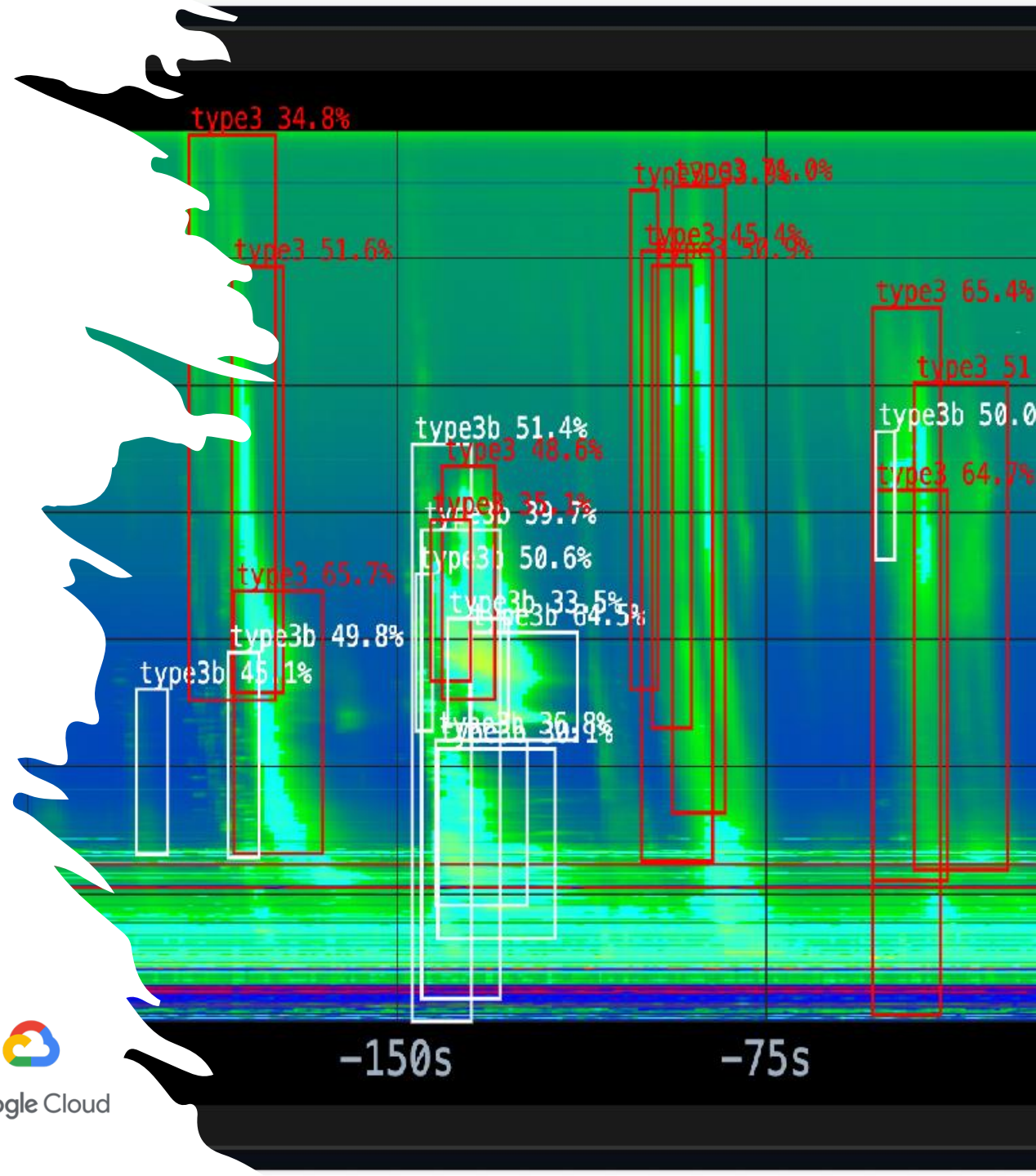


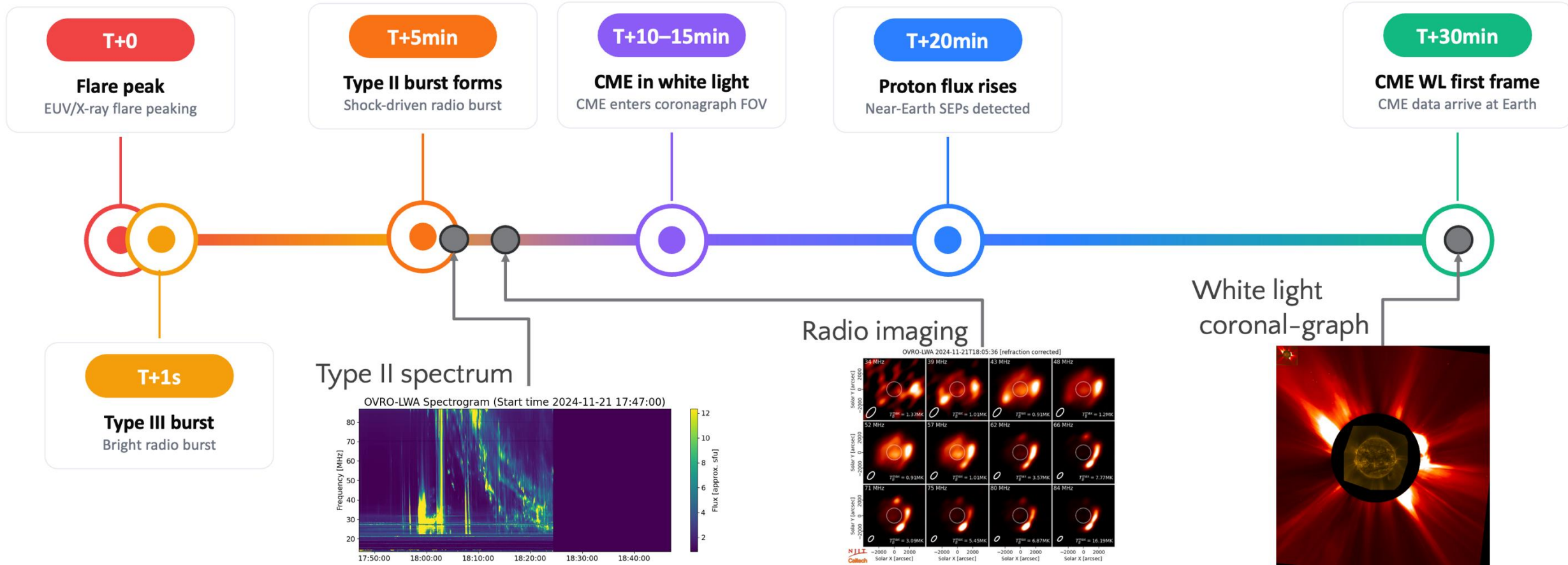
AI-Assisted Real-Time Monitoring and Reporting of Solar Radio Activity with OVRO-LWA

Peijin Zhang
and OVRO-LWA team



Motivations for realtime and AI assisted reporting

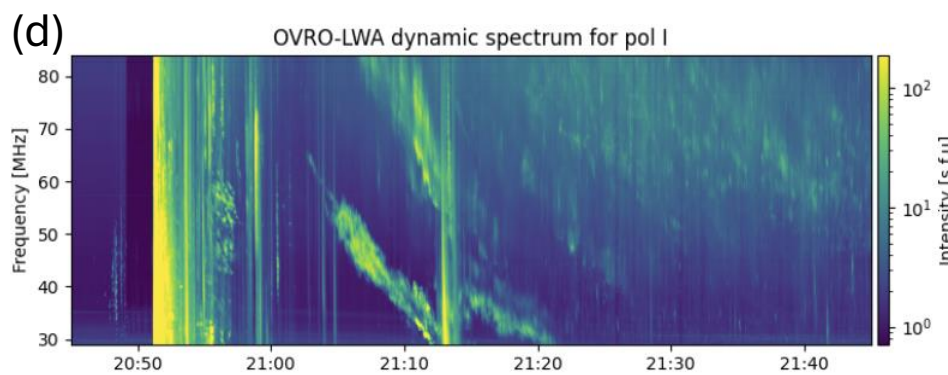
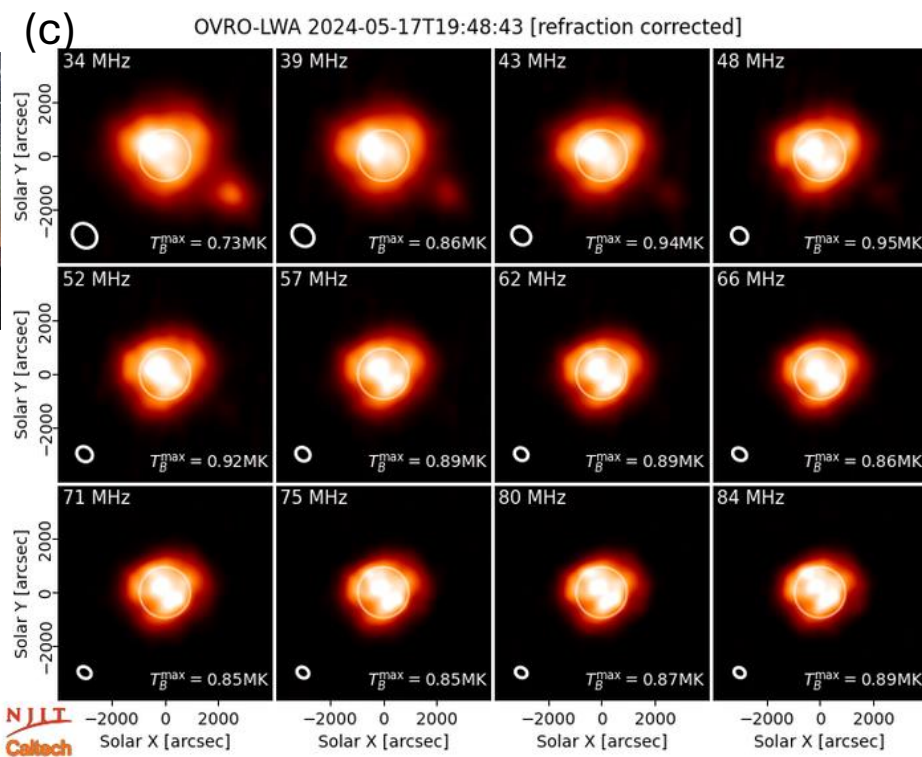
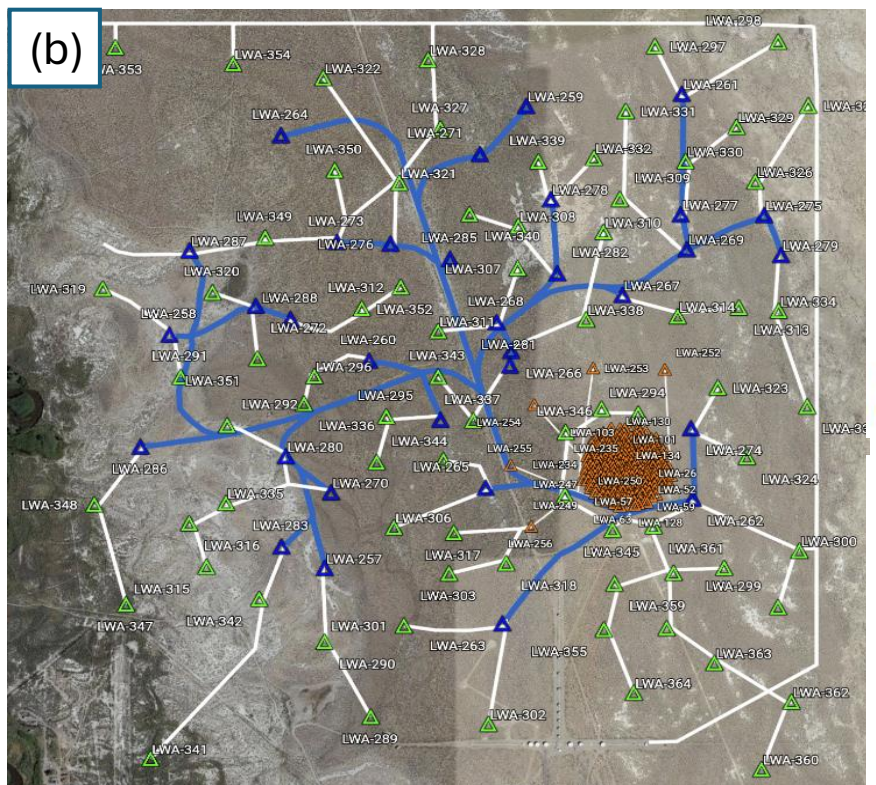
- Solar radio burst is a good indicator of the energetic particle accelerations solar activities.
- Timely reporting of radio burst type, brightness, morphology can help evaluate SEP of the event.



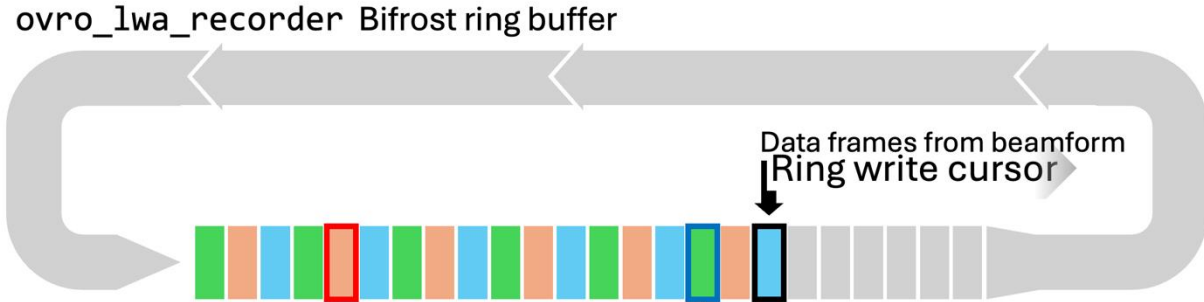
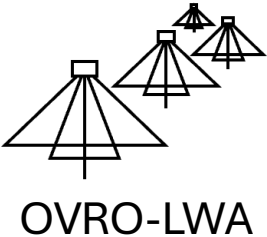
OVRO-LWA

352 Antennas (LWA dipole) -
Full array cross correlation
Layout: Dense core + Out-
riggers

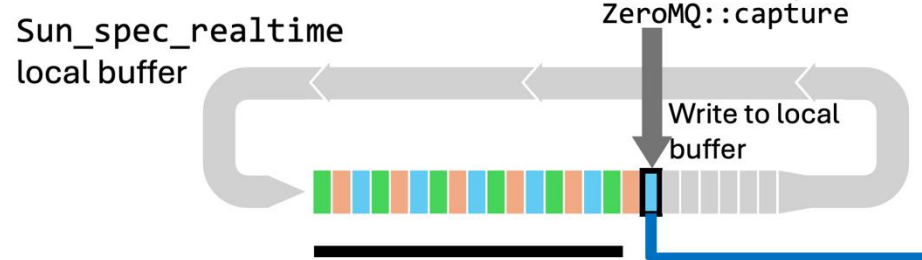
- Fast mode imaging:
Full array 10s cadence
- Slow mode imaging:
48 ant 10s cadence
- Dynamic spectrum:
256 ant beamforming
1ms, 24kHz resolution



Low latency data acquisition of the dynamic spectrum



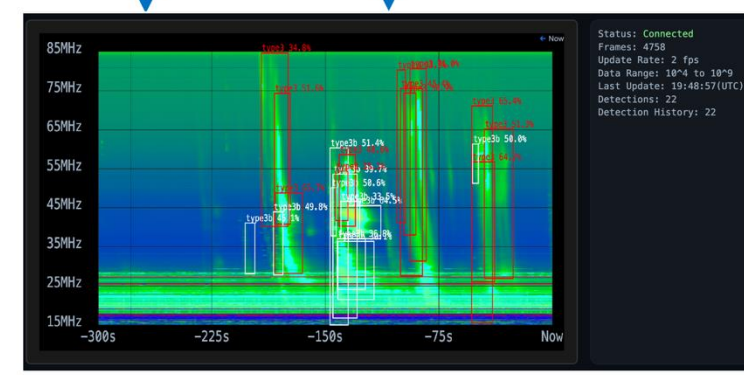
UDP packets



Accumulated spectrum
For real-time detection

Serve low latency monitoring
Delay < 0.7s

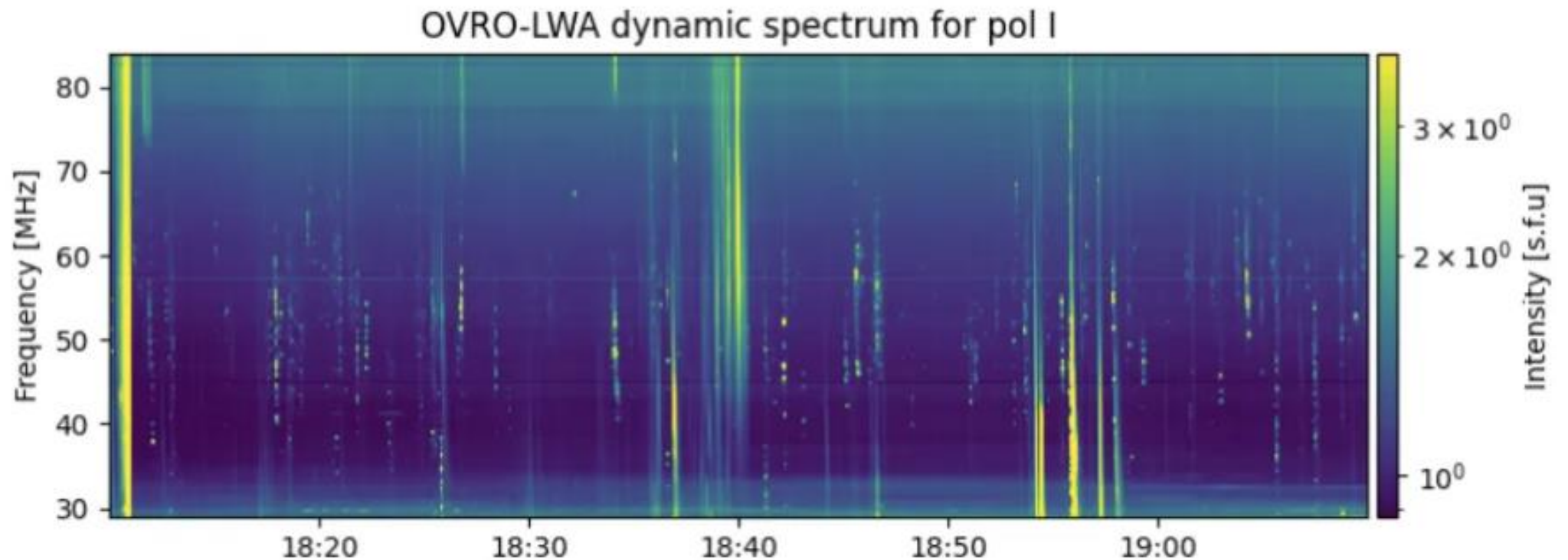
Radio burst event detection module



<https://ovsa.njit.edu/live/>

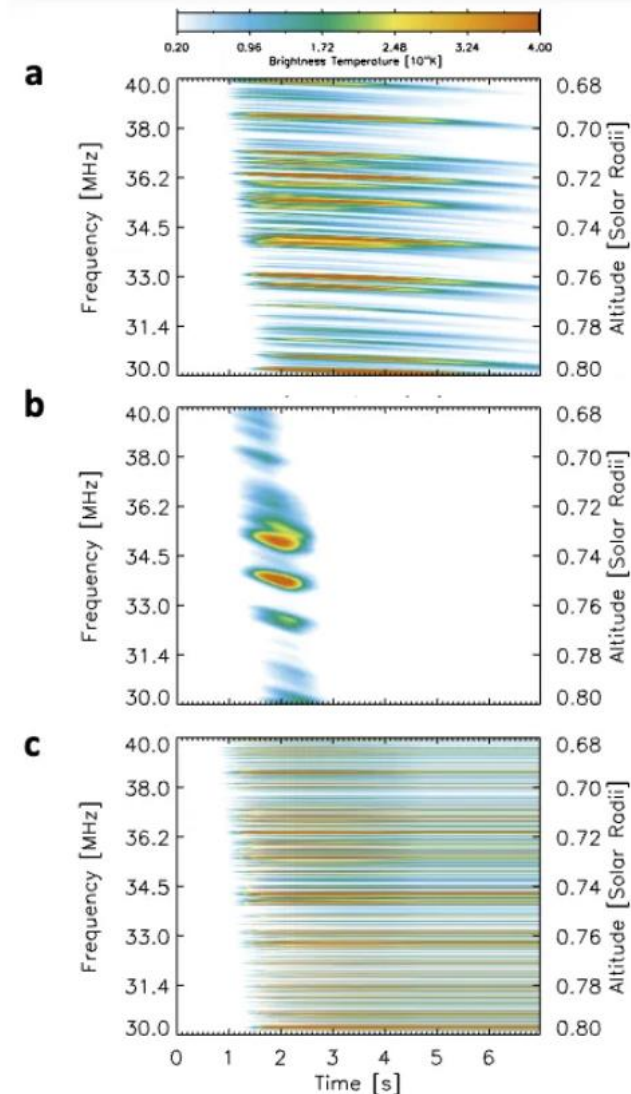
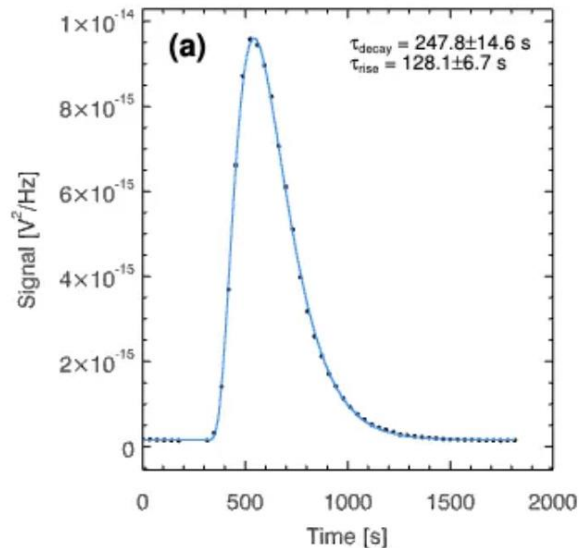
Event detection

- Radio burst detection with YOLO
- Requires large amount of precise labels of bounding box

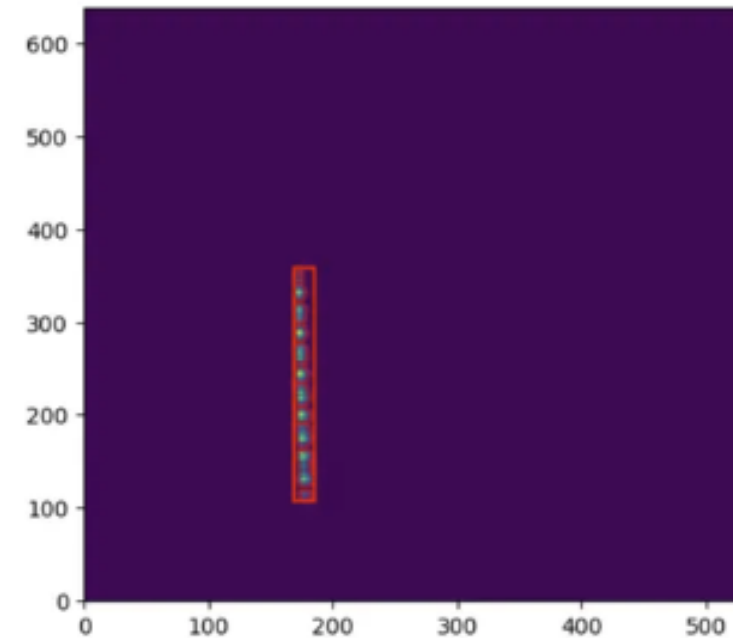


Physics informed training set generation

- Time profile
- Frequency drift rate
- Fine structures
- Starting ending freq

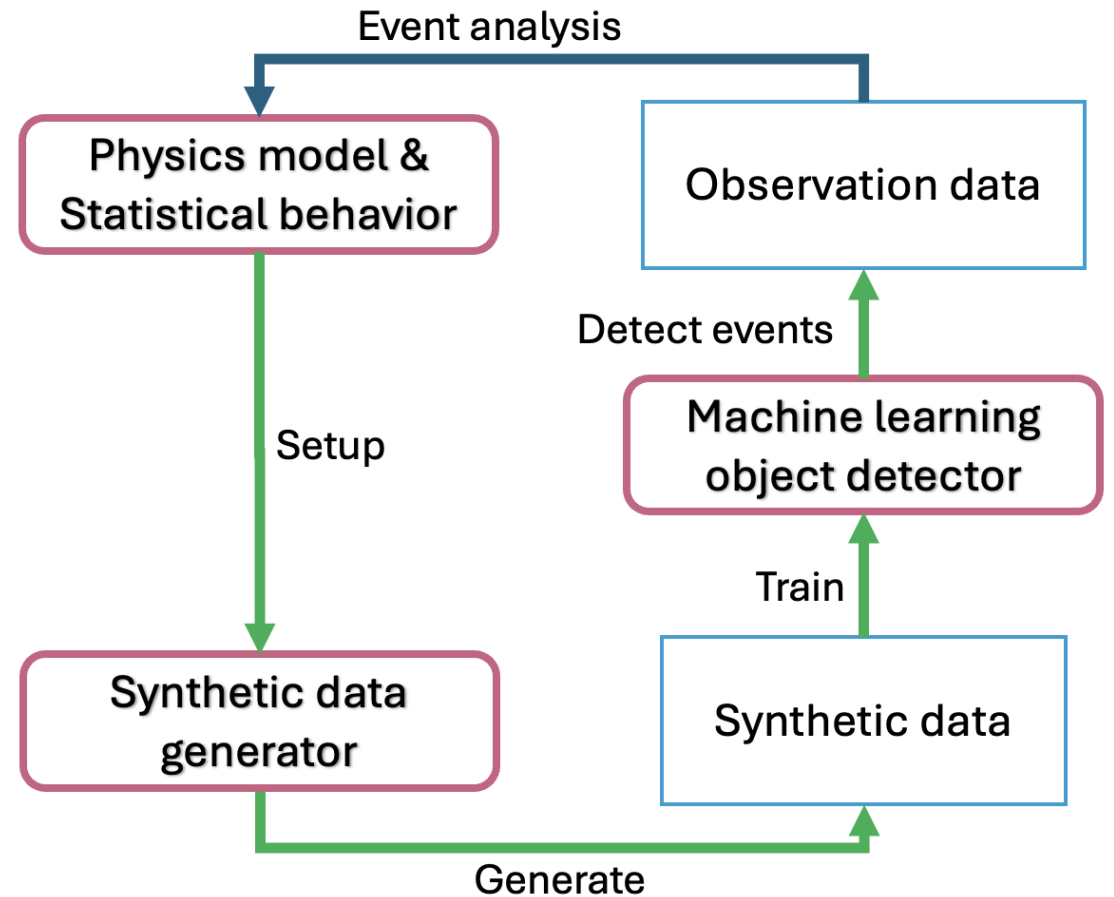
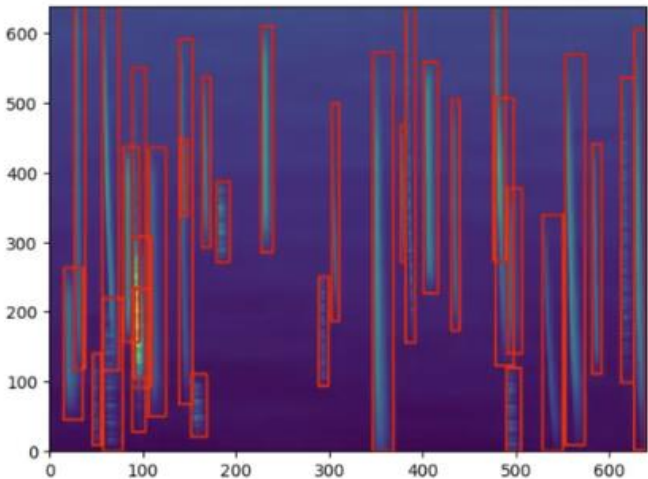
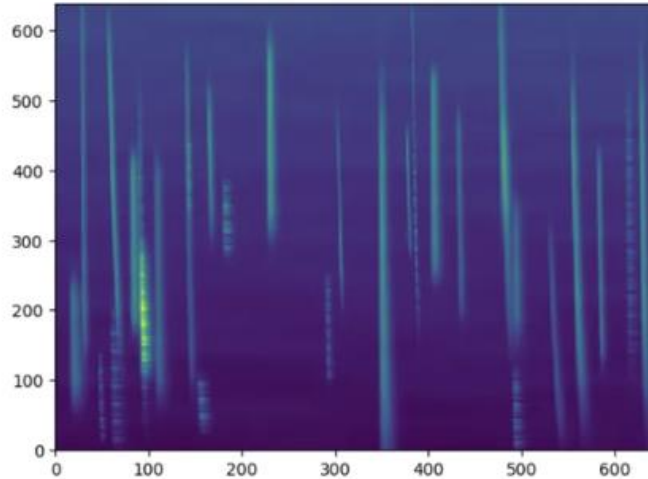


Generated burst will have the ground truth bounding box and label



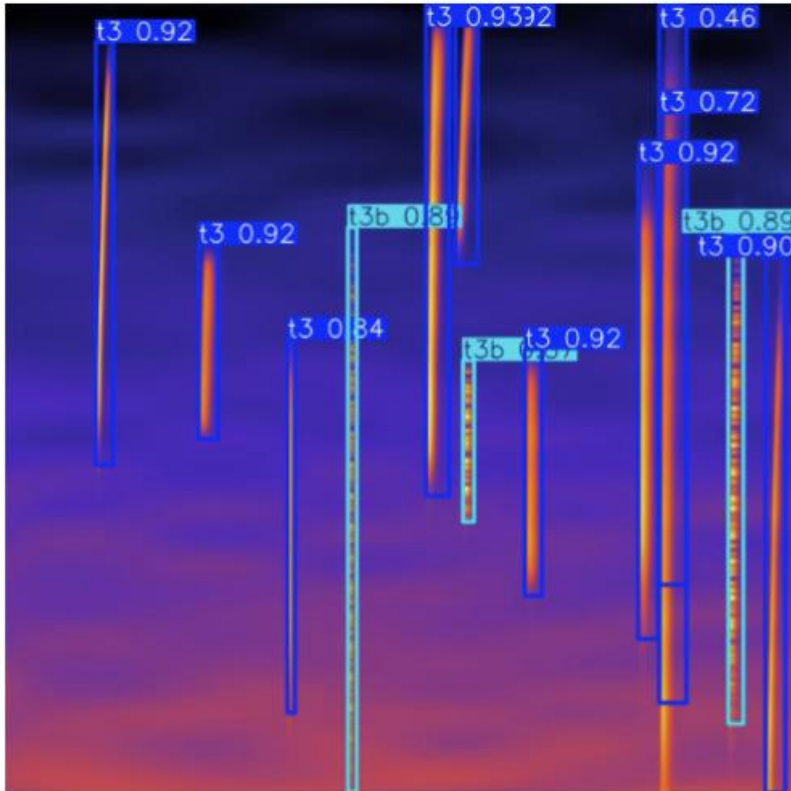
Training set and detection

Infinite amount of sample for training

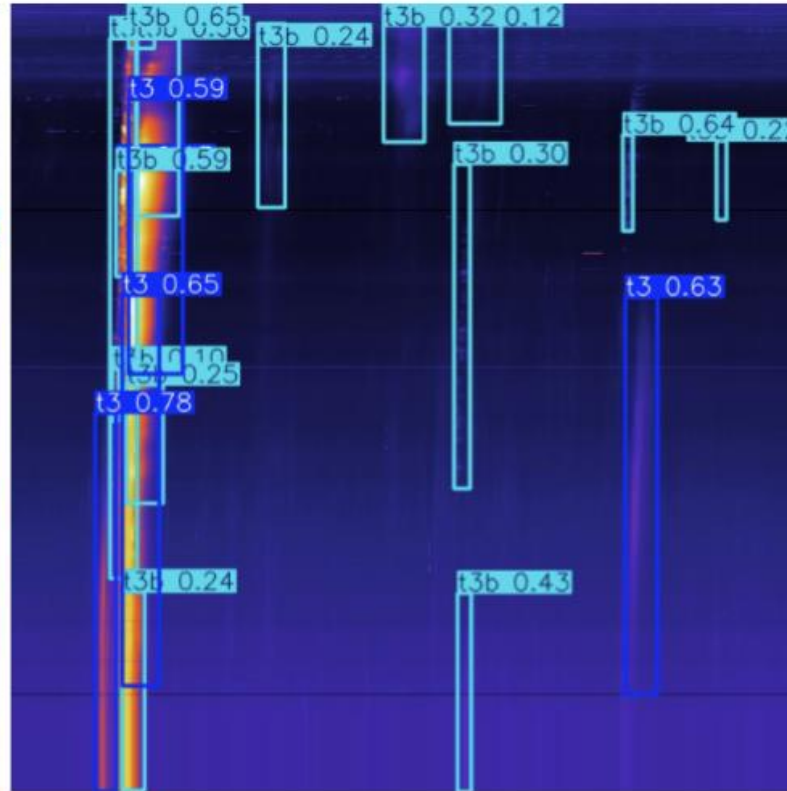


Result

Validation on generated data



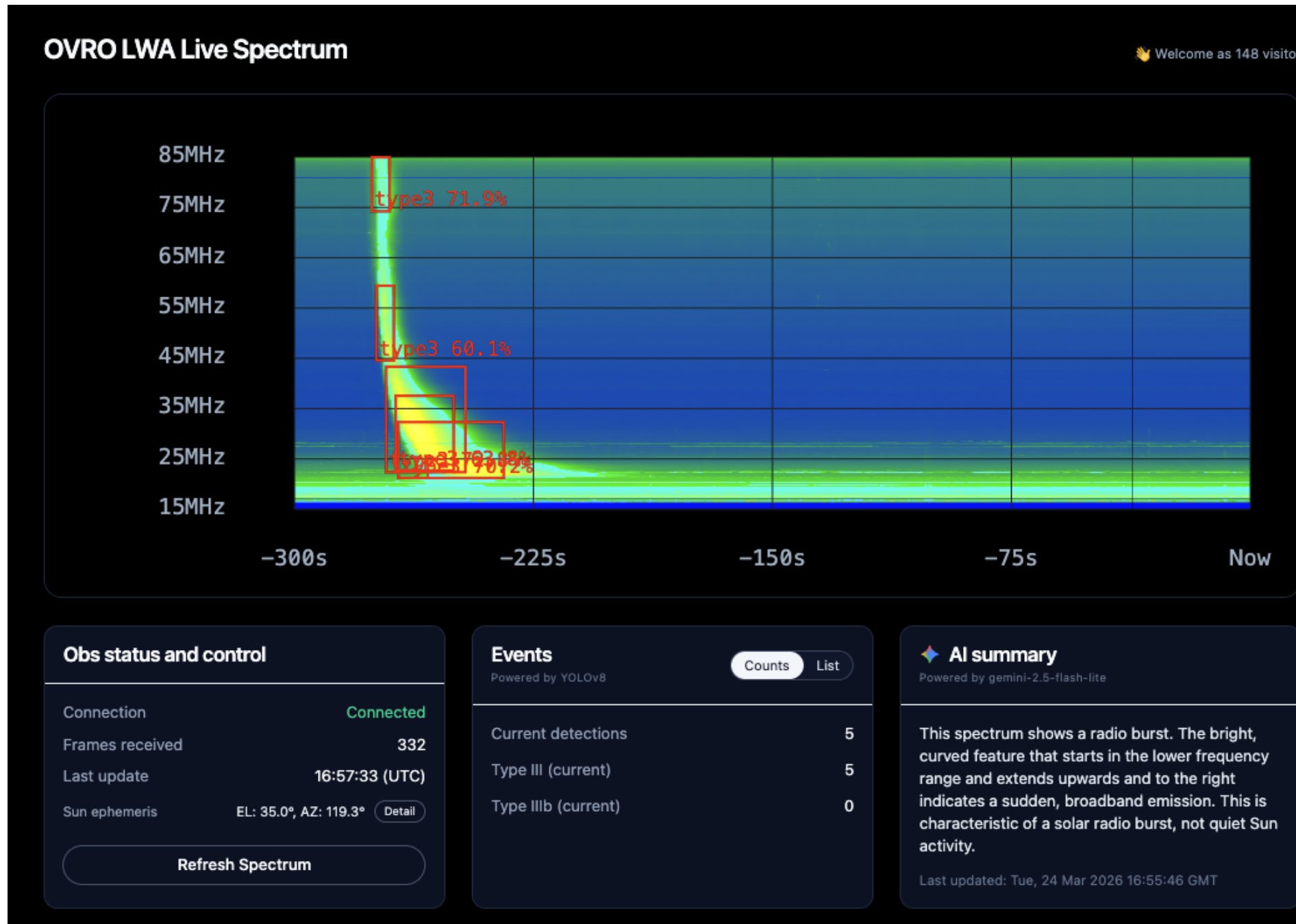
on real observed data



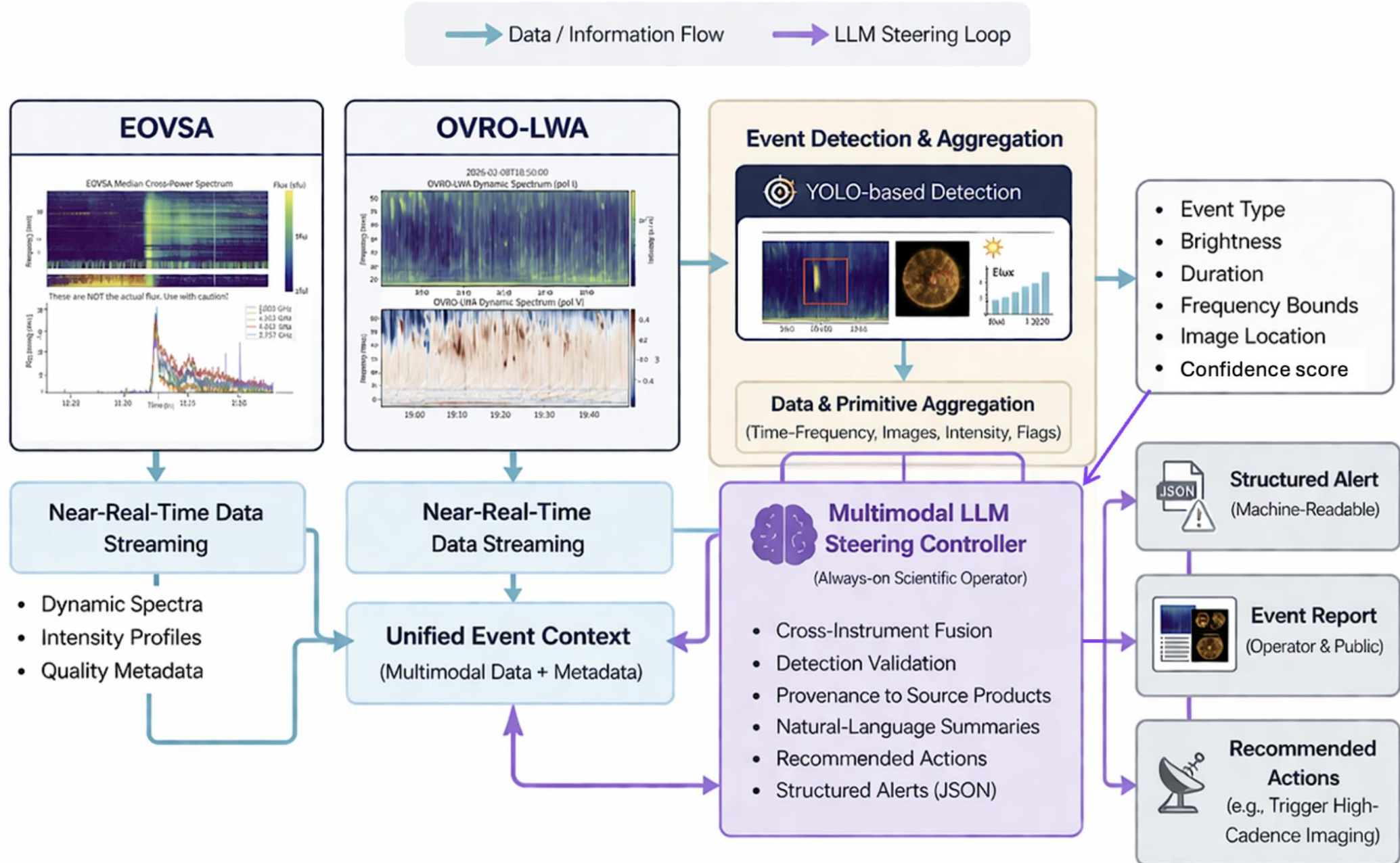
<https://ovsa.njit.edu/live/>

YOLO+LLM

- LLM assisted real-time observing
- Send data and header and event detections to AI for summary and alerting

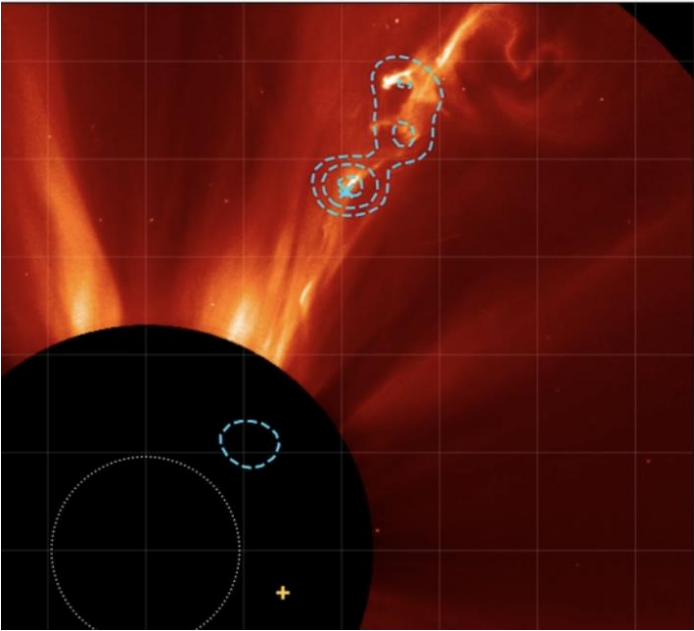


Workflow of low-latency AI-assisted framework

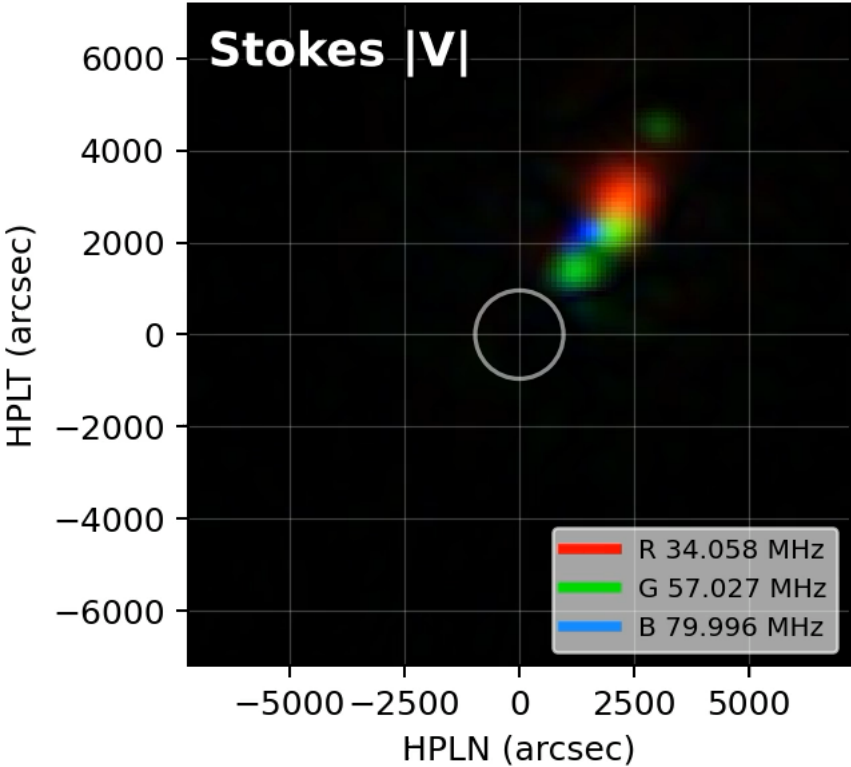
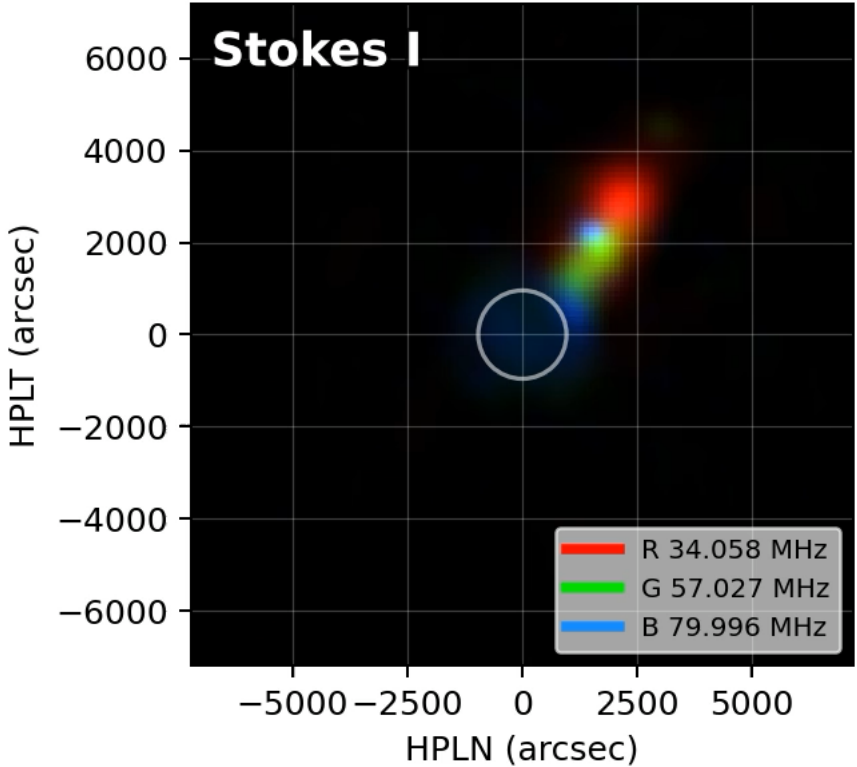


Next step, imaging and polarization (work in progress)

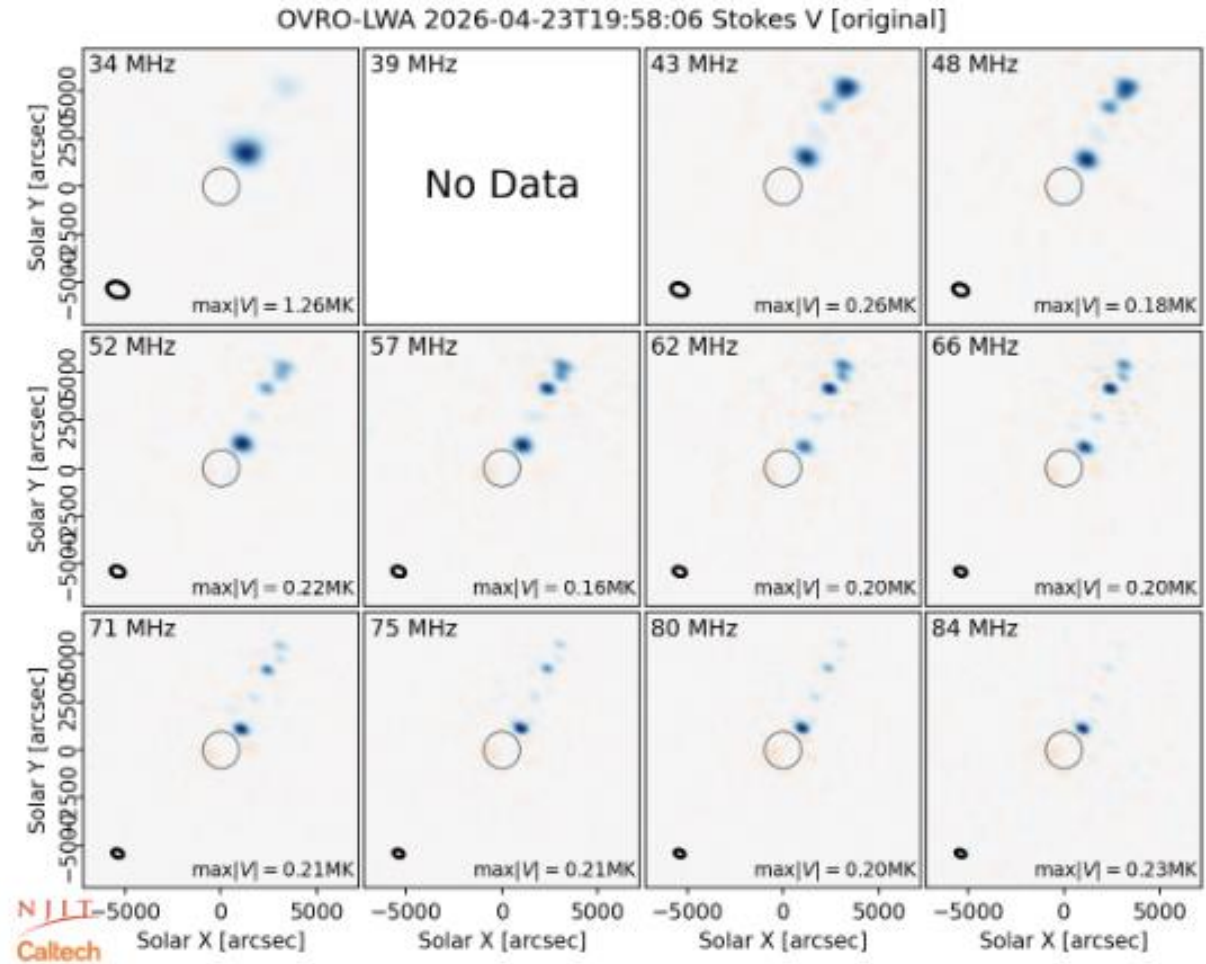
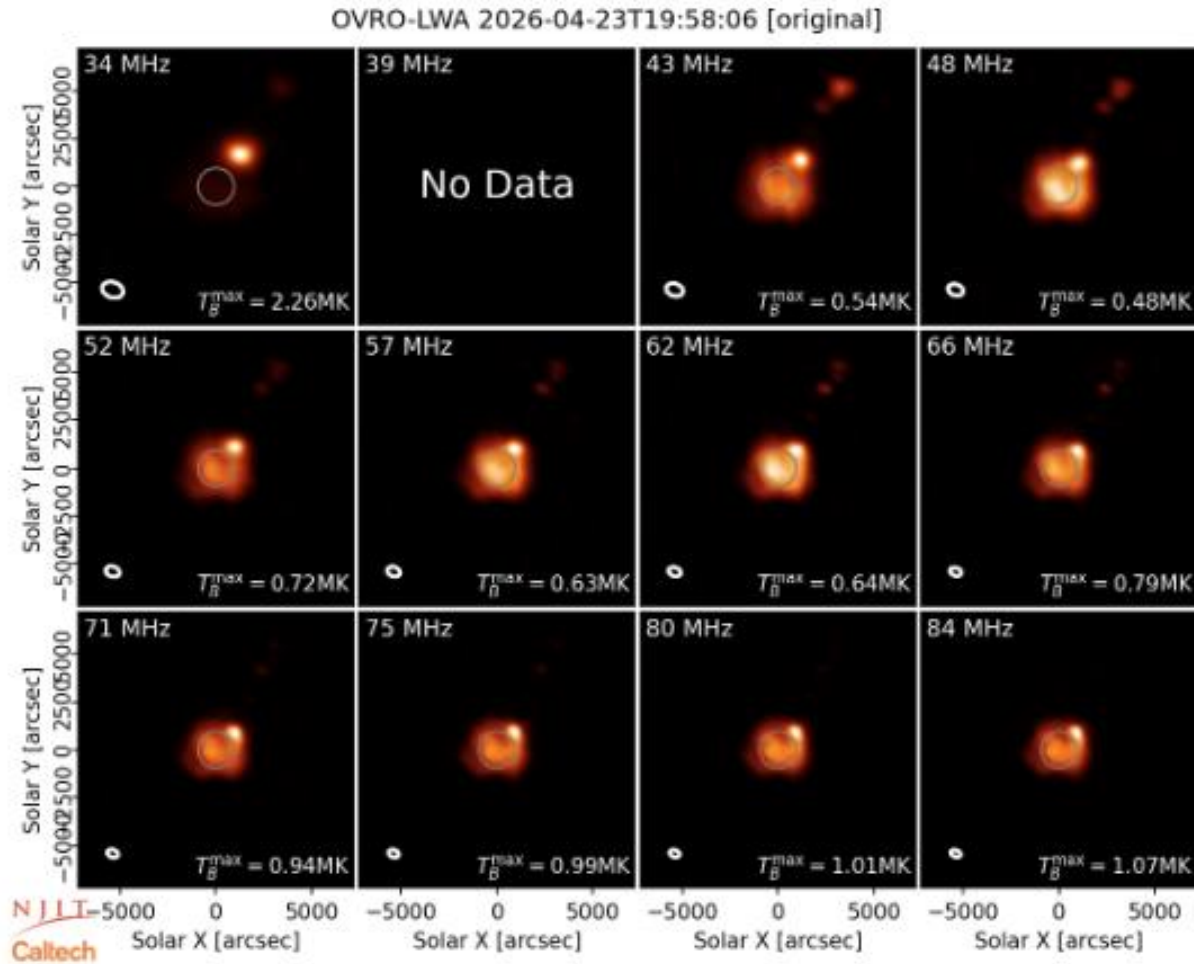
SOHO/LASCO C2 2026-04-23 19:36 UT with OVRO-LWA V-pol contours
Radio: 19:36:32.5 UT, 84.015 MHz, channel 136



OVRO-LWA MFS RGB, 2026-04-23T18:30:00.1

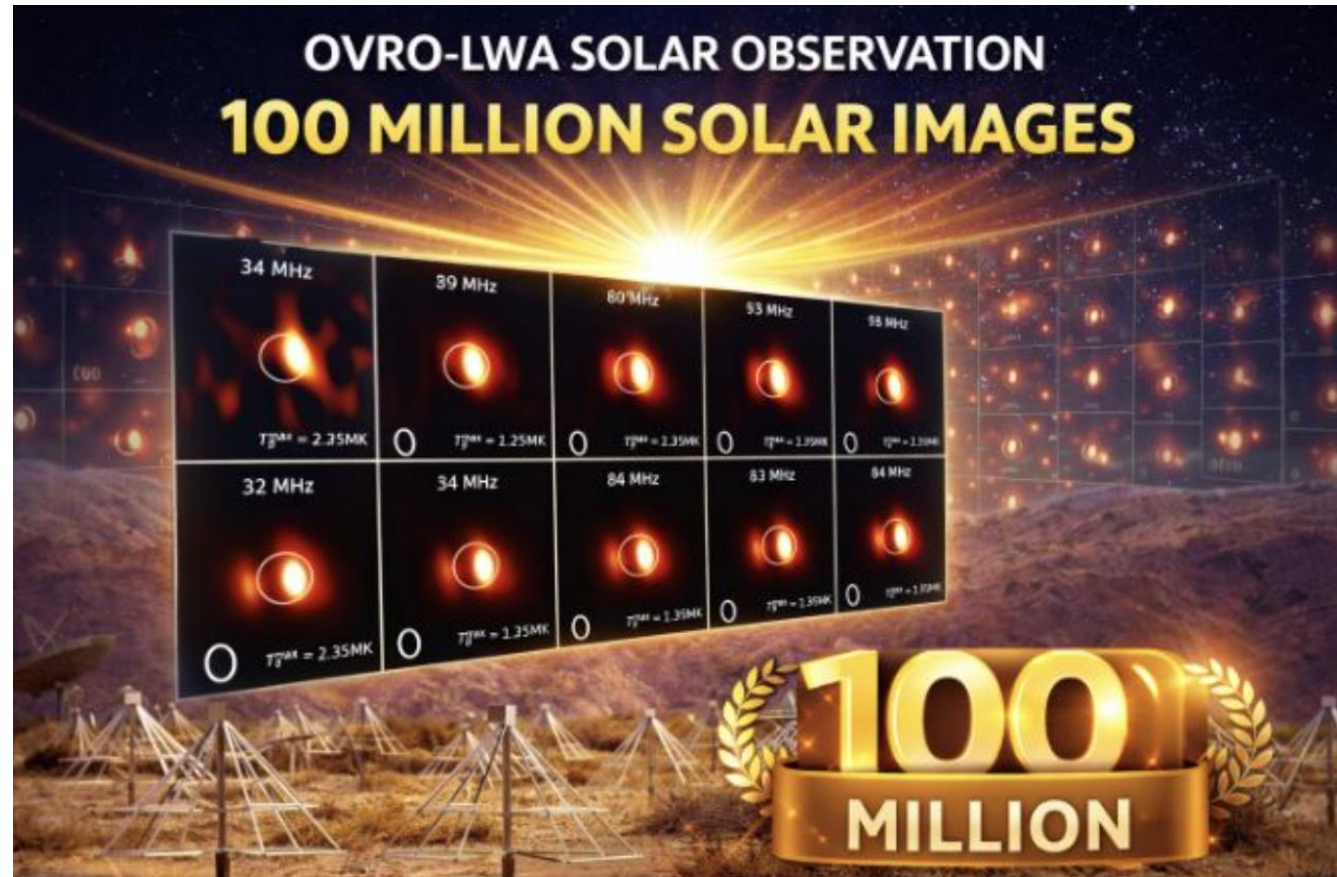


Polarization imaging is a natural coronal graph



Also next step

- Composing ML dataset of dynamic spectrum and imaging
- 100M images achieved on Feb-2026 and growing



Summary

- Buffer-streaming enables realtime datastream of solar ..(**<0.5s** delay) of dynamic spectrum
- Report radio burst **<5s** of event happening, serving space weather nowcasting and forecasting.
- Physics based training set generation can save huge amount of labeling time and giving the ground truth labeling
- LLM assisted observing interpretation and alerting