

# Joule heating scaling in thermosphere-ionosphere models compared to EISCAT incoherent scatter radar measurements

Florian Günzkofer<sup>1</sup>, Hanli Liu<sup>2</sup>, Huixin Liu<sup>3</sup>, Gunter Stober<sup>4</sup>, Gang Lu<sup>2</sup>, Haonan Wu<sup>2</sup>, Joe McInerney<sup>2</sup>, Frank Heymann<sup>1</sup>, and Claudia Borries<sup>1</sup>

<sup>1</sup> Institute for Solar-Terrestrial Physics, German Aerospace Center (DLR)

<sup>2</sup> High Altitude Observatory, NCAR

<sup>3</sup> Department of Earth and Planetary Science, Kyushu University

<sup>4</sup> Institute of Applied Physics, University of Bern

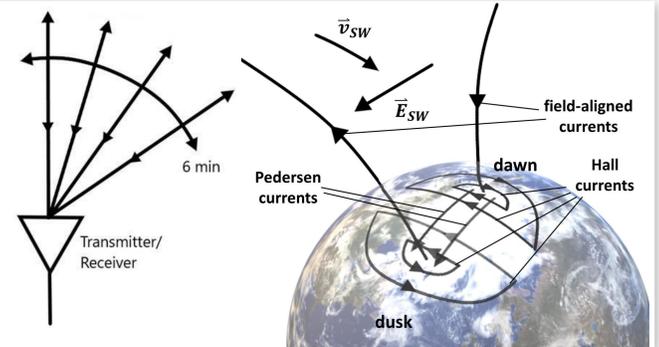
## Dynamo Region Joule Heating and EISCAT Beam-Swing Measurements

### Ionospheric Dynamo Region

- transition from **highly-collisional plasma**  $\nu_{xn} \gg \Omega_x$  to **collision-less plasma**  $\nu_{xn} \ll \Omega_x$
- electrons de-couple at  $\sim 80$  km altitude and ions de-couple at  $\sim 120$  km altitude
- Pedersen and Hall conductivity maxima  $\rightarrow$  "dynamo region"
- **Joule heating** by Pedersen currents:  $q_J = \sigma_P \cdot (\vec{E} + \vec{u} \times \vec{B})^2$

### EISCAT UHF beam-swing measurements

- EISCAT UHF antenna is rotated through a cycle of **4 pointing directions in 6 min**
- **line-of-sight ion velocity** is measured at each position
- **stochastic inversion** method is applied to obtain 3D ion velocity vector (Nygrén et al., 2011)
- **electric field** is calculated at collision-less F region:  $\vec{E} = -\vec{v}_F \times \vec{B}$



## Joule Heating Scaling Factor in TIE-GCM ( $q_J = f \cdot q_{J,m}$ default: $f = 1.5$ )

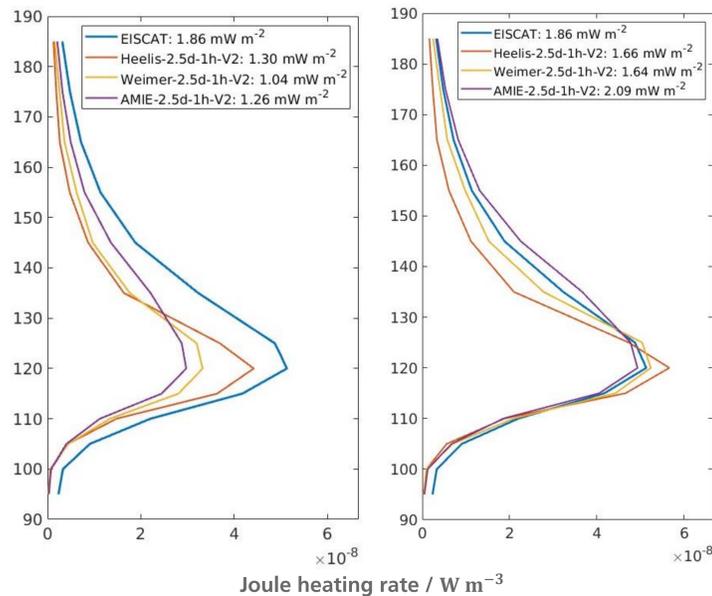
September 09-16, 2005 ( $Kp > 2$ ;  $\sim 180$ hr):

$\bar{f} = 1.505$

$f_H = 1.28$

$f_A = 1.66$

$f_W = 1.58$



- convection models applied for high-latitude plasma potential: **Heelis** (Heelis et al., 1982)  $\rightarrow$  empirical, **Weimer** (Weimer, 2005)  $\rightarrow$  empirical, **AMIE** (Richmond and Kamide, 1988)  $\rightarrow$  assimilative

- **short-scale variability** of electric fields increases  $q_J$  (Codrescu et al., 1995)

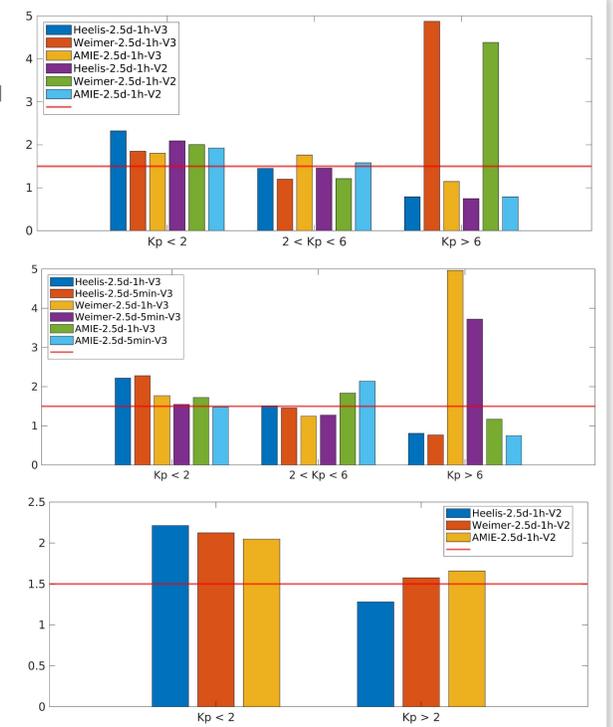
$$E = e_m + x \cdot e_v \quad e_v \sim 1.5 \cdot e_m \quad f(x) = \frac{1}{2} \cdot \theta(|x| - 1)$$

$$q_J \propto E^2 = \int (e_m + x \cdot e_v)^2 \cdot f(x) dx = 1.5 \cdot E^2$$

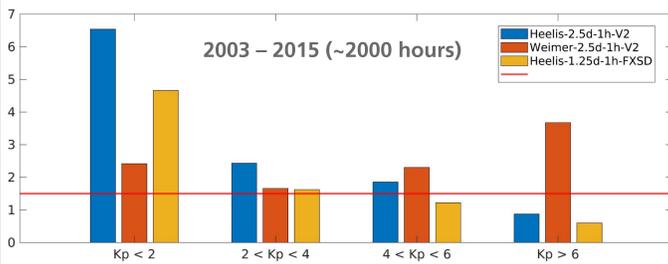
Günzkofer et al., 2024:

**default scaling very accurate on average**, but does not consider variations with **geomagnetic activity**, **magnetic local time**, or **plasma convection parameterization**

- Joule heating rates compared for **Heelis**-, **Weimer**-, and **AMIE**-driven model runs (AMGeO to be added...)
- variations of scaling factor  $f$  investigated with respect to:
  - geomagnetic activity (**Kp index**)
  - TIE-GCM version (**Version 2 and Version 3**)
  - model resolution (**2.5° and 1.25°**; not shown)
  - output sampling frequency (**1 hour and 5 min**)

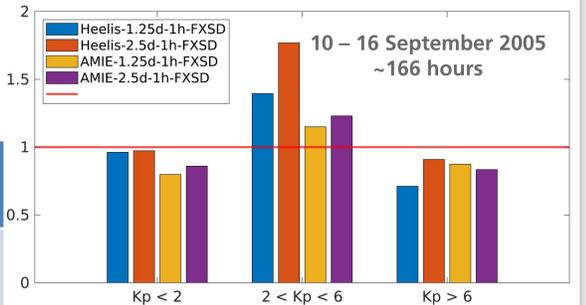


## Joule Heating in (SD) WACCM-X



- (SD) WACCM-X (Heelis) Joule heating rates are larger than TIE-GCM (Heelis) rates **by a factor of 1.4 – 1.55**
- WACCM-X **does not apply** a scaling factor ( $f = 1$ )
- (SD) WACCM-X Joule heating rates also **very accurate on average**

	Heelis 1.25°	Heelis 2.5°	AMIE 1.25°	AMIE 2.5°
$\bar{f}$	1.02	1.22	0.94	0.98



## Conclusions

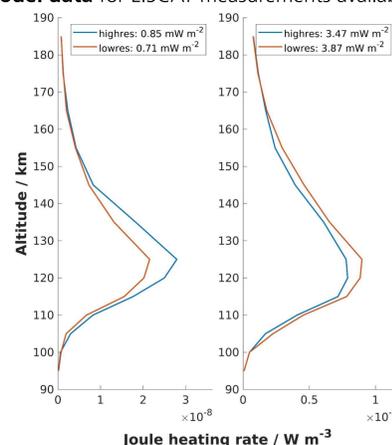
1. Joule heating rates in both TIE-GCM and (SD) WACCM-X agree very well with EISCAT measurements on average.
2. Additionally to the variations described in Günzkofer et al., 2024, model accuracy also varies with model version, resolution, and output sampling frequency.
3. Different convection models are the main source of differences.
4. Both maximum Joule heating rate and shape of profile are significantly different for different convection models.

## References:

Liu et al., *J. Adv. Mod. Earth Syst.*, **16**, e2023MS004024, 2024  
Günzkofer et al., *Earth Space Sci.*, **11**, e2023EA003447, 2024  
Nygrén et al., *J. Geophys. Res.*, **116**, A05305, 2011  
Weimer, *J. Geophys. Res.*, **110**, A05306, 2005  
Lyon et al., *J. Atmos. Sol.-Terr. Phys.*, **66**, 1333-1350, 2004  
Codrescu et al., *Geophys. Res. Lett.*, **22**, 2393-2396, 1995  
Richmond and Kamide, *J. Geophys. Res.*, **93**, A6, 5741-5759, 1988  
Heelis et al., *J. Geophys. Res.*, **87**, A8, 6339-6345, 1982

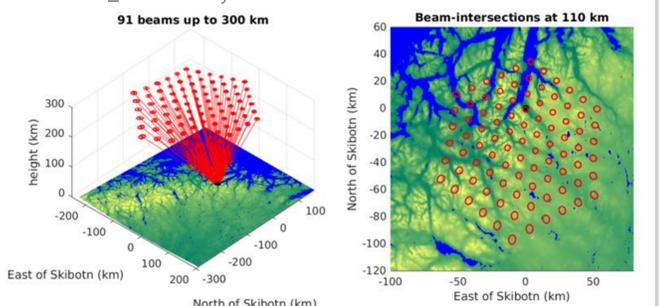
## Outlook – (HR) WACCM-X

- Liu et al, 2024 performed **0.25° and 1.25°** resolution runs with WACCM-X spectral element (SE) dynamical core
- high-latitude electrodynamics forced with **GAMERA** (Lyon et al., 2004)
- **no model data** for EISCAT measurements available so far



## Outlook – PFISR and EISCAT\_3D

- **phased array ISRs** allow for electronic beam steering
- significantly faster cycle times and **higher temporal resolution** of ion velocity measurements
- **Poker Flat ISR (PFISR)** is a phased array radar on approximately the same **geomagnetic latitude** as EISCAT
- test scaling factors for **longitudinal variations**  $\rightarrow$  scaling factors only applicable if **constant** at same magnetic latitude
- **EISCAT\_3D** currently under construction



from "Recommendation for initial stage of EISCAT\_3D common programmes and their future directions" by the EISCAT\_3D Common Programme working group

## Acknowledgments:

EISCAT is an international association supported by research organizations in China (CRIRP), Finland (SA), Japan (NIPR and ISEE), Norway (NFR), Sweden (VR), and the United Kingdom (UKRI). The TIE-GCM and related Thermosphere-Ionosphere models have been developed by the "Atmosphere Ionosphere Magnetosphere" (AIM) Section of the High Altitude Observatory (HAO) at NCAR.

## Contact: Florian Günzkofer

Institute for Solar-Terrestrial Physics, DLR  
Kalkhorstweg 53, 17235 Neustrelitz  
Tel.: +49 3981 480-118, Email: florian.guenzkofer@dlr.de