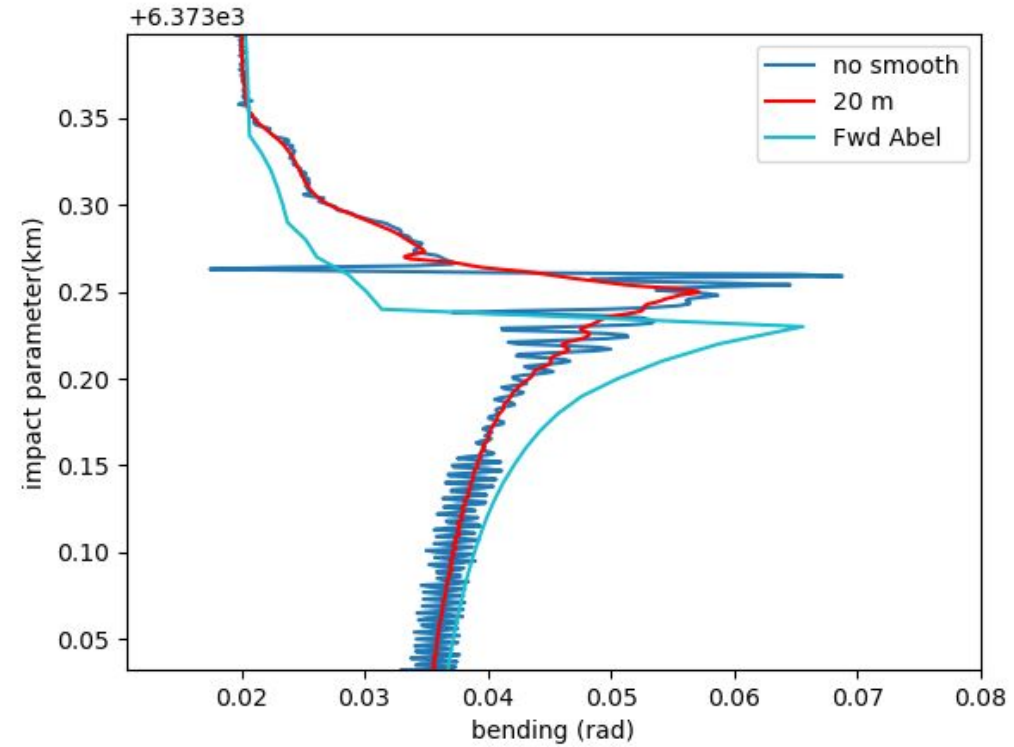
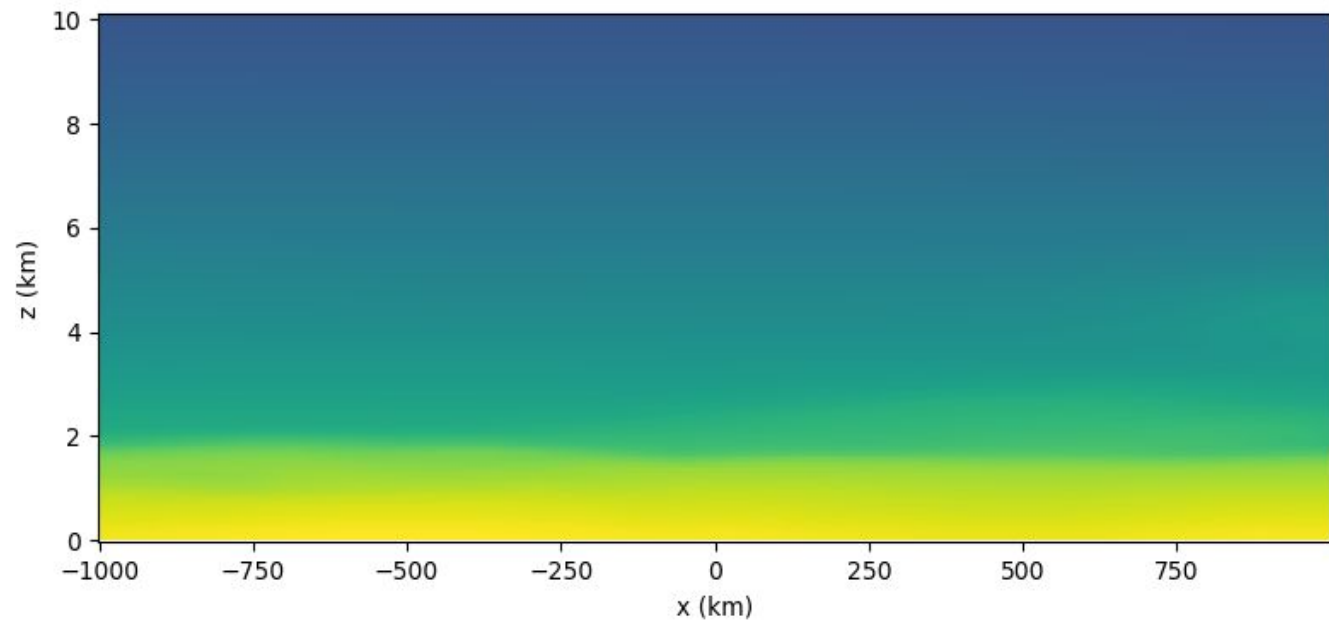
Summary

- Negative RO bending angle bias w.r.t NCEP analysis can be found at $\sim 1\text{km}$, and is more significant in high refractivity gradient regions
- Based on the MPS analysis the negative bias can be connected to the vertical smoothing, noise level, and the receiver model. The smoothing and noise contribute most of bending angle bias around the peak
- The horizontal inhomogeneity in different scales could also influence the bending angle bias, which needs further studies.

Thank you for your attention!

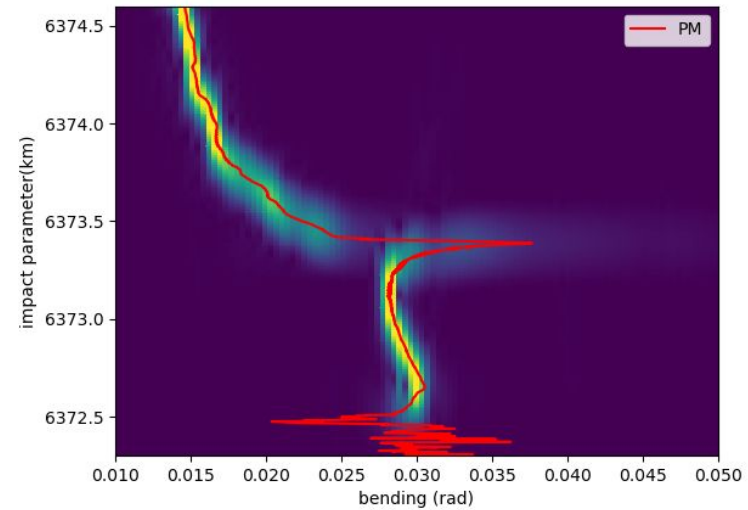
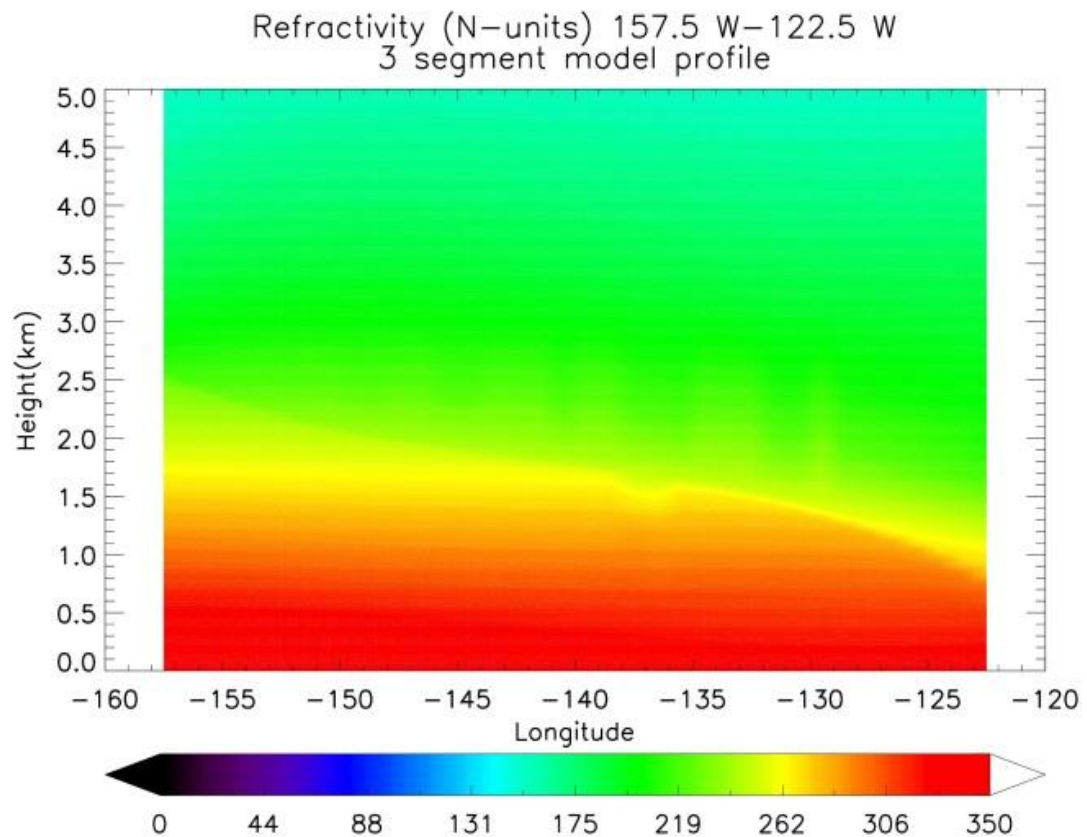
MPS simulation

- Horizontal inhomogeneity

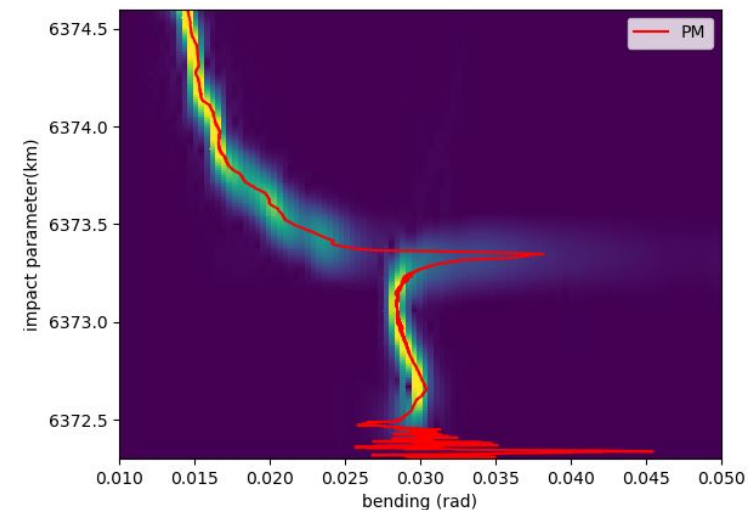


MPS simulation

- Horizontal inhomogeneity (20m smoothing) (Courtesy to Thomas Winning)



1D

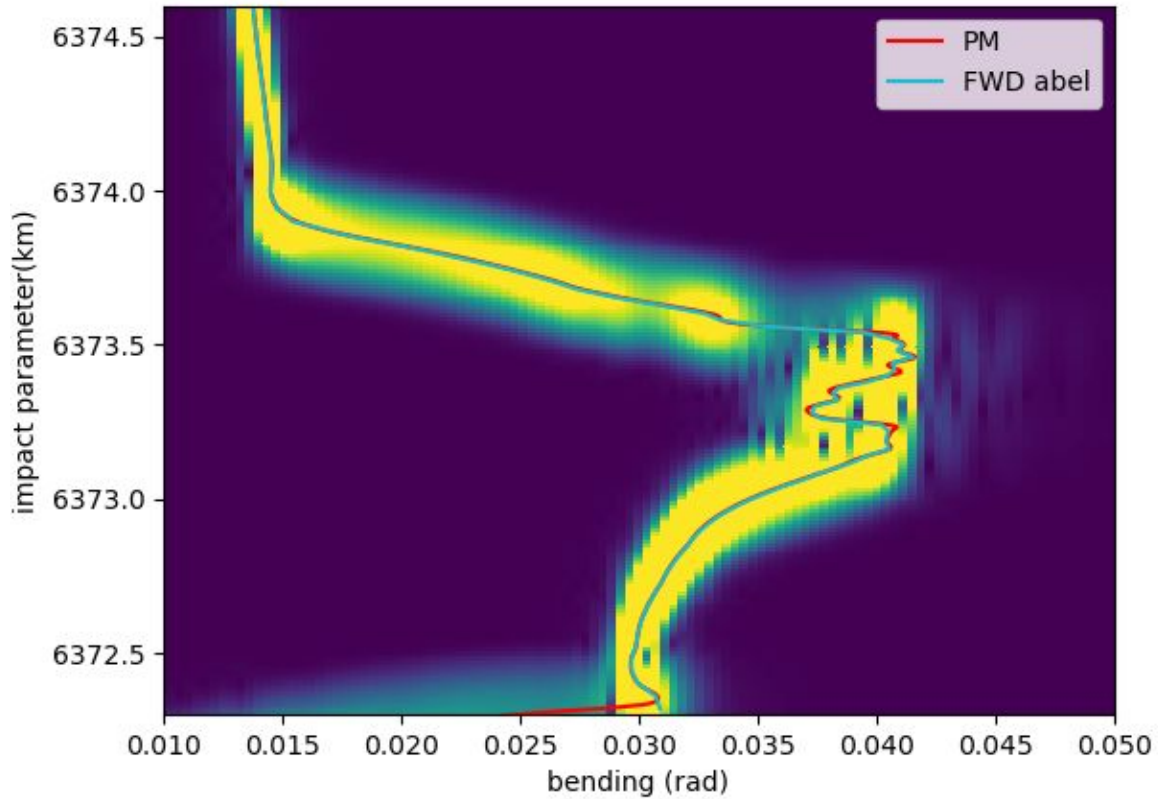


2D

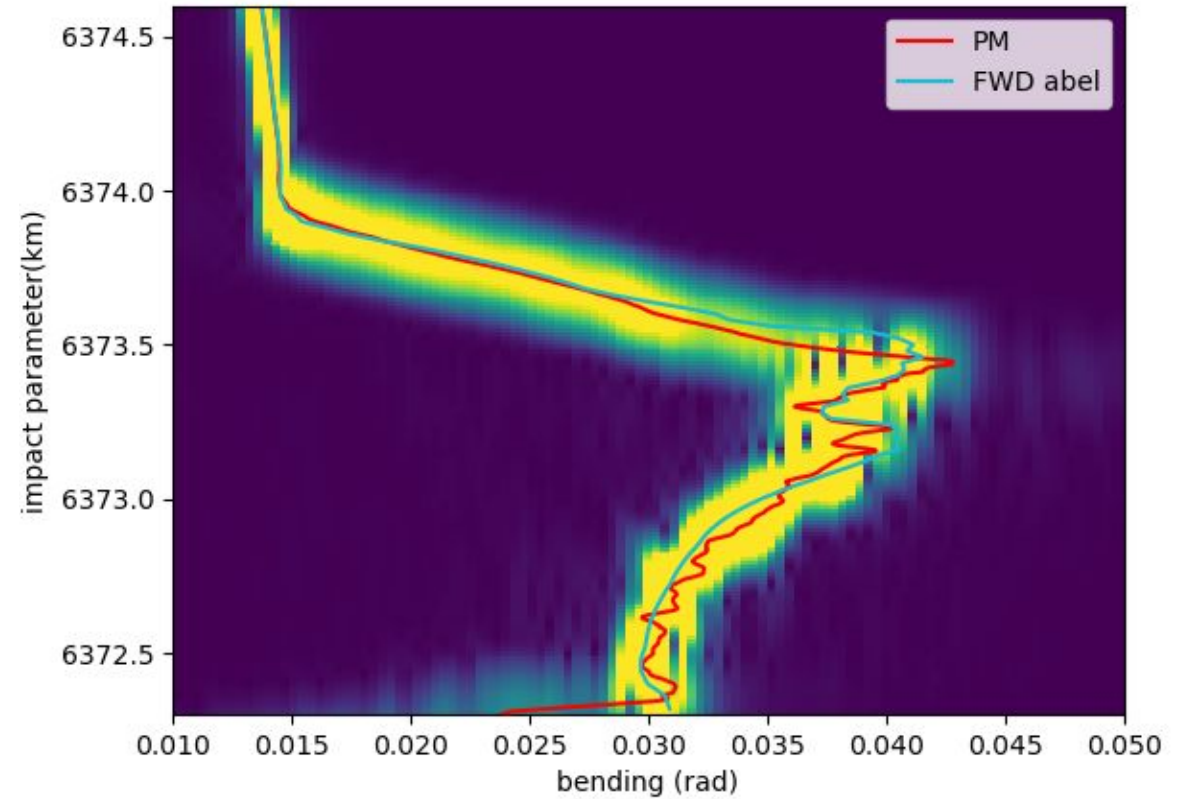
MPS simulation

- Turbulence (20m smoothing) (LES)

1D



2D



Forward Operator

- ROPP approach [Healy and Thepaut, 2006]
 - When the N decreases with height: Assuming the refractivity profile is exponential between each levels

$$\Delta\alpha = 10^{-6}\sqrt{2\pi a k_i N_i} \exp(k_i(x_i - a)) \left[\operatorname{erf}\left(\sqrt{k_i(x - a)}\right) \right] \Big|_{x_i}^{x_{i+1}}$$

$$k_i = \frac{\ln(N_i/N_{i+1})}{(x_{i+1} - x_i)} > 0$$

- When the N increases with height: Assuming the refractivity gradient is constant between each levels

$$\Delta\alpha = -2\sqrt{2a}10^{-6} \frac{(N_{i+1} - N_i)}{(x_{i+1} - x_i)} \left[\sqrt{(x - a)} \right] \Big|_{x_i}^{x_{i+1}}$$

Forward Operator for Ducting

- Separate the Abel integral

$$\alpha(a_0) = -2a \left[\int_{r_0}^{r_m} + \int_{r_m}^{r_t} + \int_{r_t}^{\infty} \right] \frac{dr}{\sqrt{x^2 - a_0^2}} \frac{d \ln n}{dr}$$

[Ao et al, 2007] [Xie et al, 2006]

$$k_i = \frac{\ln(N_i/N_{i+1})}{(x_{i+1} - x_i)}$$

