



SUNCET

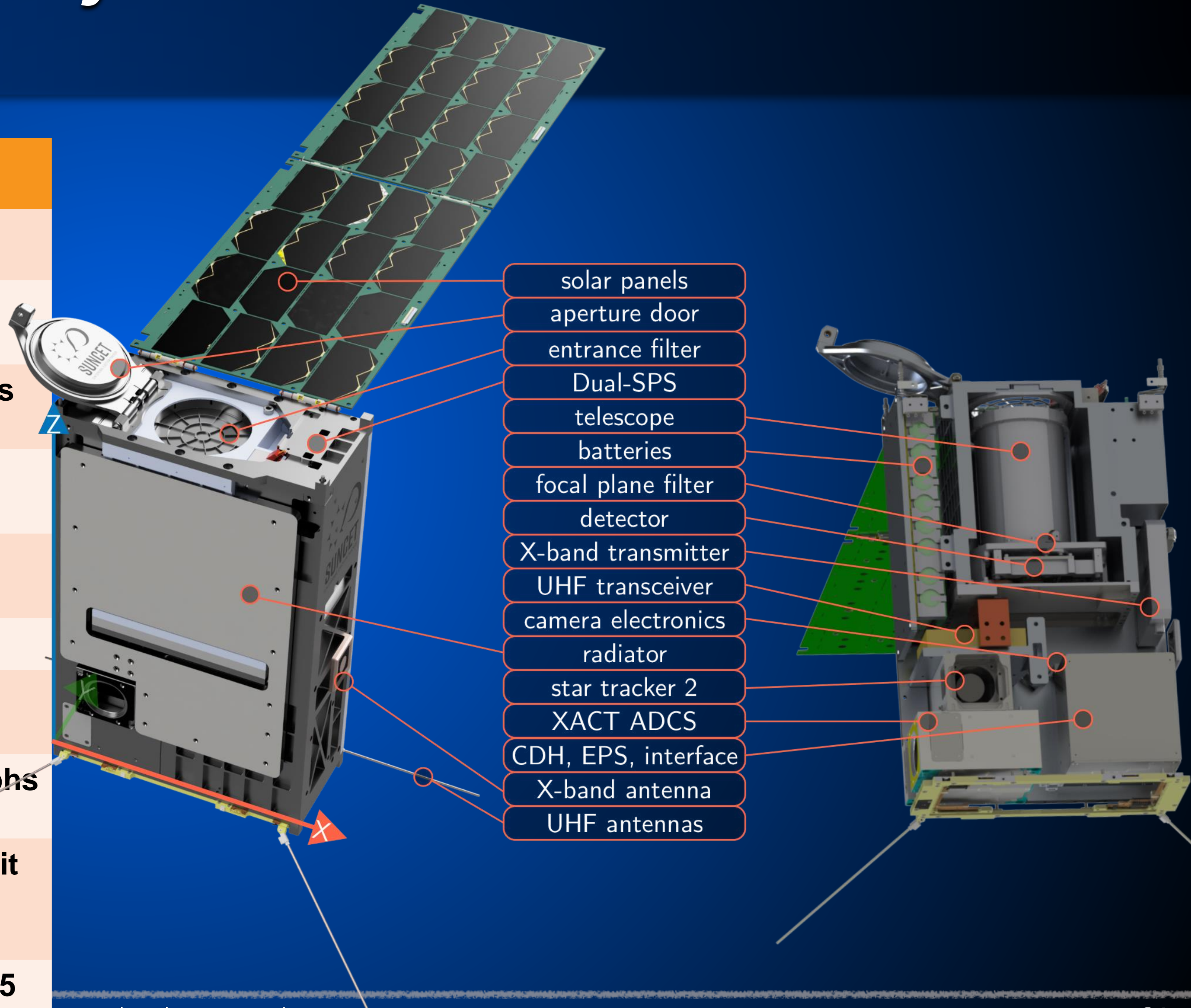
SUN CORONAL EJECTION TRACKER



Concise summary

NASA mission tech specs

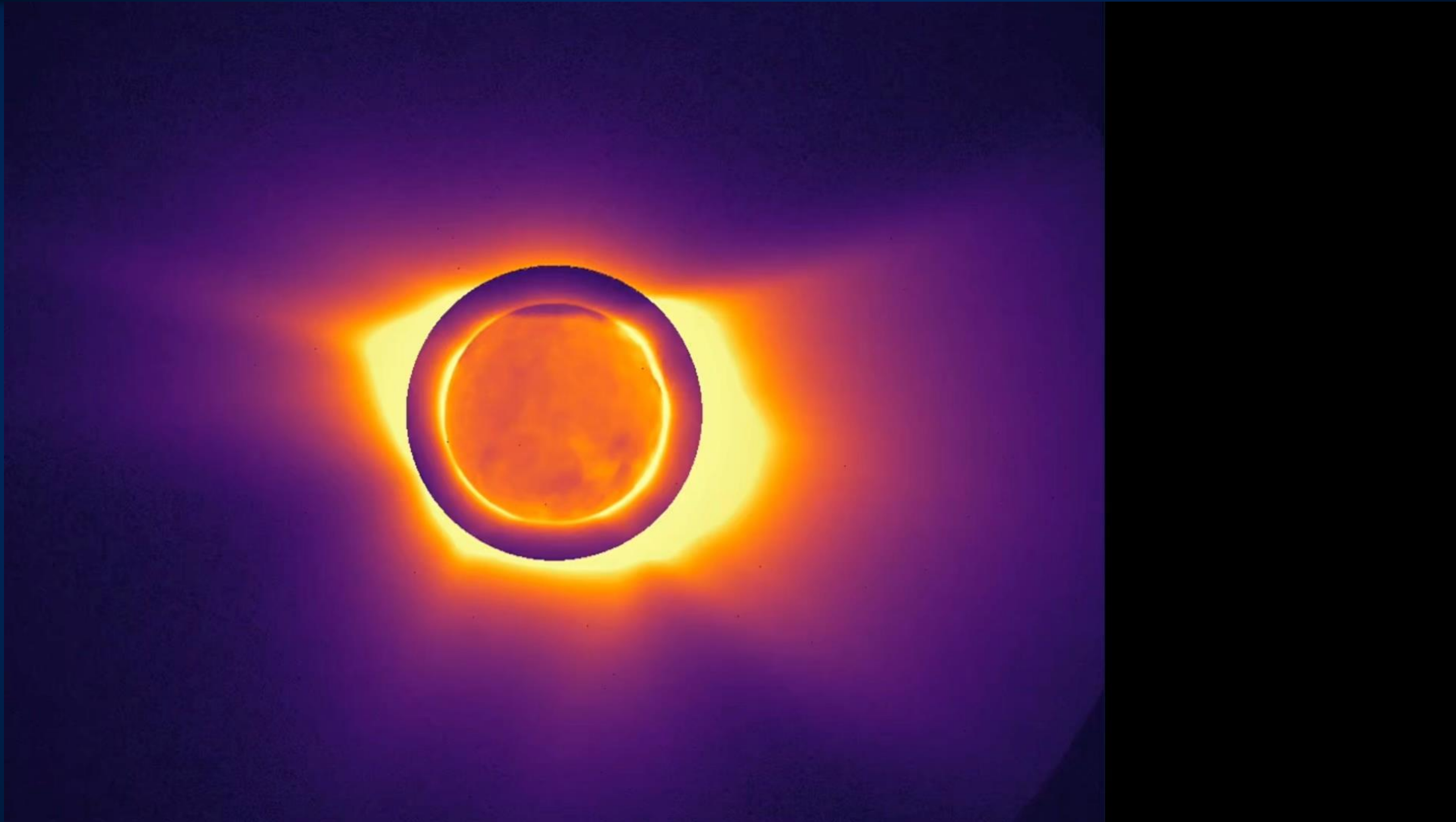
Parameter	Value	Note
Form factor	6U CubeSat	~Shoebox size
Launch date	Sometime 2026	Space X Falcon 9 TSIS-2 Rideshare
Prime mission	8 months	To extend as long as able
Orbit	Altitude $\leq \sim 550$ km inclination $\geq 35^\circ$	sun-sync 98° Inclination
FOV	$\pm 5.34 R_\odot \times \pm 4 R_\odot$	That's wide! ± 1.5 is typical
Bandpass	170-200 Å	Wider than typical
Dynamic range	2×10^6	SDO/AIA's is 10^4 (typical)
Spatial resolution	20 arcsec	Similar to coronagraphs
Exposure times	0.035 sec (on disk) 15 seconds (off disk)	Configurable on orbit
Cadence	1 minute	Nominally; can do 15 sec





Why SunCET is needed

Bulk of CME acceleration occurs in the “middle” corona (e.g., Bein et al., 2011; D’Huys et al., 2014)

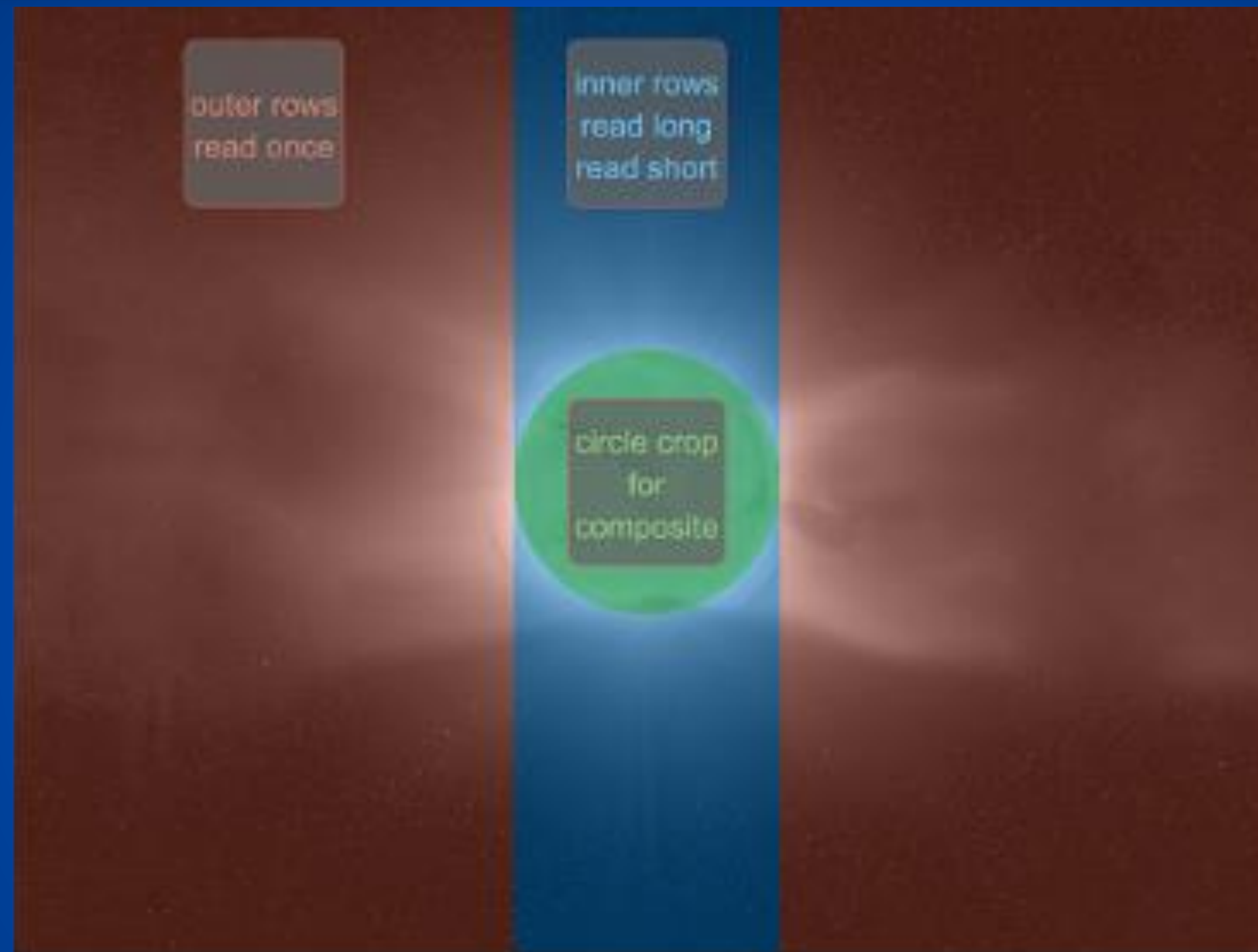


Mason+, ApJ, 924, 63 (2022)

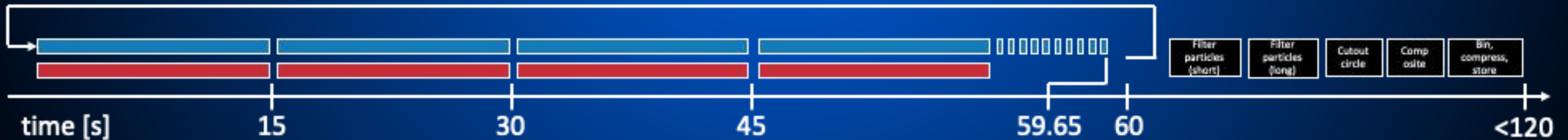


How SunCET handles $10^4:1$ disk:off-disk photons

A novel, simultaneous high dynamic range (SHDR) detector + algorithm



*image dimensions and colored regions to scale

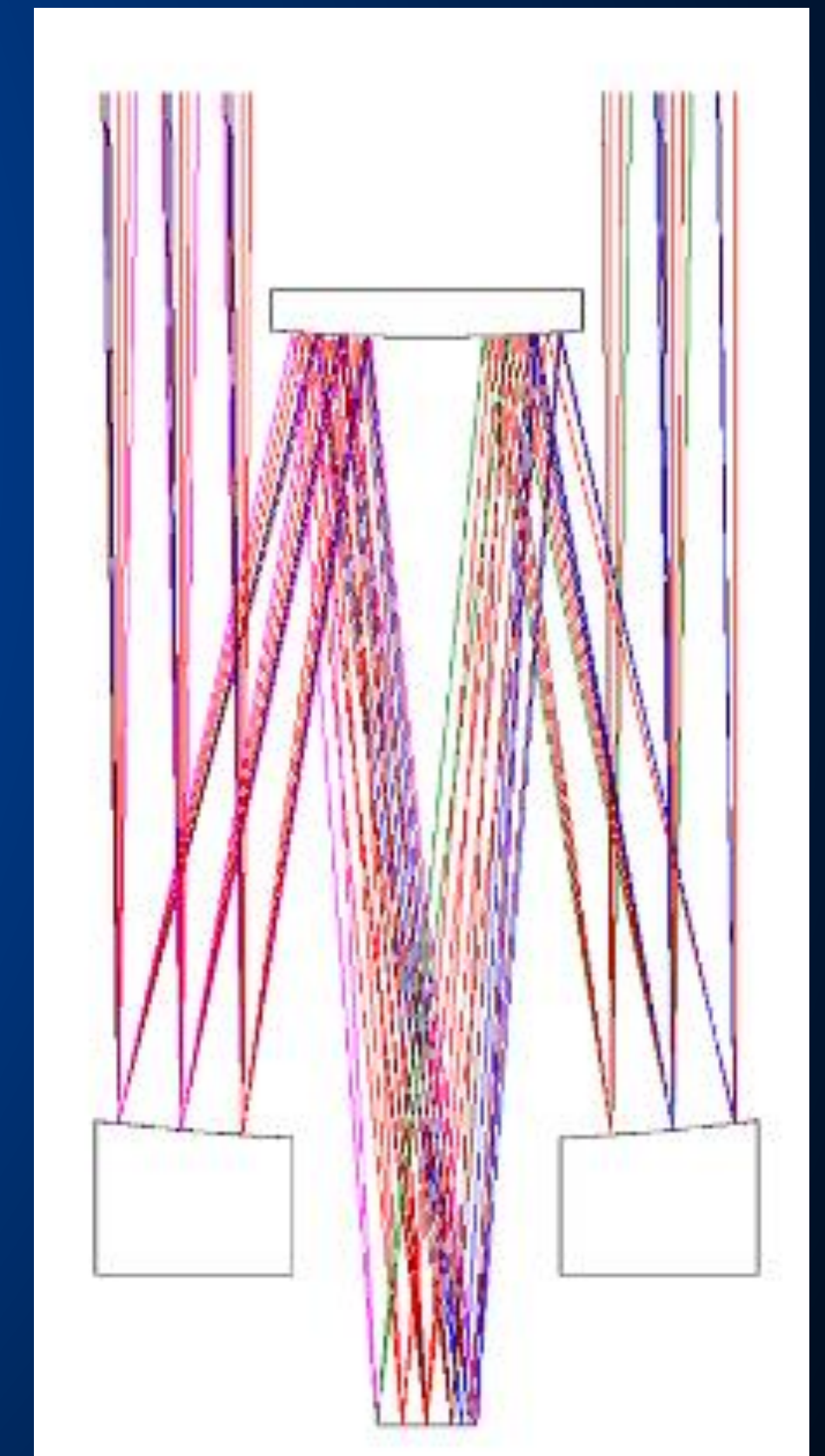
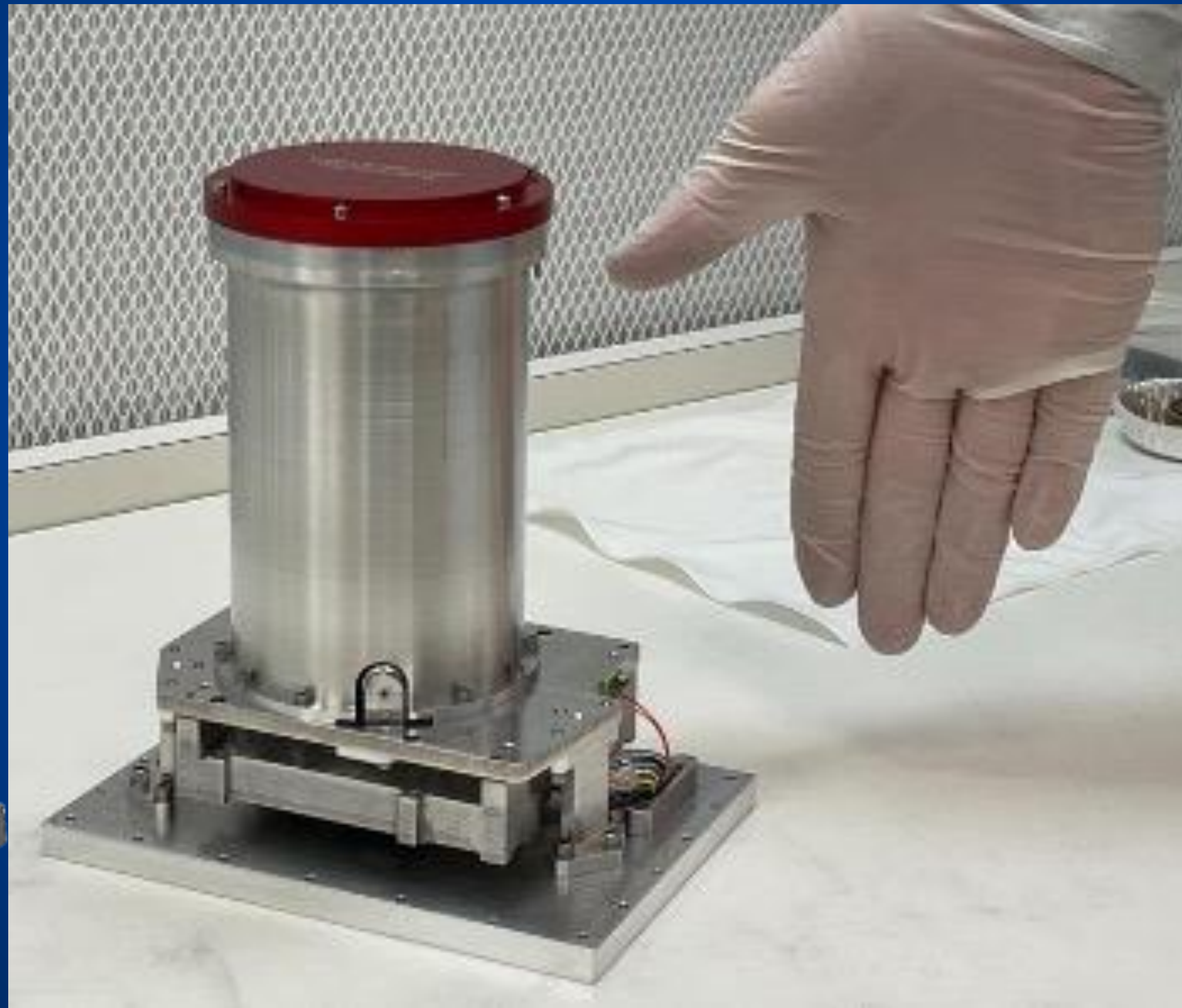
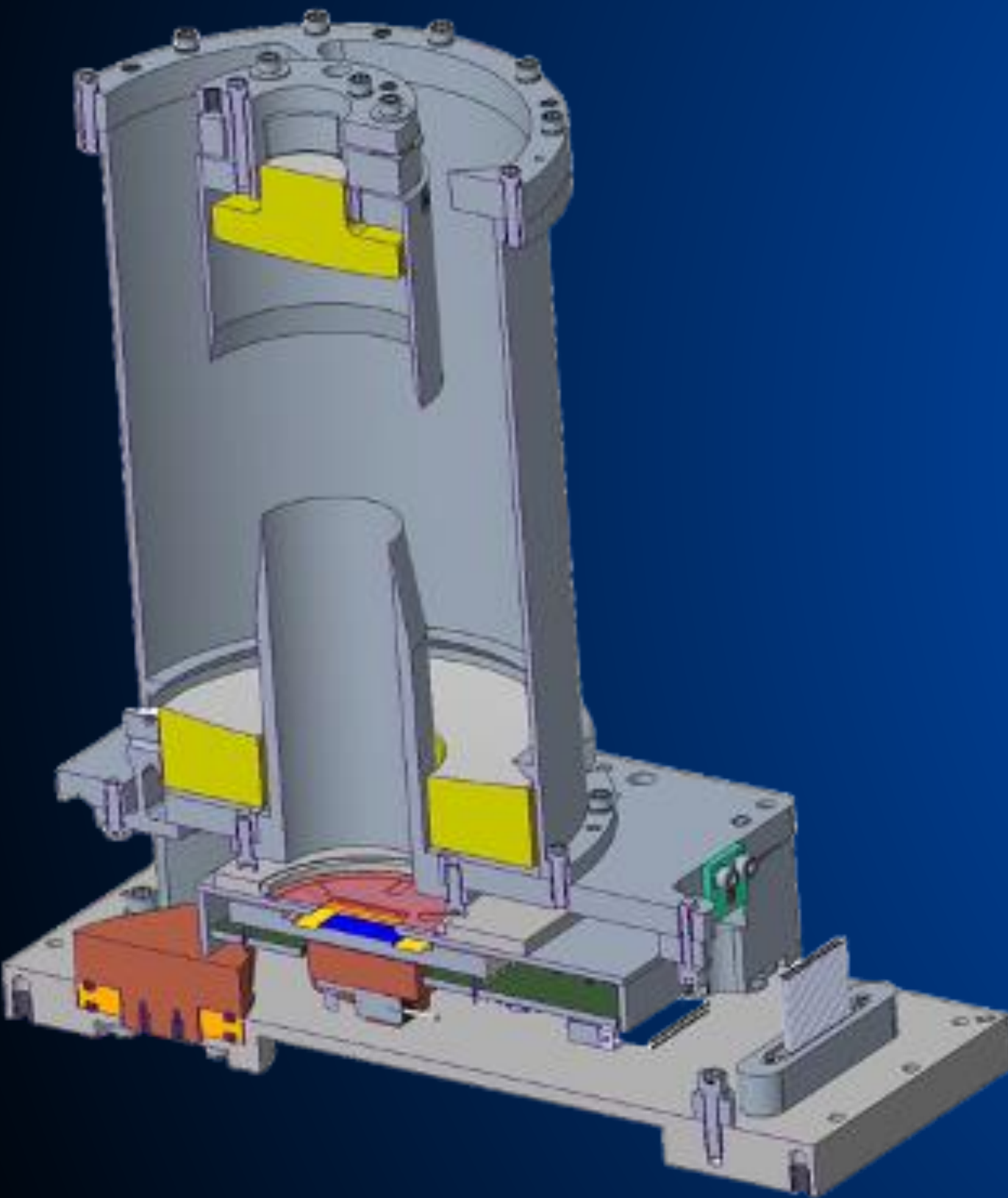


Sun image is MHD simulation run on NASA's Pleiades supercomputer of solar CME (dimmiest of the set we ran = worst case), then passed through our SunCET instrument simulator to create realistic synthetic SunCET image composites (right FOV, spatial/temporal/spectral resolution, QE, QY, reflections, etc)

Mason+, ApJ, 924, 63 (2022)

Standard Ritchey-Chrétien Telescope

Provides wide field of view, and can be made *tiny*

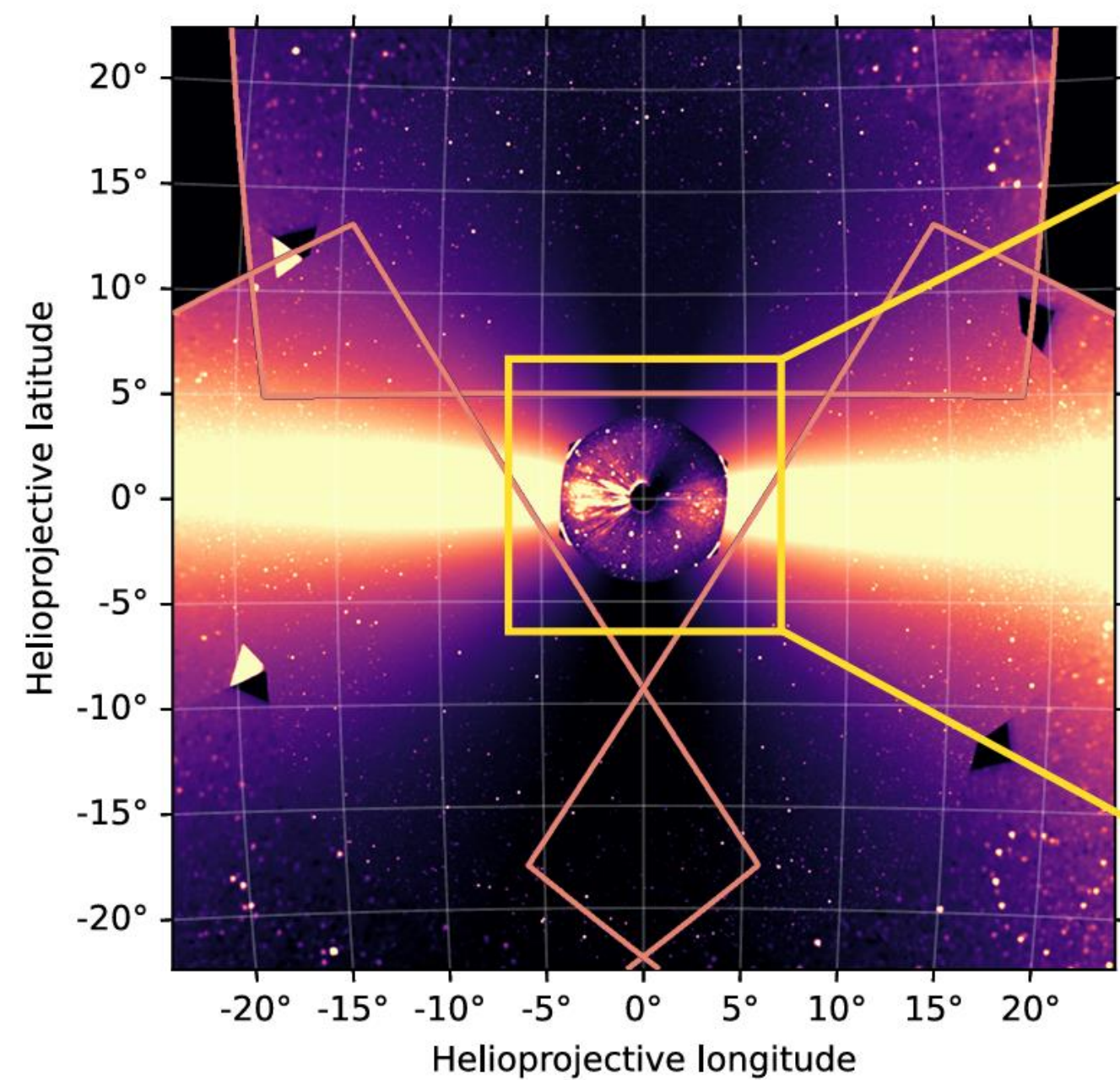




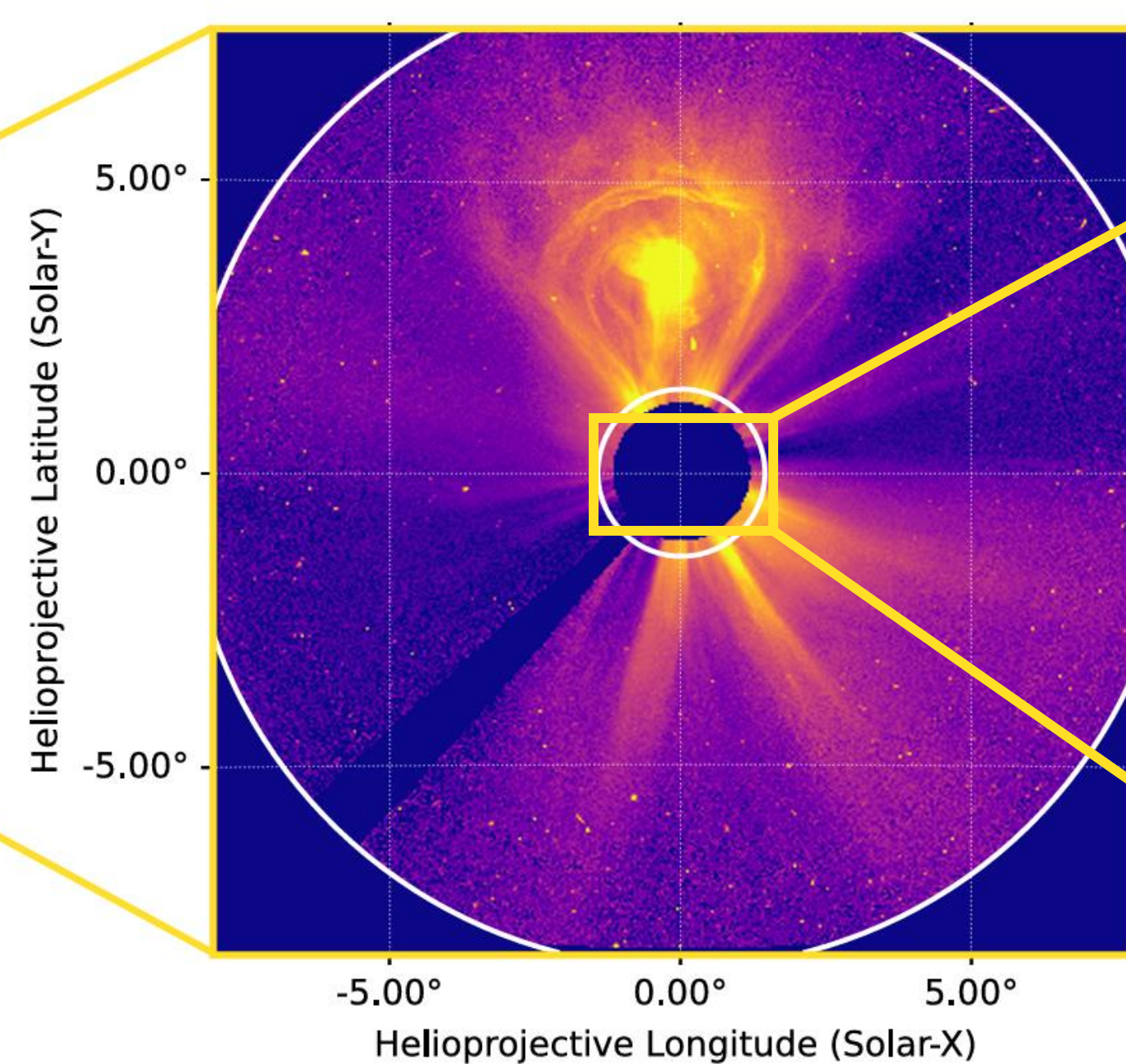
SunCET-PUNCH Overlap

Fields of view

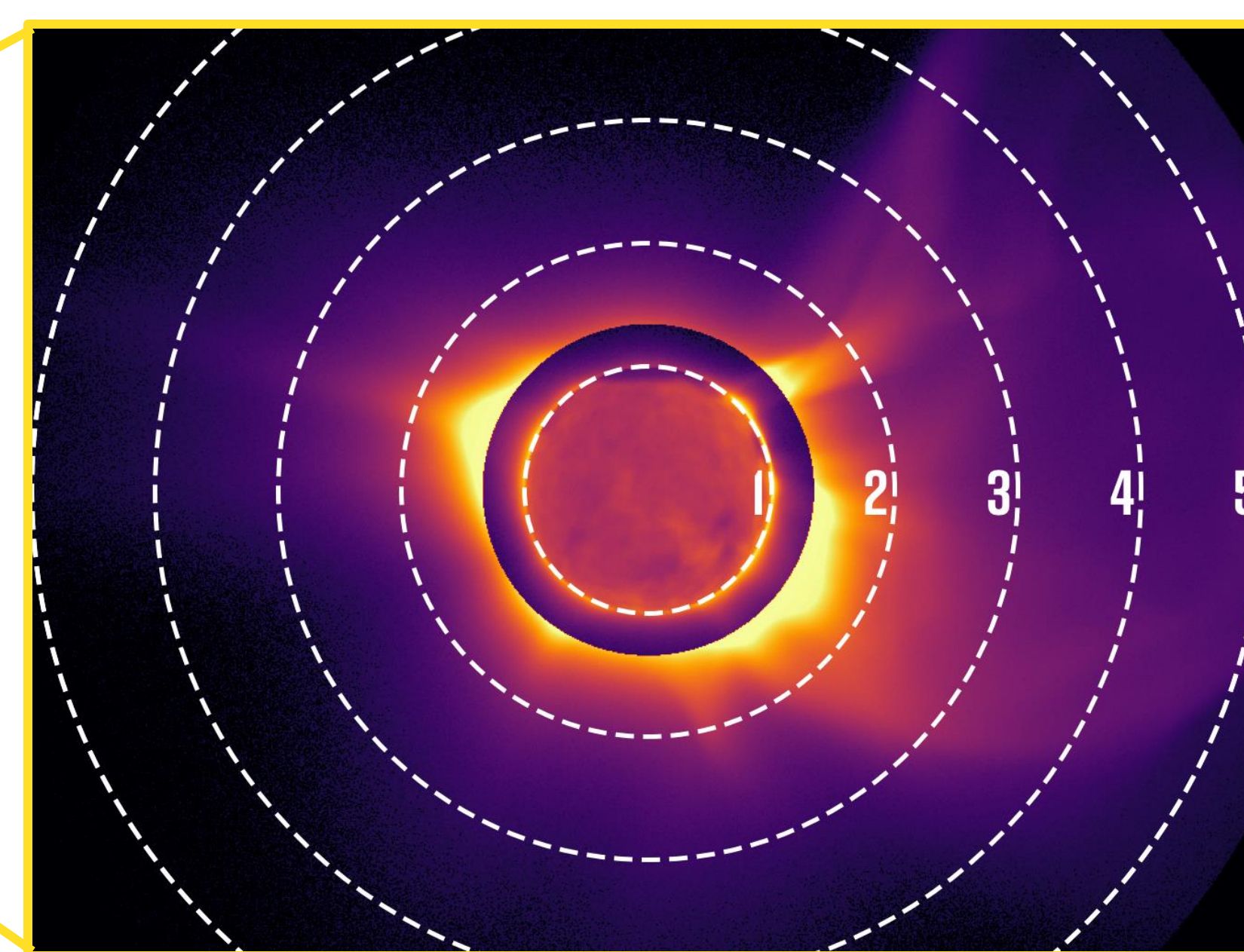
WFI



NFI



SunCET



$17.4 R_{\odot} - 180 R_{\odot}$

$5.4 - 32 R_{\odot}$

$\pm 5.34 R_{\odot} \times \pm 4 R_{\odot}$