Polarimeter to Unify the Corona and Heliosphere



PUNCH 6 Science Meeting February 25-26, 2025 Cal Poly

WFI Instrument Status

Glenn Laurent WFI Instrument Lead









WFI Overview

- WFI/NFI provide first complete, photometric, high resolution views of corona/solar wind transition.
 - WFI 5-45°, NFI 1.5-8°
- WFI provides first wide-field polarimetric solar wind images.
- Design based on STEREO/HI, SoloHI heliospheric imagers.
- 3 observatories in 620 km polar orbit (95.95 min)
- Rotating trefoil pattern orbit separated by 120° ±30°.
 - Continuous observations 4 min observing cadence (2x per roll)
 - Full coverage in 32 min
 - 30° roll every 8 min



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| Resource | СВЕ | Cont. | Total | | |
|--------------------|-------|--------|-------|--|--|
| Mass (kg) | 16.88 | 7.05% | 18.07 | | |
| Power (W) | 15.49 | 12.00% | 17.35 | | |
| Length (mm) | 889 | - | 889 | | |
| Width (mm) | 438 | - | 438 | | |
| Height (mm) | 149 | - | 149 | | |
| Data Rate (GB/day) | 1.41 | 34.20% | | | |

* LV update provides additional margin

WFI Instrument



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PUDGH

Integrated Instrument (WFI-3)





WFI Block Diagram



PUDGH

WFI Level 2 Driving Requirements

• All Requirements validated

| ID | Requirement | Value | Performance | Status |
|------|---------------------|--|-------------------------------------|--------|
| 1057 | Passband | Width: 300+/-100nm Center: 550+/-75nm | 450-750nm | PASS |
| 1063 | Field of View (FOV) | 20 R $_{\odot}$ – 160 R $_{\odot}$ | 17.4 R $_{\odot}$ – 180R $_{\odot}$ | PASS |
| 1064 | Instantaneous FOV | 40 deg [°] square truncated by 50 deg [°] circle | 40.2° FOV Baffle, >50° OLA FOV | PASS |
| 1068 | Angular Resolution | 4 arcmin | 2.4 arcmin | PASS |
| 1071 | Norm. Sensitivity | 7E-17 B _☉ | 3.7E-17 B $_{\odot}$ | PASS |
| 1076 | Polarization | 3 angles | -60°, 0°, +60° | PASS |

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WFI Conops

PUNA

| | | | PUNCH Observing Sequence Schedule | | | | | | |
|-----|--|--------------|-----------------------------------|-----------------------------------|---------------------------------|--|--|--|--|
| | Conops common to WFI & NFI | Time (s) | Length + margin | NFI Action | WFI Action | | | | |
| | Two sate of polarization coguanase por 9 min | 0 | 47+4 | Roll & set PFW to -60 $^{\circ}$ | Roll & set PFW to -60° | | | | |
| | • Two sets of polarization sequences per o min | 51 | 48+1(*) | Expose 3x13s at -60° | Expose 45s at -60° | | | | |
| | roll cadence | 98 | 15+5 | Set PFW to 0° & settle | Set PFW to 0° & settle | | | | |
| | 20 seconds PFW rotation time | 118 | 48+1(*) | Expose 3x13s at 0° | Expose 45s at 0° | | | | |
| | | 165 | 15+5 | Set PFW to 60° & settle | Set PFW to 60° & settle | | | | |
| | | 185 | 48+1(*) | Expose 3x13s at 60° | Expose 45s at 60° | | | | |
| | | 232 | 15+5 | Set PFW to CL & settle | Set PFW to CL & settle | | | | |
| | | 252 | 22+1(*) | Expose 3x5s at CL | Expose 19s at CL | | | | |
| | | 273 | 15+5 | Set PFW to -60° & settle | Set PFW to -60° & settle | | | | |
| | ◄ Roll Cadence: 480s ±2s (8 min.) | 293 | 48+1(*) | Expose $3x13s$ at -60° | Expose 45s at -60° | | | | |
| | | 340 | 15+5 | Sep PFW to 0° & settle | Sep PFW to 0° & settle | | | | |
| WFI | -60° -0° +60° Clear -60° -0° +60° | 360 | 48+1(*) | Expose 3x13s at 0° | Expose 45s at 0° | | | | |
| NFI | -60° -0° +60° Clear -60° -0° +60° | 407 | 15+5 | Set PFW to 60° & settle | Set PFW to 60° & settle | | | | |
| (| 0 100 200 300 400 | 427 | 48+1(*) | Expose 3x13s at 60° | Expose 45s at 60° | | | | |
| 1 | Elapsed Time (sec) | - 474 | 1 to 11 | Sync for next roll | Sync for next roll | | | | |
| ļ | Legence Maneuver Exposure Wheel Motion Margin | (*) 2- | second ove | rlap with following event | | | | | |
| | | | | | | | | | |

WFI Instrument Status

| Milestone | WFI-1 | WFI-2 | WFI-3 | | |
|------------------------|--------------------------|-------------------------|-------------------------|--|--|
| Camera Focus | PASS | PASS | PASS | | |
| Vibe | PASS | PASS | PASS | | |
| TVAC / TBAL | PASS | PASS | PASS | | |
| SCOTCH | PASS | N/A (Descoped) | N/A (Descoped) | | |
| Optical Performance | PASS | PASS | PASS | | |
| PSR / EIDP | Complete (10/24/2023) | Complete (3/8/2024) | Complete (4/12/2024) | | |
| Delivery | Complete (1/22/2023) | Complete (3/22/2024) | Complete 5/14/2024 | | |

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WFI Requirements Verification

WFI Requirements Verification Summary Table

| Level | Total Requirements | Passed | Deferred | Waiver | Open | Percent Verified Items |
|-----------------|-----------------------|--------|----------|--------|------|---------------------------|
| WFI Level 3 / 4 | 113 | 110 | 1 | 2 | 0 | 100% |
| WFI Level 5 | 20 | 20 | 0 | 0 | 0 | 100% |
| Total | 133 | 130 | 1 | 2 | 0 | |

- Two Waivers (Metering Bracket Reflectivity, Solar Shield Position) approved -- Negligible performance impact
- No Outstanding MIUL Open Items
- No MUAs

IDdi

WFI Structural Verification (Vibration Testing)



WFI-3 Vibration Testing Completed

WFI Thermal Verification (TVAC / TBAL)



WFI-FM003 TVAC / TBAL Complete (meeting ERD requirements)

WFI Optical Focus Testing



WFI FM1 5x5_231011



WFI-1 Optical Performance testing (all filters) Meets requirements

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Point Source Regularization







PSF is well understood and Software Pipeline is ready to create Mosaics

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Incl

TVAC / SOC Calibration -- WFI



Photon Transfer Curve



SOC De-streaking Analysis Validated on Cold Operational Data

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WFI Instrument/Observatory Focus Repeatability



Radial distance [deg]

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TIDAT

WFI Optical Vignetting Testing



WFI-1 Flat Fields meet Vignetting Requirements

IIDGH

WFI Observatory Vignetting Model



WFI Vignetting Model Ready for PUNCH Software Pipleline

Inch

WFI Post-Vibe LED Testing



WFI-3 LED Testing completed (Meets science requirements)

PLIDGH

WFI Post-Vibe LED/PFW Testing





WFI-3 LED Testing completed (Verifies minimal CCD Obscured pixels)

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Indi

WFI-3 Polarizer Testing completed (Verifies Polarizer Angles)



NFI/WFI Observatory TVAC Bias/Read Noise Trending

- Observatory TVAC testing examined Bias & Read Noise
- Cold CCD, S/C Systems Operating (Radio link, reaction wheel, torque rod active)
- Bias values are consistent with nominal CCD temperature fluctuations (< +/- 0.5 DN)
- Read Noise trends are small (< +/- 0.02 DN)
- Both HotOp & ColdOp show similar trends

Bias & Read Noise Stable through end-to-end Observatory Testing (Verifies CCD EMI/EMC Performance)

WFI01 NFI **WFI-1 WFI-2** Hot-on m WFI01 WFI-3

NFI/WFI Observatory TVAC Dark Trending

- Observatory TVAC testing examined
 Dark current / Noise
- Cold CCD, S/C Systems Operating (Radio link, reaction wheel, torque rod active)
- Dark Current of ~0.7-1.3 e-/pix/s meets requirement of <5 e-/pix/s.
- Dark Current fluctuations are consistent with nominal CEB temperature fluctuations (~ +/- 0.1 e-/pix/s)
- Dark Current rms trends are small (< +/-0.01 e-/pix/s)

Dark Current Stable through end-to-end Observatory Testing (Verifies CCD EMI/EMC Performance)



WFI Delivery to Observatory

- All WFIs Successfully Delivered to Observatory I&T
 - WFI-1 (1/22/2023)
 - WFI-2 (3/22/2023)
 - WFI-3 (5/14/2023)
- Functional Testing
- Comprehensive Performance Testing (CPT)
 - Bias/Read/Dark Noise
 - LED / Flat Field
 - PFW
 - Door Testing
 - Optical Focus Testing





PUNCH I&T @ VSFB

- Vandenberg SFB I&T
 - Delivery & Inspection





PUNCH I&T @ VSFB

- Vandenberg SFB I&T
- Functional Testing
- Comprehensive Performance Testing (CPT)
 - Bias/Read Noise
 - LED / Flat Field
 - PFW
 - Door Testing



PUNCH I&T @ VSFB

- Vandenberg SFB I&T
- Functional Testing
- Comprehensive Performance Testing (CPT)
 - Bias/Read Noise
 - LED / Flat Field
 - PFW
 - Door Testing











• WFI-2 Solar Array Deployment



Commissioning / Calibration

- On-orbit calibration of WFI & NFI Instruments individually
- On-orbit commissioning/validation of the "single virtual observatory" cross-instrument calibration and SOC testing

• Commissioning phases:

Spacecraft Commissioning:L to L+30days(includes WFI/NFI functional checks)Instrument Commissioning:L+30days to L+60days(instrument calibration)Constellation Commissioning:L+60days to L+90days(validate "virtual" instrument)

- Instrument commissioning plan well defined
 - Identifies ADCS/Camera/PFW Commanding, Imaging types, Required Polarizers

Closed-Door Activities

- Thermal Verification (PRTs, Heaters, Heater Control)
- Camera Verification (Bias, Darks, LED imaging, Linearity)
 - Continue against trending of noise from TVAC
- PFW Verification (motion, resolver function)



LED Composite Flat Field: similar for both NFI and WFI since the LEDS are integrated into the camera housing

Open Door Activities

- Starfield Imaging validates several instrument parameters
 - WFI Baffle alignment to S/C
 - Offset measured between predicted & measured pointing
 - Update boresight pointing in ADCS
 - PSF performance
 - Analyze stars in field of view, compare to predictions from test data and update PSF mapping for each instrument
 - Platescale, FOV, Distortion
 - Perform astrometry fit to determine optical parameters
 - Done previously on EM WFI night sky observations
 - Photometric performance
 - From starfield, determine CCD response / vignetting function as function of position.
 - Repeat imaging over a minimum of two weeks to ensure stars have drifted over the field of view.









EM WFI Starfield astrometry & distortion

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Open Door Activities

- Multiple images of the star field provide
 - Baffle solar attenuation performance
 - Nominal pointing
 - Starfield photometry provides attenuation performance
 - Optimize WFI performance by adjusting WFI pitch angle
 - Polarizer efficiency

That

- Image polarized source at each PFW filter position
- Compare stellar photometry over an orbit, (including observatory rolls) to measure PFW polarizer efficiencies



WFI Vignetting Functions

Performing full exposure sequence in preparation for Phase E

Optimize exposure times / summing



Figure shows Solar Orbiter Metis polarization calibration. PUNCH will preform a full roll during every orbit.

| Polarizer | Axis | Mapp | ing (D | egree | es) | | | | | | | | | | |
|-----------|------|-------|--------|-------|-------|-------|-------|-------|-------|-------|-------|-------|-------|-------|------|
| x,y in mm | -21 | -18 | -15 | -12 | -9 | -6 | -3 | 0 | 3 | 6 | 9 | 12 | 15 | 18 | 21 |
| 21 | | | | | | | | 0.11 | | | | | | | |
| 18 | | | | | -0.01 | -0.01 | 0.21 | 0.09 | 0.03 | 0.06 | 0.03 | | | | |
| 15 | | | | 0.60 | -0.06 | 0.03 | 0.00 | 0.02 | -0.01 | 0.07 | -0.01 | 0.04 | | | |
| 12 | | | 0.01 | -0.02 | 0.00 | -0.01 | 0.06 | 0.09 | 0.05 | 0.05 | 0.01 | -0.01 | 0.03 | ÷ | |
| 9 | | 0.00 | 0.00 | 0.06 | -0.03 | 0.02 | 0.05 | 0.01 | -0.01 | -0.02 | -0.02 | 0.02 | -0.01 | 0.02 | |
| 6 | | 0.01 | -0.01 | -0.03 | 0.04 | 0.04 | 0.03 | 0.07 | 0.04 | 0.02 | 0.00 | -0.01 | -0.01 | 0.05 | |
| 3 | | -0.03 | 0.06 | 0.01 | -0.04 | 0.03 | 0.01 | 0.03 | 0.07 | -0.02 | 0.05 | -0.04 | -0.02 | 0.05 | |
| 0 | 0.00 | 0.09 | 0.00 | 0.02 | 0.03 | 0.01 | 0.13 | 0.00 | 0.00 | -0.02 | 0.01 | 0.00 | -0.08 | -0.01 | 0.05 |
| -3 | | -0.04 | 0.01 | 0.06 | -0.05 | 0.05 | 0.02 | -0.01 | 0.02 | 0.00 | -0.01 | 0.01 | -0.03 | 0.00 | 1 |
| -6 | | 0.01 | -0.01 | -0.03 | 0.04 | 0.04 | 0.03 | 0.07 | 0.04 | 0.02 | 0.00 | -0.01 | -0.01 | 0.05 | 0 |
| -9 | | 0.06 | 0.02 | 0.01 | 1.61 | -0.03 | 0.36 | 0.01 | 0.06 | -0.01 | -0.22 | 0.10 | 0.33 | -0.02 | |
| -12 | | | 0.03 | -0.26 | -0.40 | 0.25 | 0.02 | 0.09 | -0.05 | -0.05 | 0.04 | -0.01 | 0.04 | 6 | |
| -15 | | | | 0.08 | 0.02 | -0.02 | -0.04 | 0.02 | -0.04 | -0.02 | 0.05 | -0.18 | | Ŷ | |
| -18 | | | | | 0.00 | -0.01 | -0.01 | 0.05 | 0.00 | 0.03 | -0.02 | 0.21 | | | |
| -21 | | | | | | | | 0.21 | | | | | | 1 | |

Polarizer uniformity

Single Virtual Observatory Commissioning

- Each instrument is *individually* calibrated for
 - PSF
 - FOV/plate scale/distortion
 - Photometric performance (Occulter performance / vignetting function)
 - Polarimetry
- Constellation calibration ensures *relative* calibration across the cameras, by directly comparing simultaneous measurements in overlapping fields of view.
- The four principal calibration quantities are refined to produce higher quality relative calibration to ensure the cameras function as a single "virtual instrument" after SOC processing.

Single Virtual Observatory Commissioning

- All constellation calibrations are performed at the SOC level
- Uses same science data as used to calibrate the individual instruments





Single Virtual Observatory Commissioning

- Cross check and verify:
 - PSF correction
 - Tune L1 PSF calibration to ensure instruments are matched.
 - Photometric performance (instrument photometry and vignetting)
 - Verify standalone calibration curves
 - Adjust calibration to improve relative precision
 - Distortion calibration
 - Adjust derived instrument calibration functions to achieve required co-alignment across each field
 - Polarimetric calibration
 - Verify that known catalog high-polarization stars and solar wind features are measured the same, independent of source camera
- All FOVs overlap, providing cross-calibration "patches" between instruments for each cross-check









